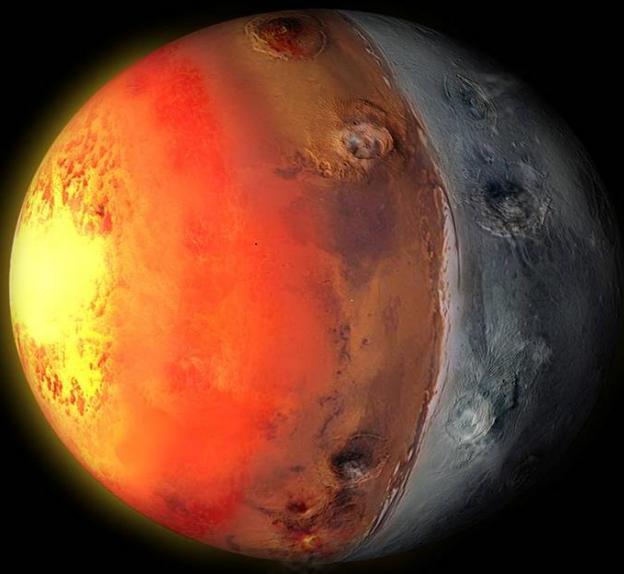


Planetas Extra-Solares ou Exoplanetas



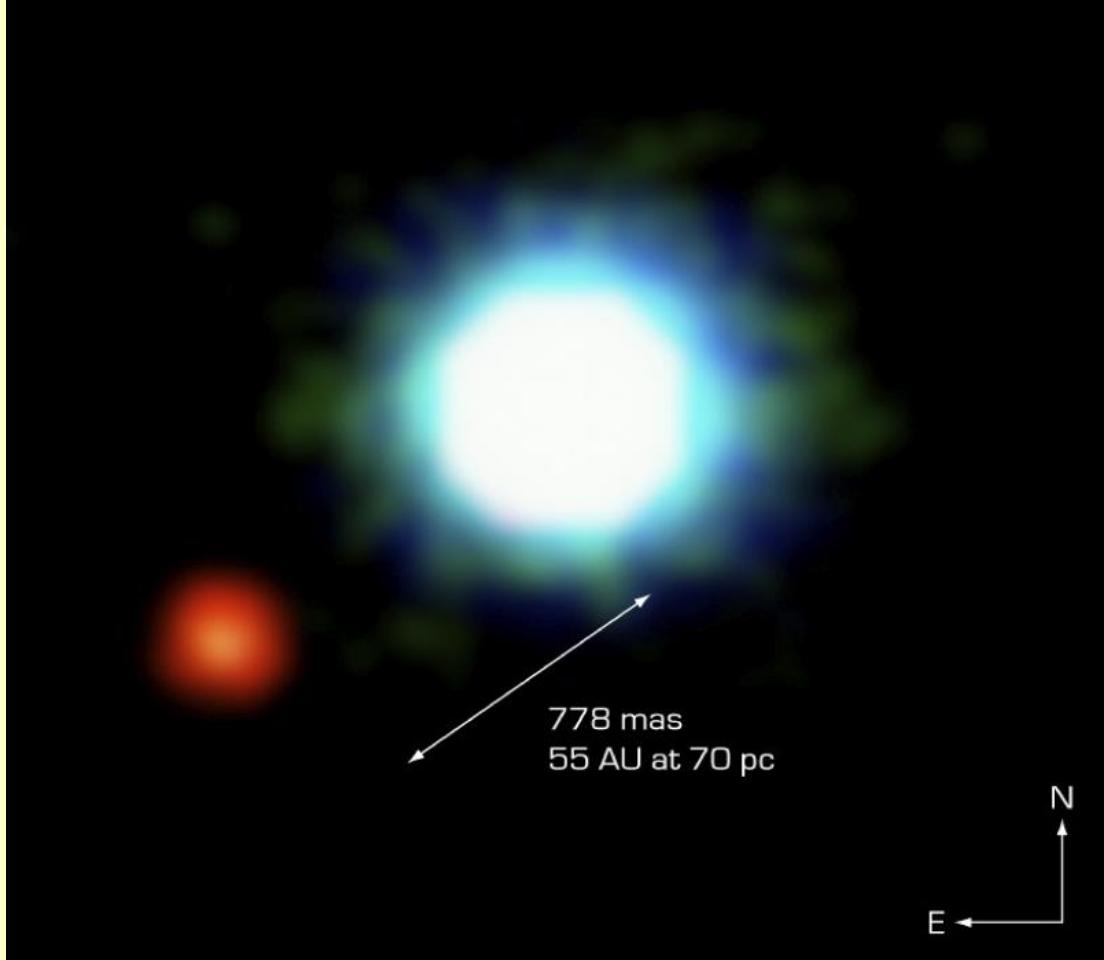
Sylvio Ferraz Mello
Instituto de Astronomia,
Geofísica e Ciências
Atmosféricas
Universidade de São
Paulo

IMAGENS

Luminosidade
de um planeta
igual a Júpiter

=

Luminosidade do Sol
1 000 000 000

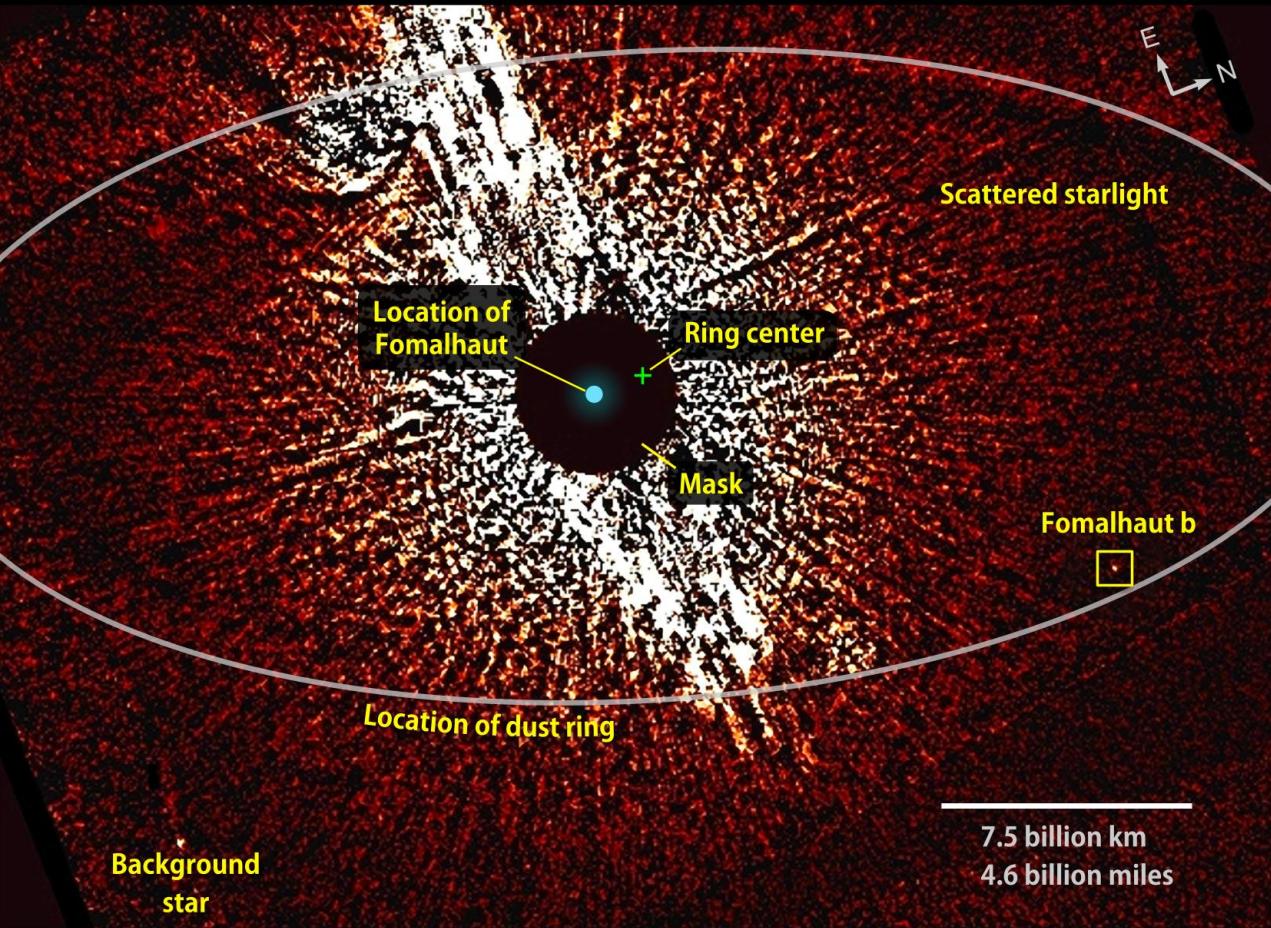


The Brown Dwarf 2M1207 and its Planetary Companion
(VLT/NACO)

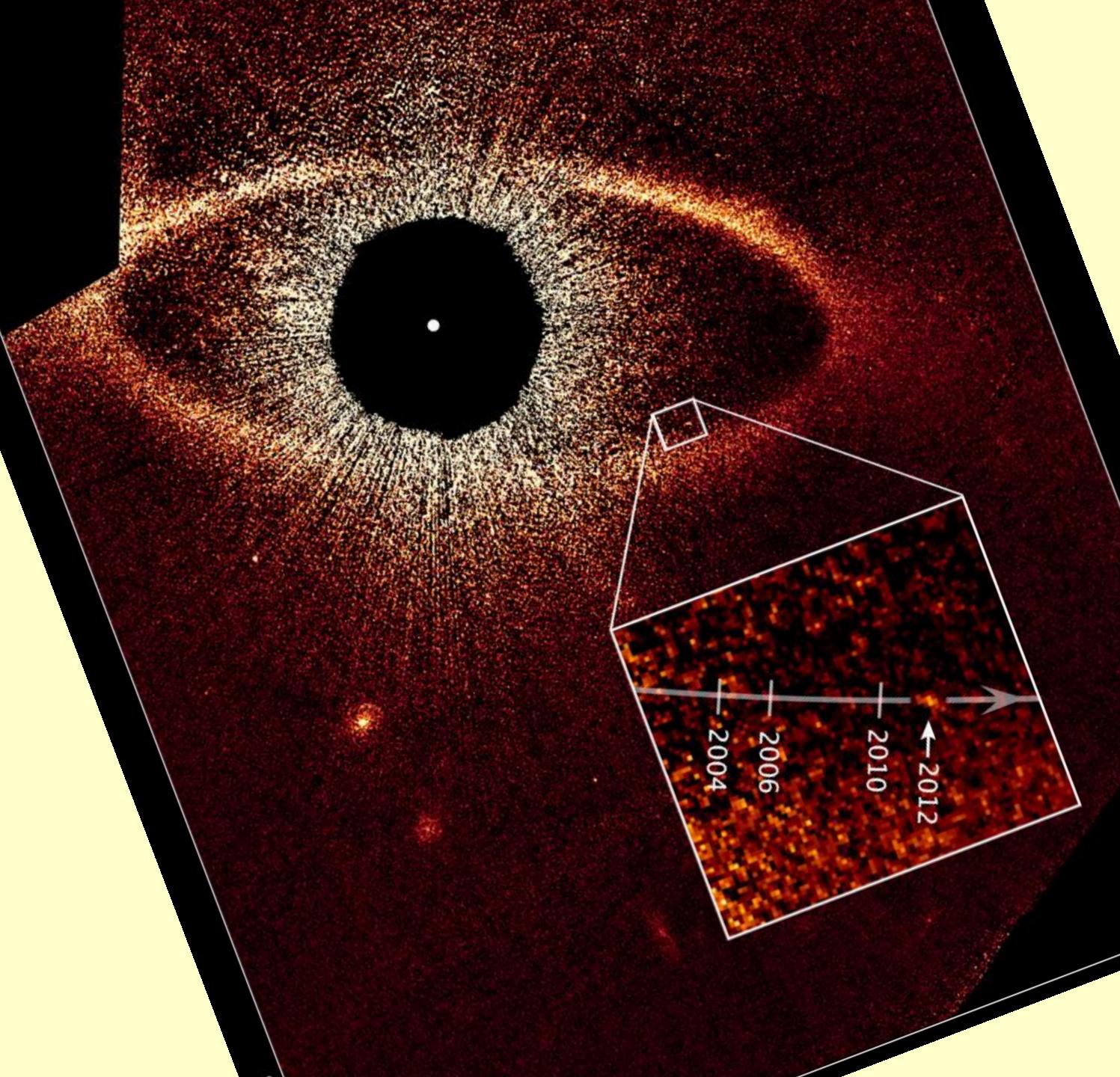
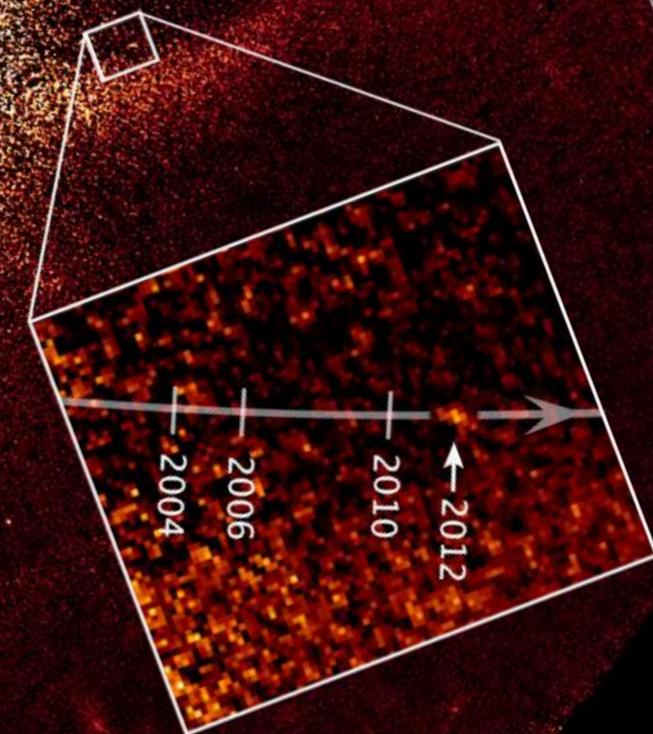
ESO PR Photo 14a/05 (30 April 2005)



© ESO



Hubble Space Telescope • 31



NASA and ESA

Fomalhaut's ring shepherd planets



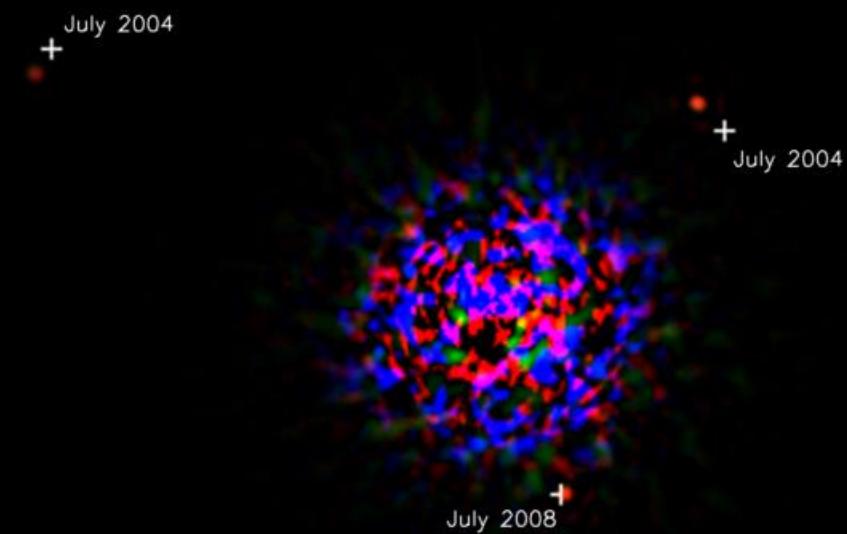
HD 19467

N
E



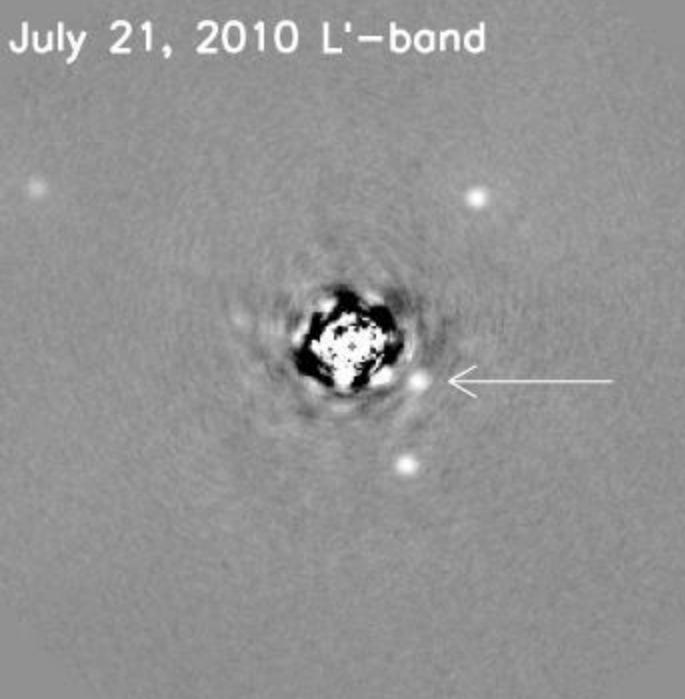
$1''$
 $\overline{30.9 \text{ AU}}$

Planets Orbiting HR 8799 (Sept. 2008)

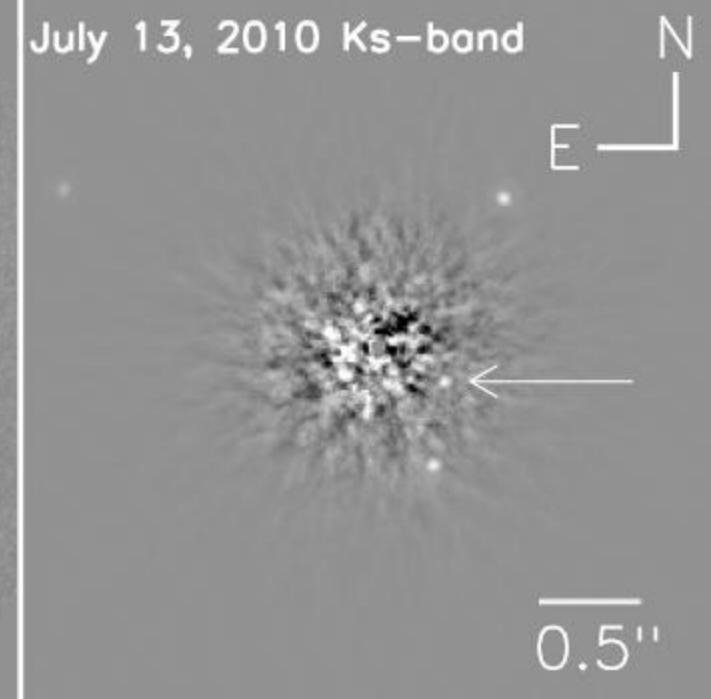


0.5 arcsec
20 AU

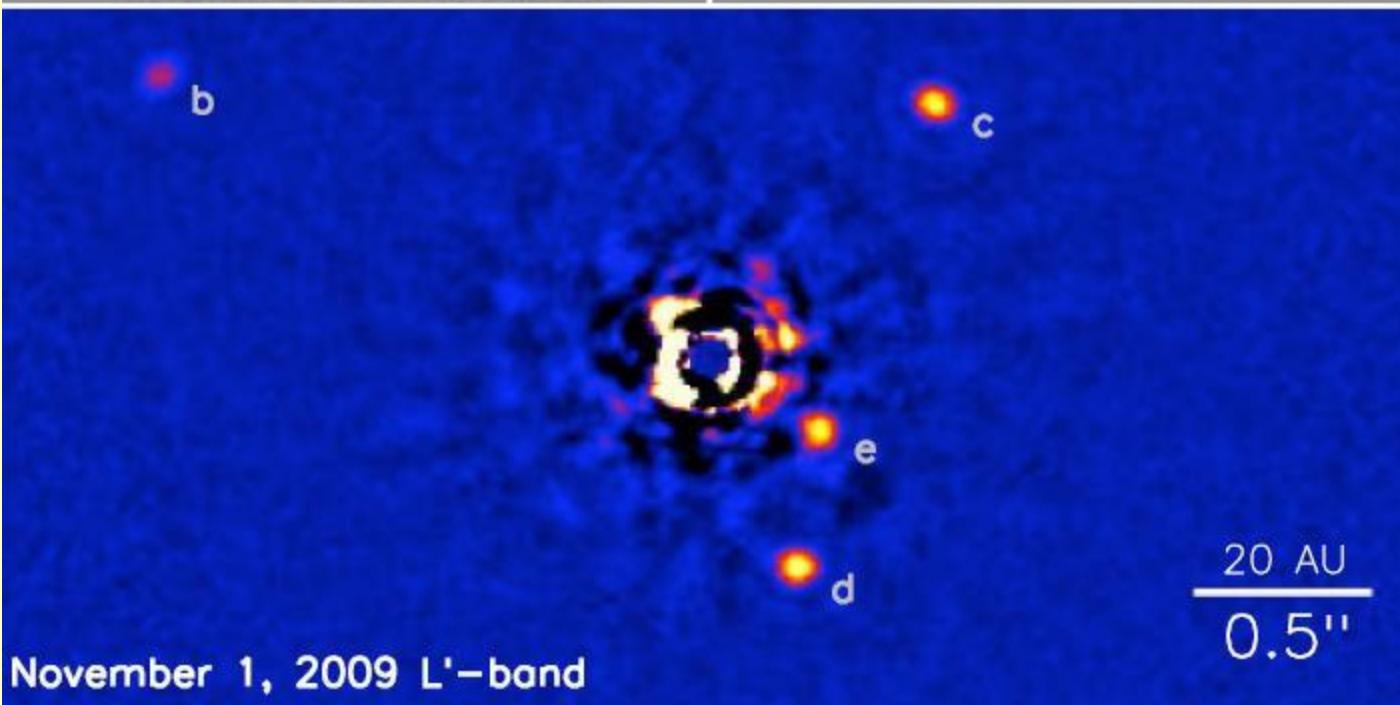
July 21, 2010 L'-band



July 13, 2010 Ks-band

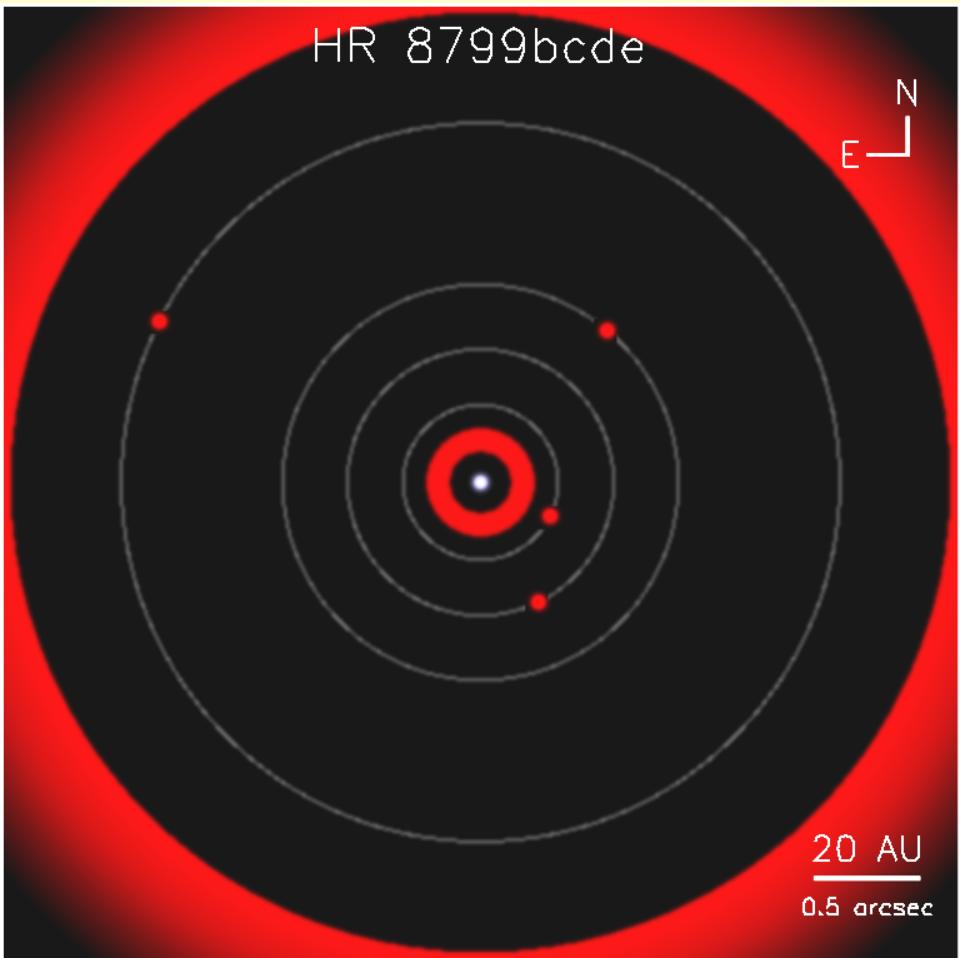


0.5''

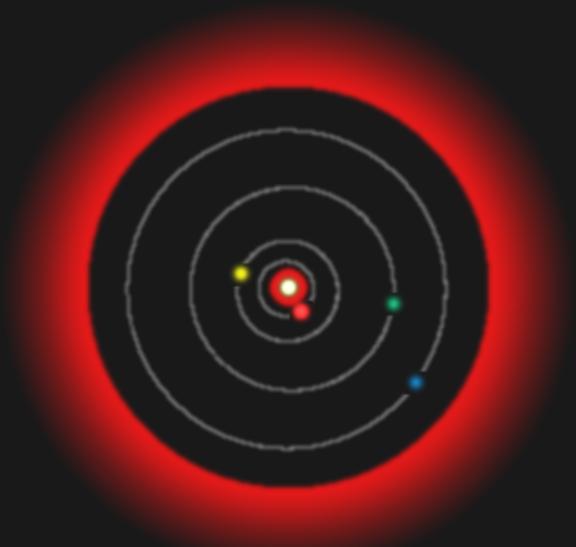


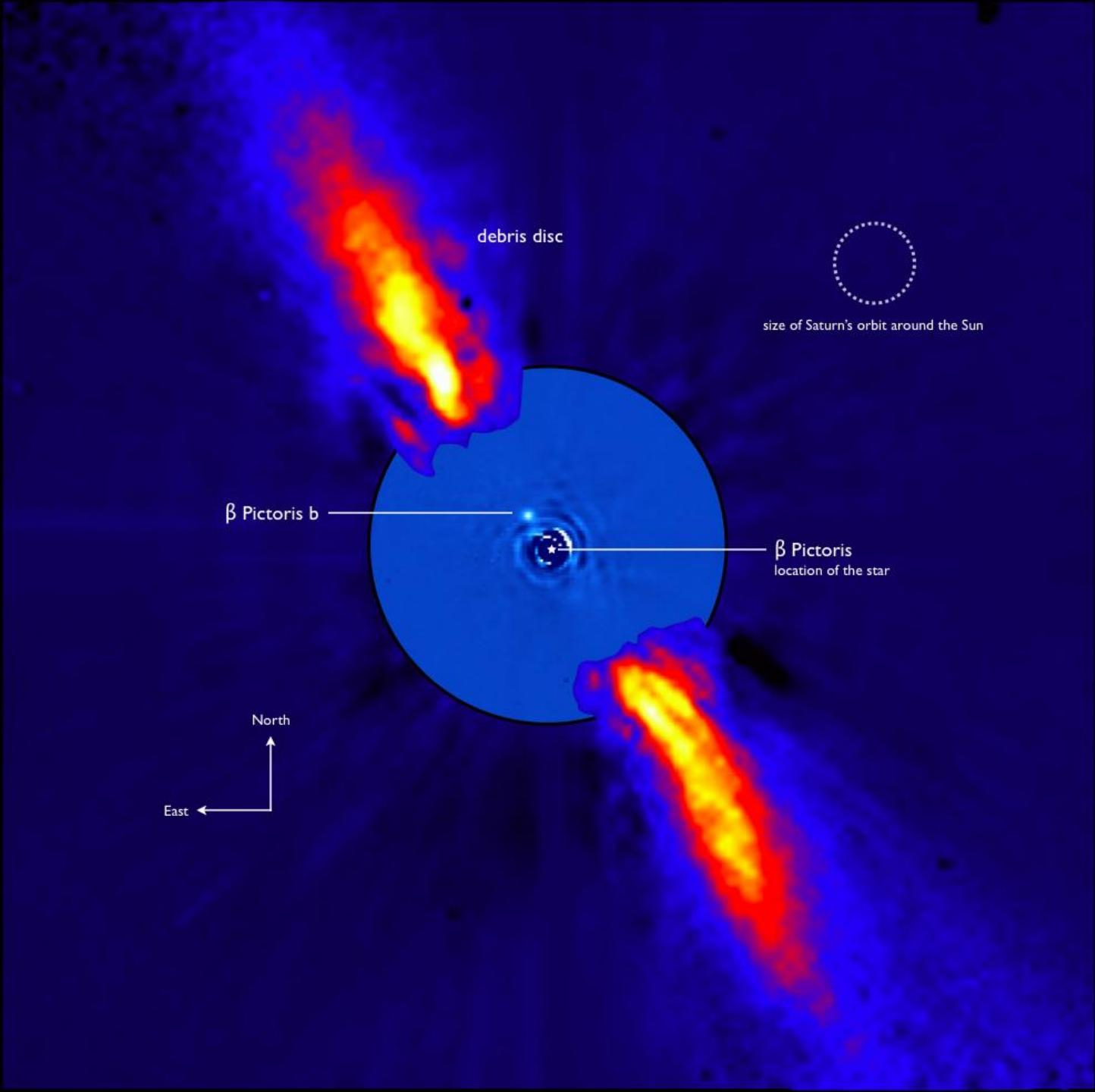
November 1, 2009 L'-band

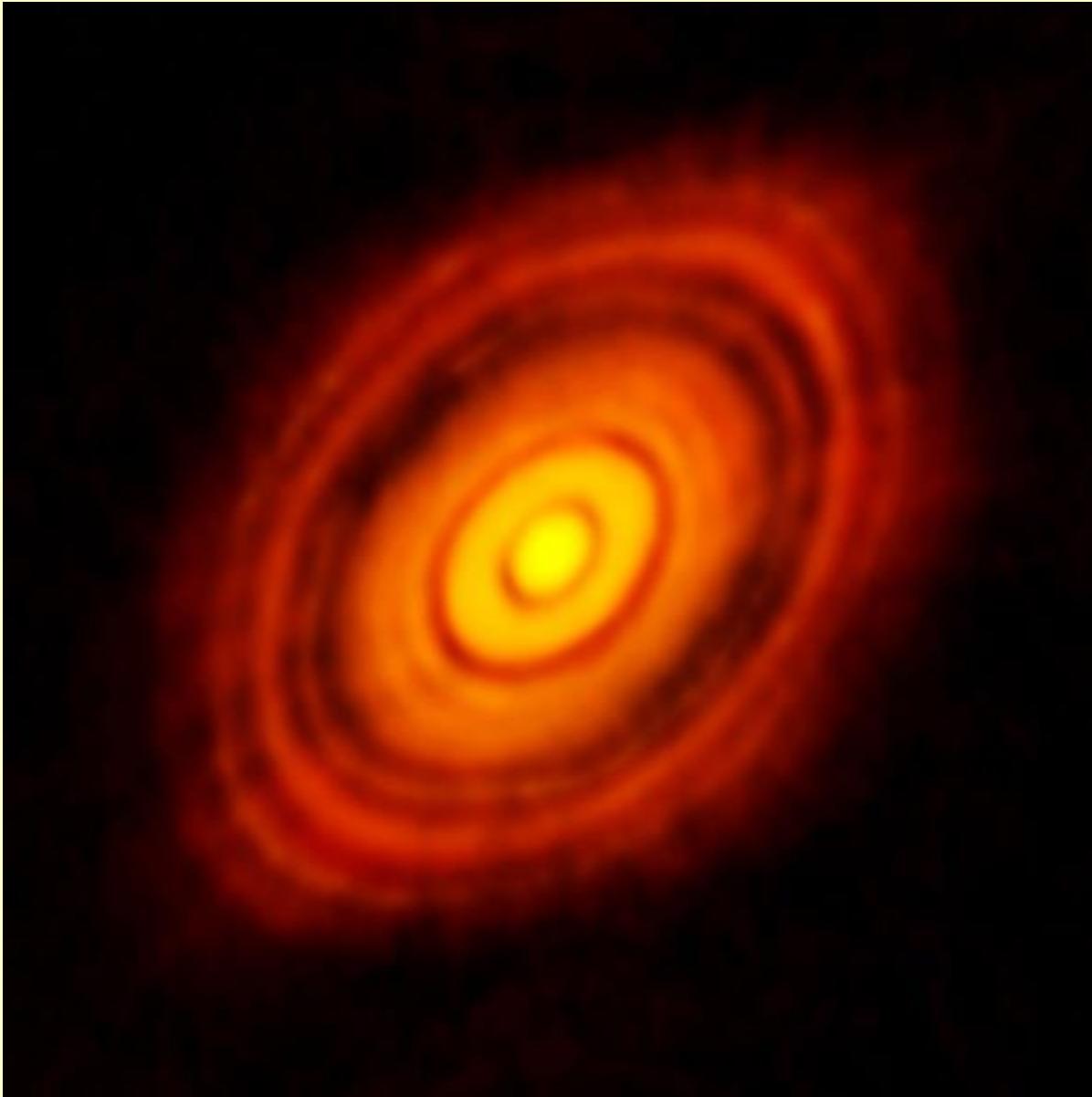
HR 8799bcde



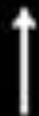
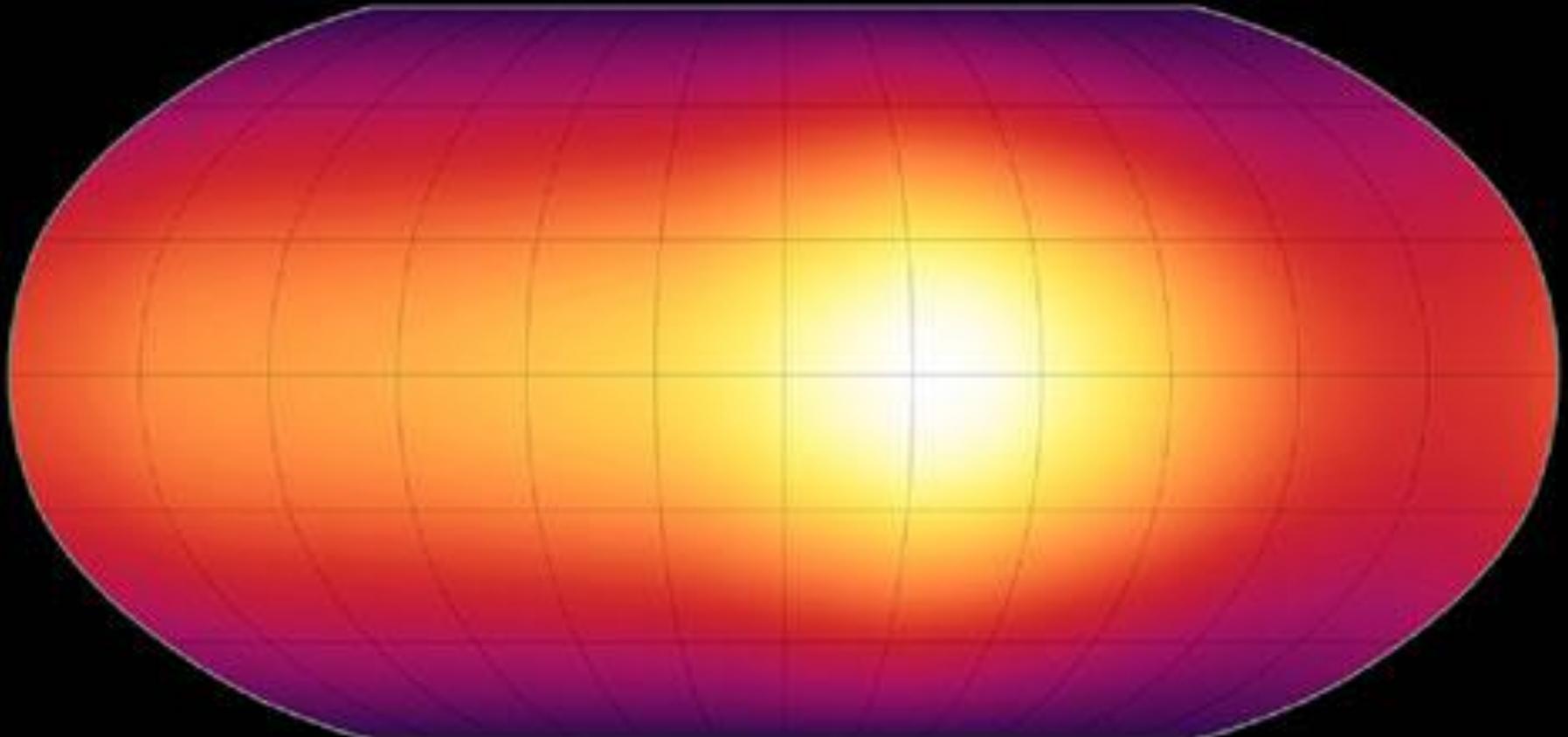
Solar System







HL Tauri



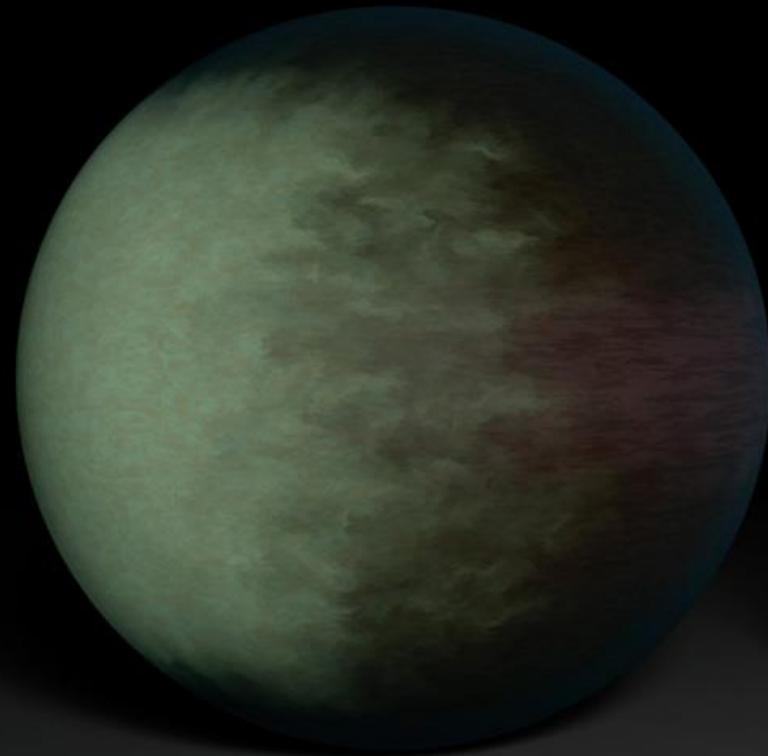
Sun-Facing Longitude

(Grid Spacing: 30°)

Global Temperature Map for Exoplanet HD 189733b
NASA / JPL-Caltech / H. Knutson (Harvard-Smithsonian CfA)

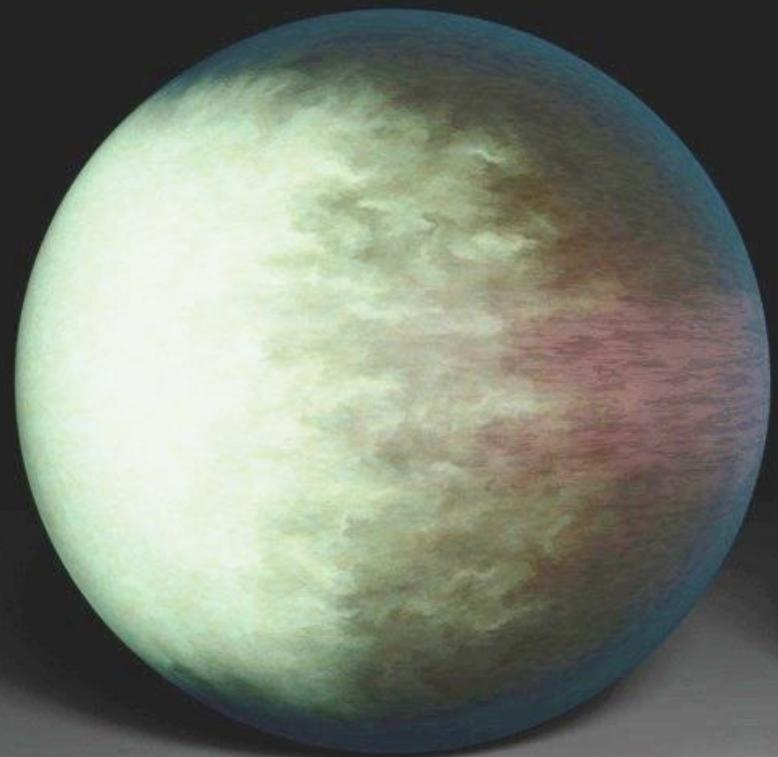
Spitzer Space Telescope • IRAC
ssc2007-09a

Kepler 7b (IR/Spitzer)



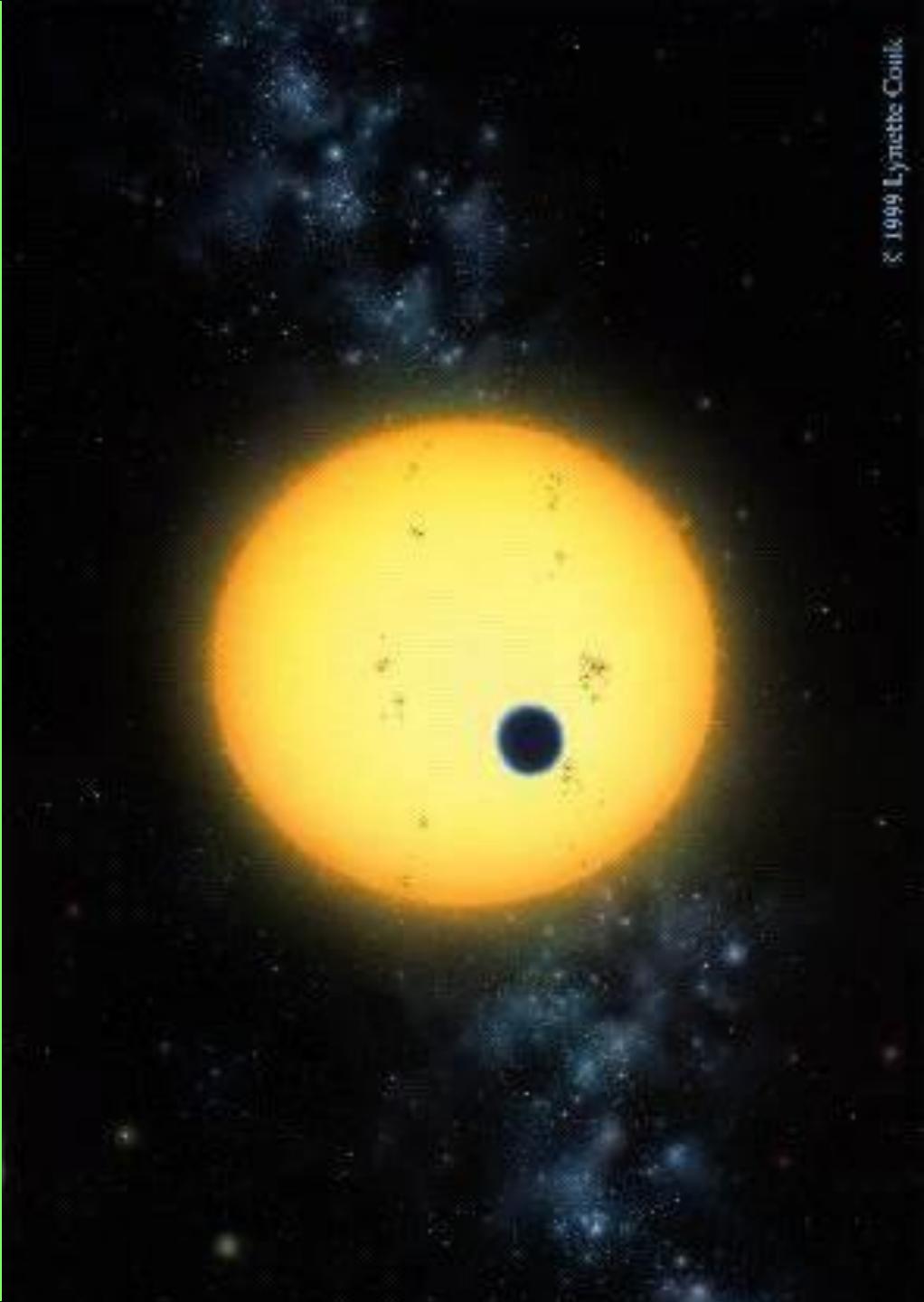
$T \sim 1400$ K

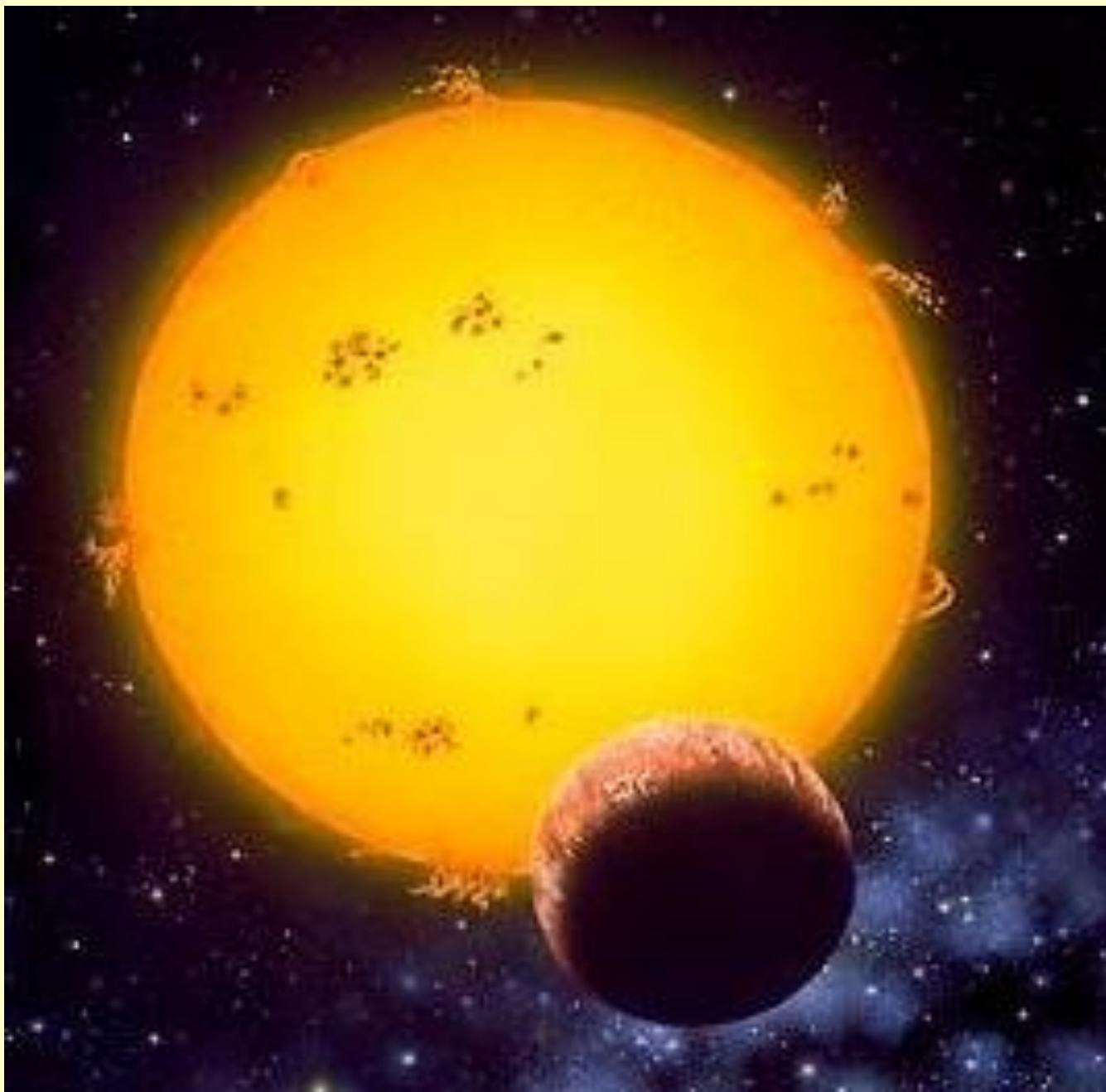
$a \sim 9$ million km

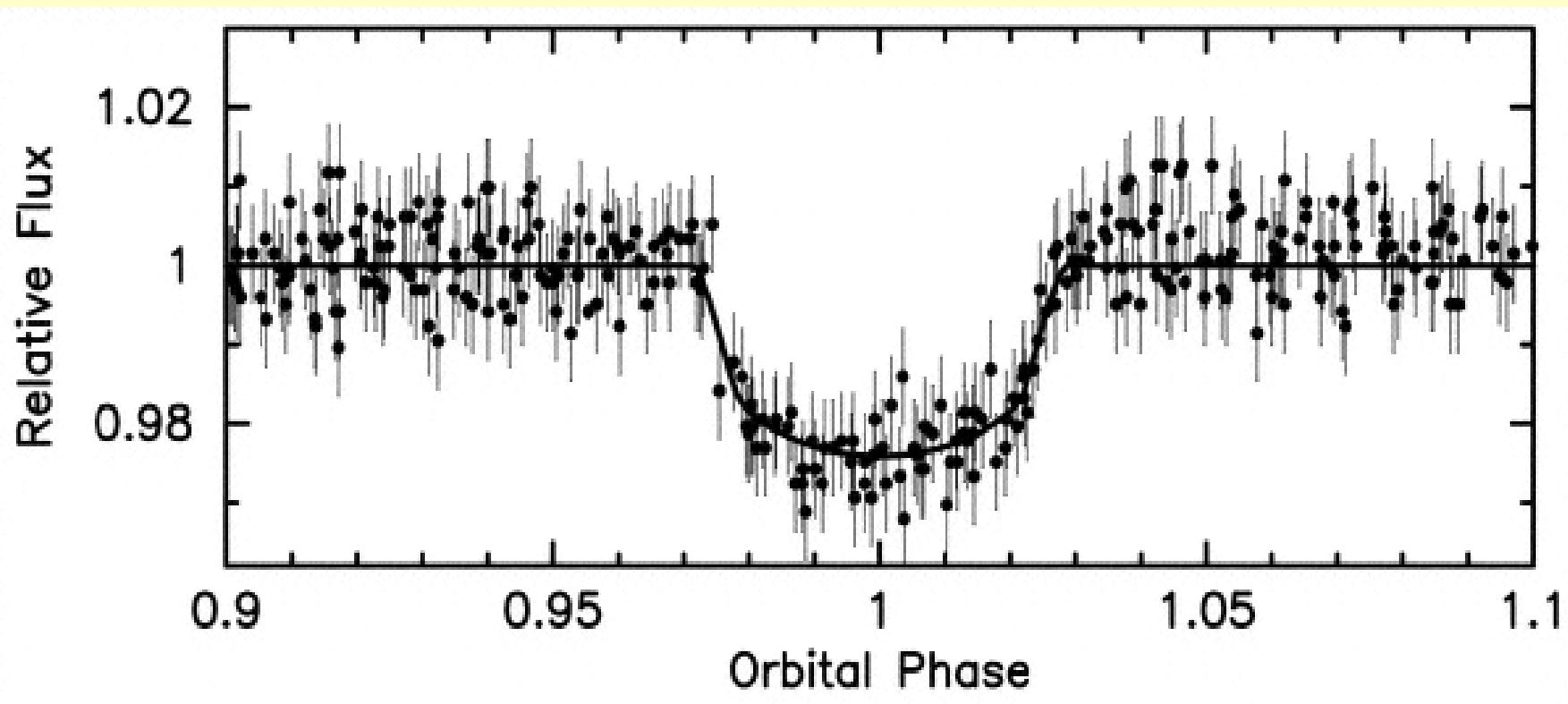


TRÂNSITOS

(O tamanho dos Planetas)





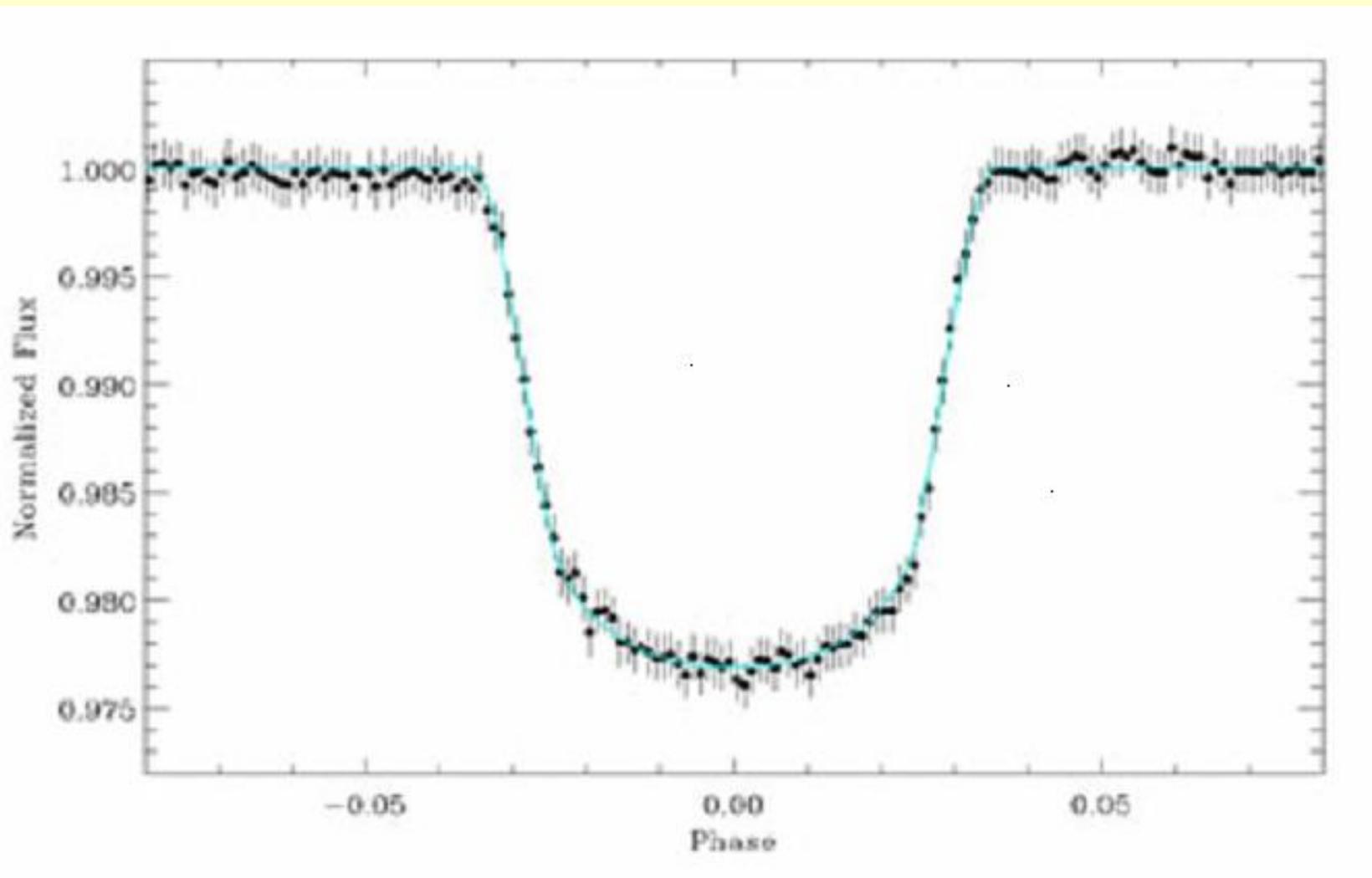




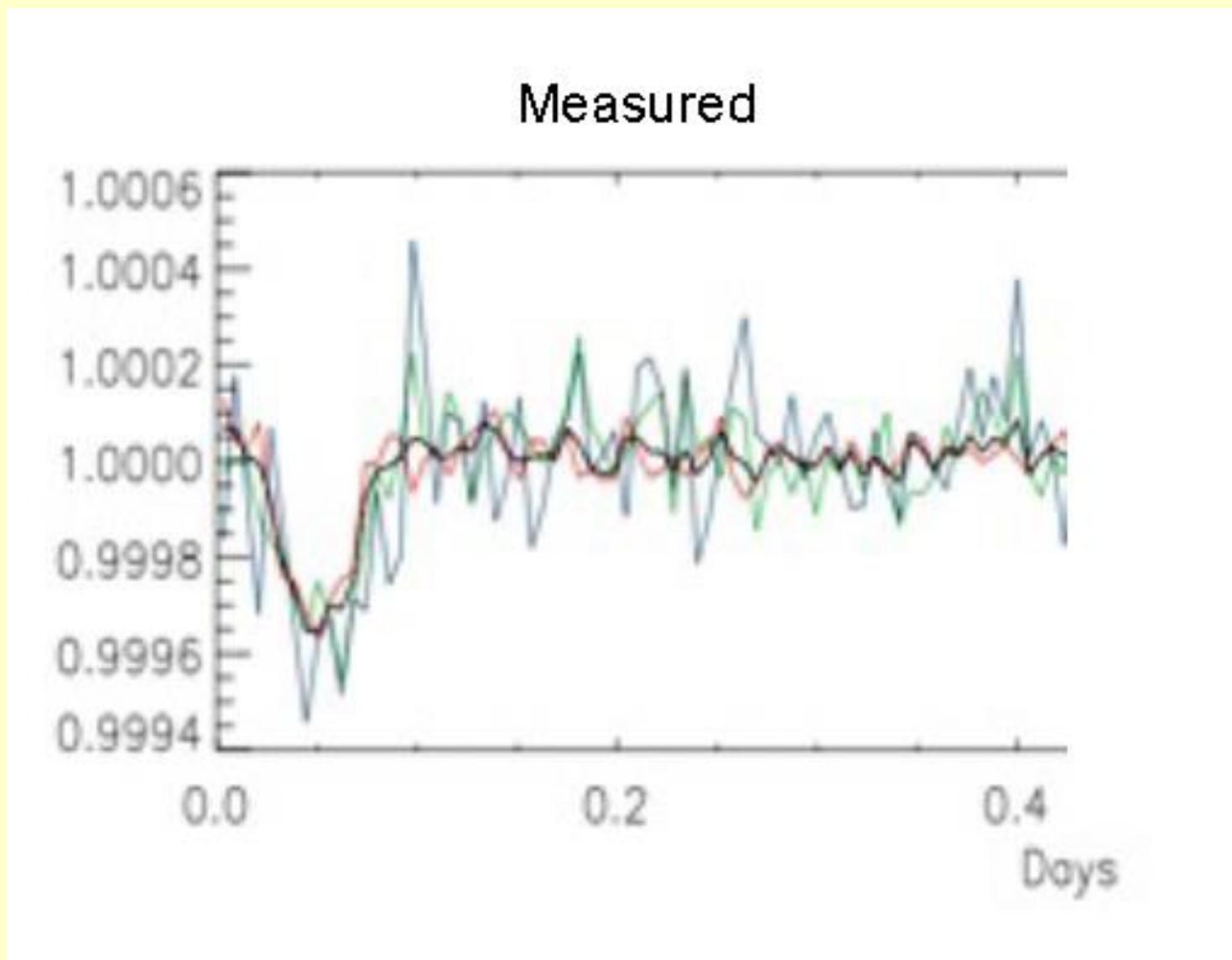
(CNES+ESA+ ... + Brasil)
2006 – 2014
ø 20 cm telescope
4 fields



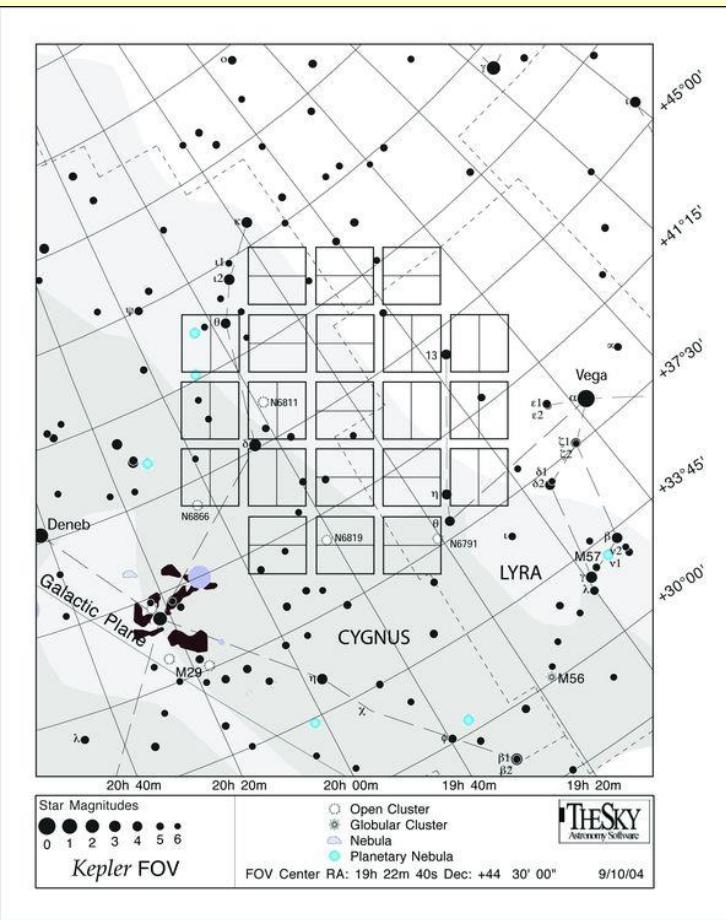
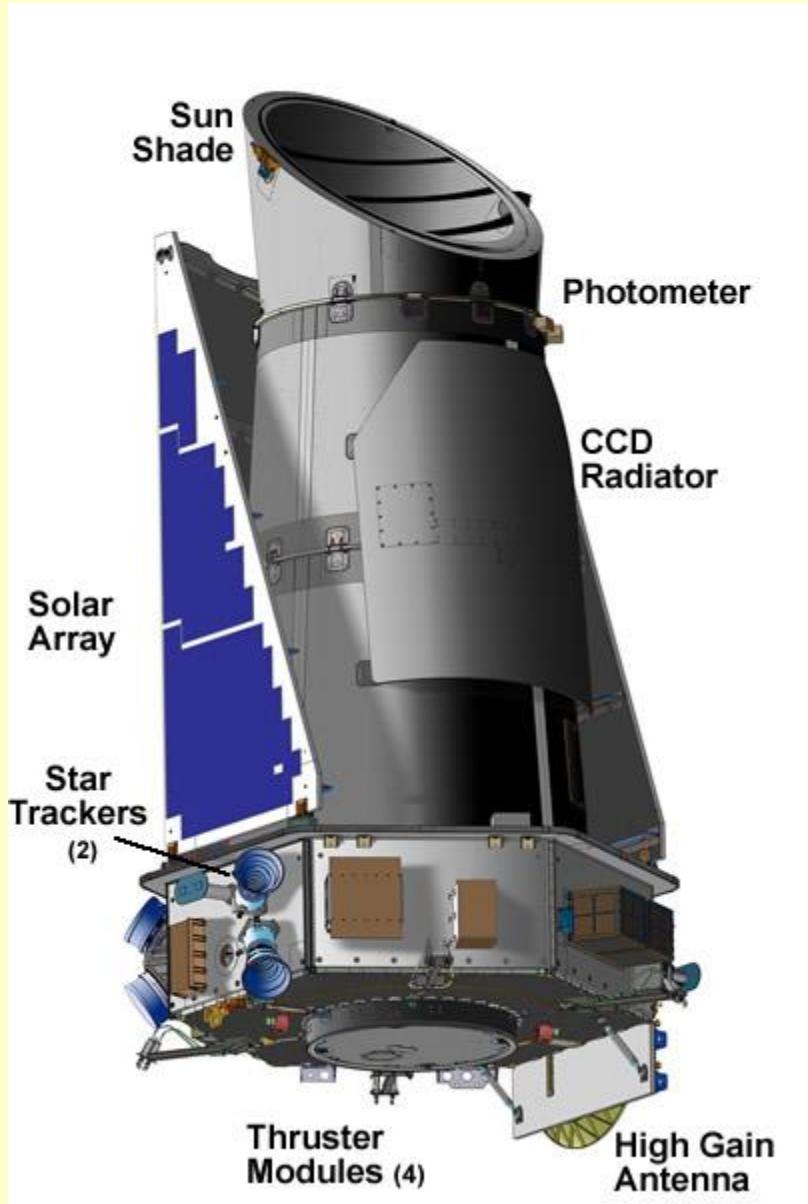
CoRoT-exo-1b



CoRoT 7b (Variação 0.0004)



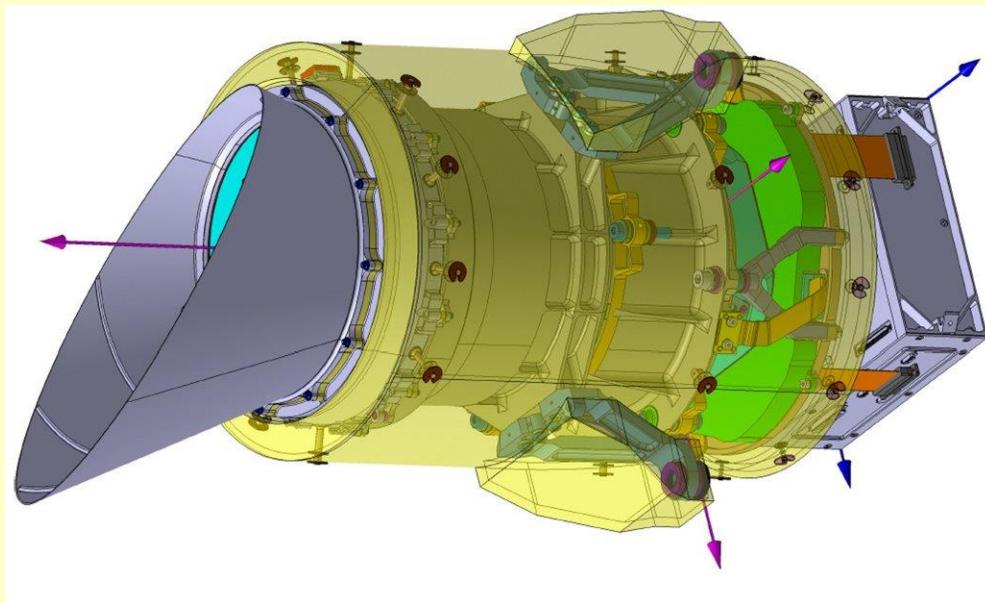
V=11.7



KEPLER (NASA)
2009 -
ø 1m telescope
21 fields



2018
TESS (NASA)
 \varnothing 10.5 cm telescope
4 fields



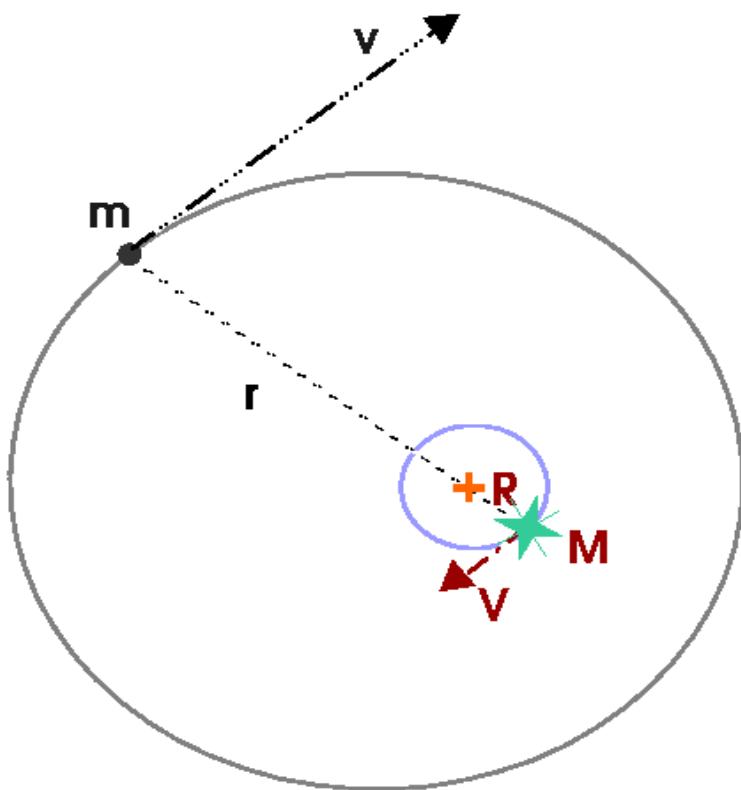
2026

PLATO (ESA + ... + Brasil)

Ø 12 cm telescope

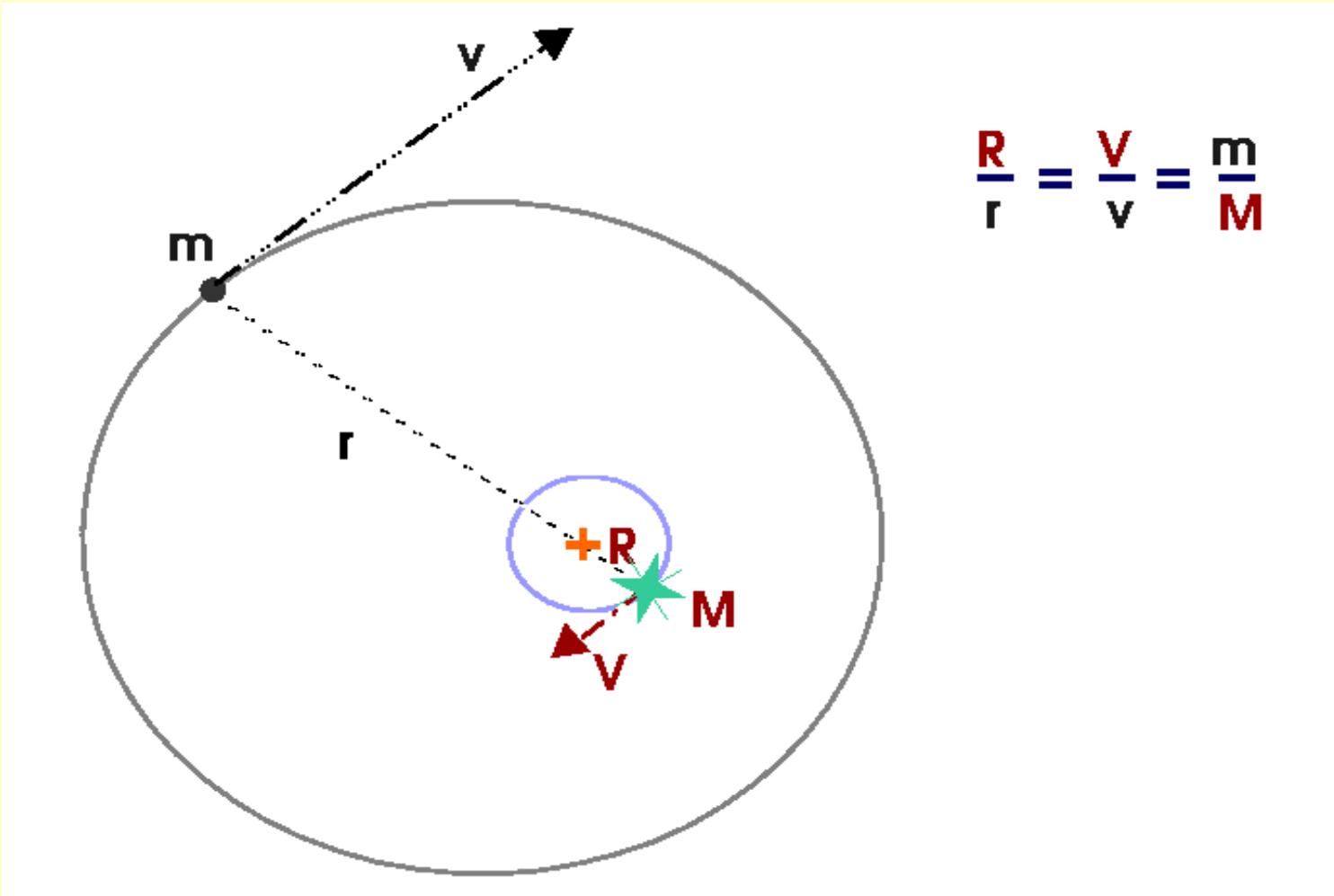
24 cameras

VELOCIDADES RADIAIS (Massas dos planetas)



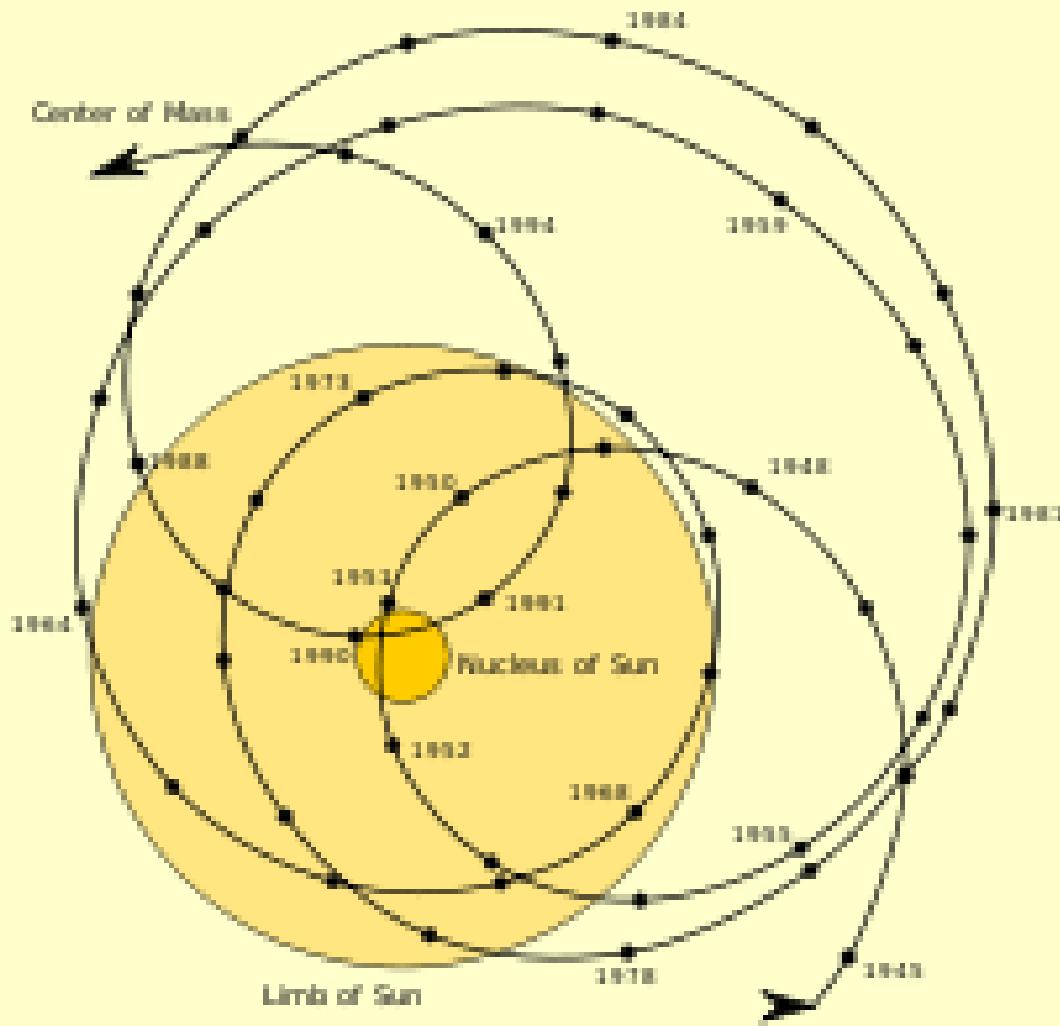
$$\frac{R}{r} = \frac{V}{v} = \frac{m}{M}$$

r, v, R, V são coordenadas e velocidades baricêtricas

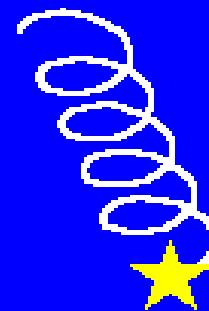
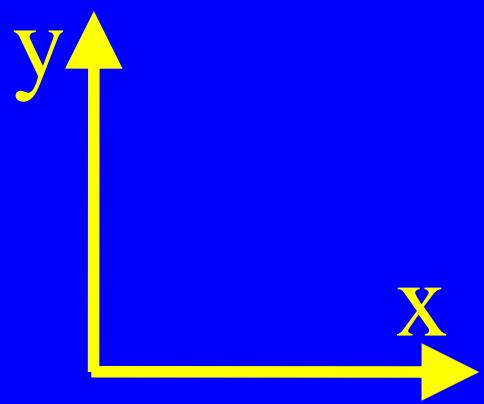


$$\frac{R}{r} = \frac{v}{v} = \frac{m}{M}$$

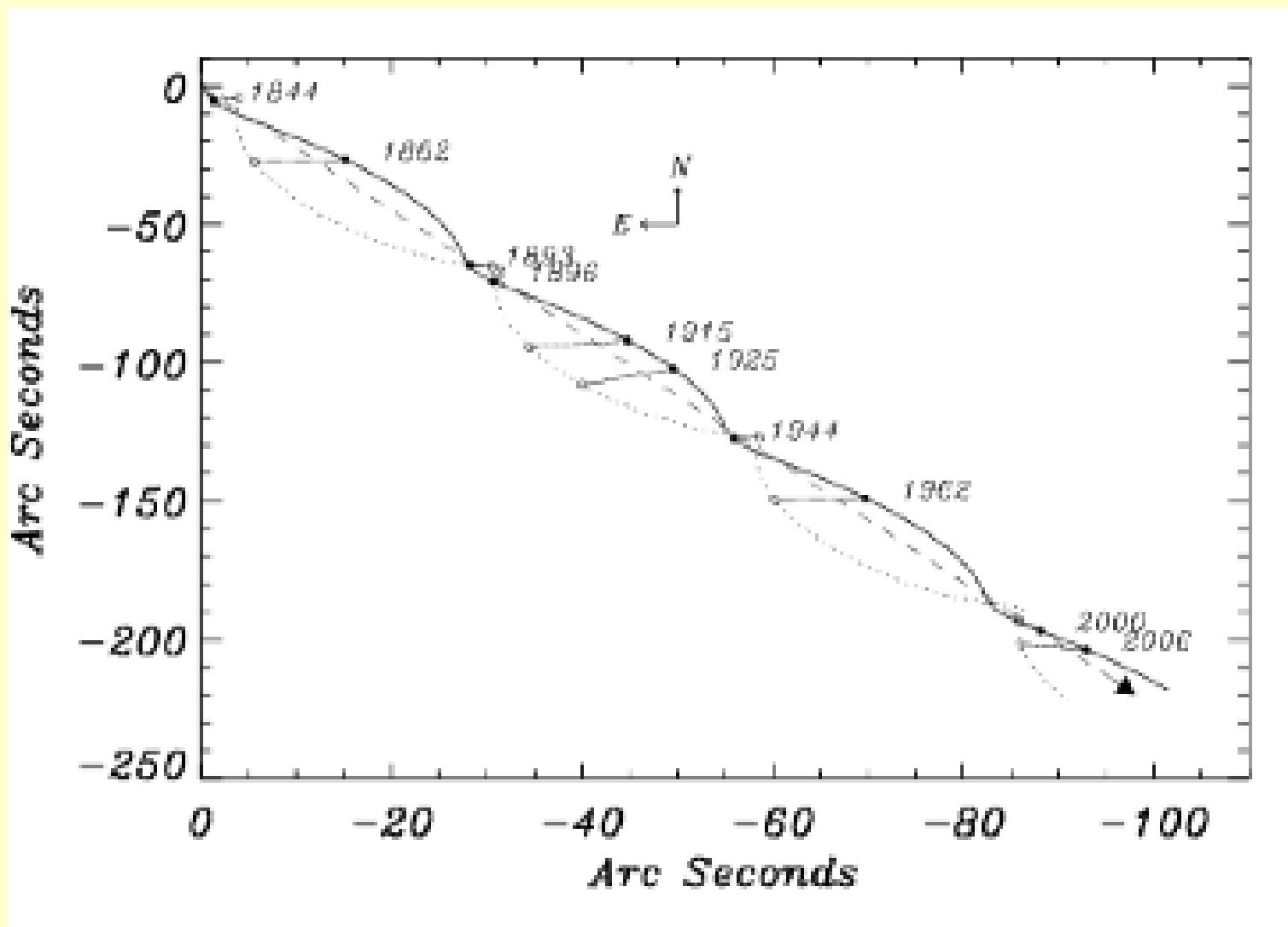
r, v, R, V são coordenadas e velocidades baricêtricas

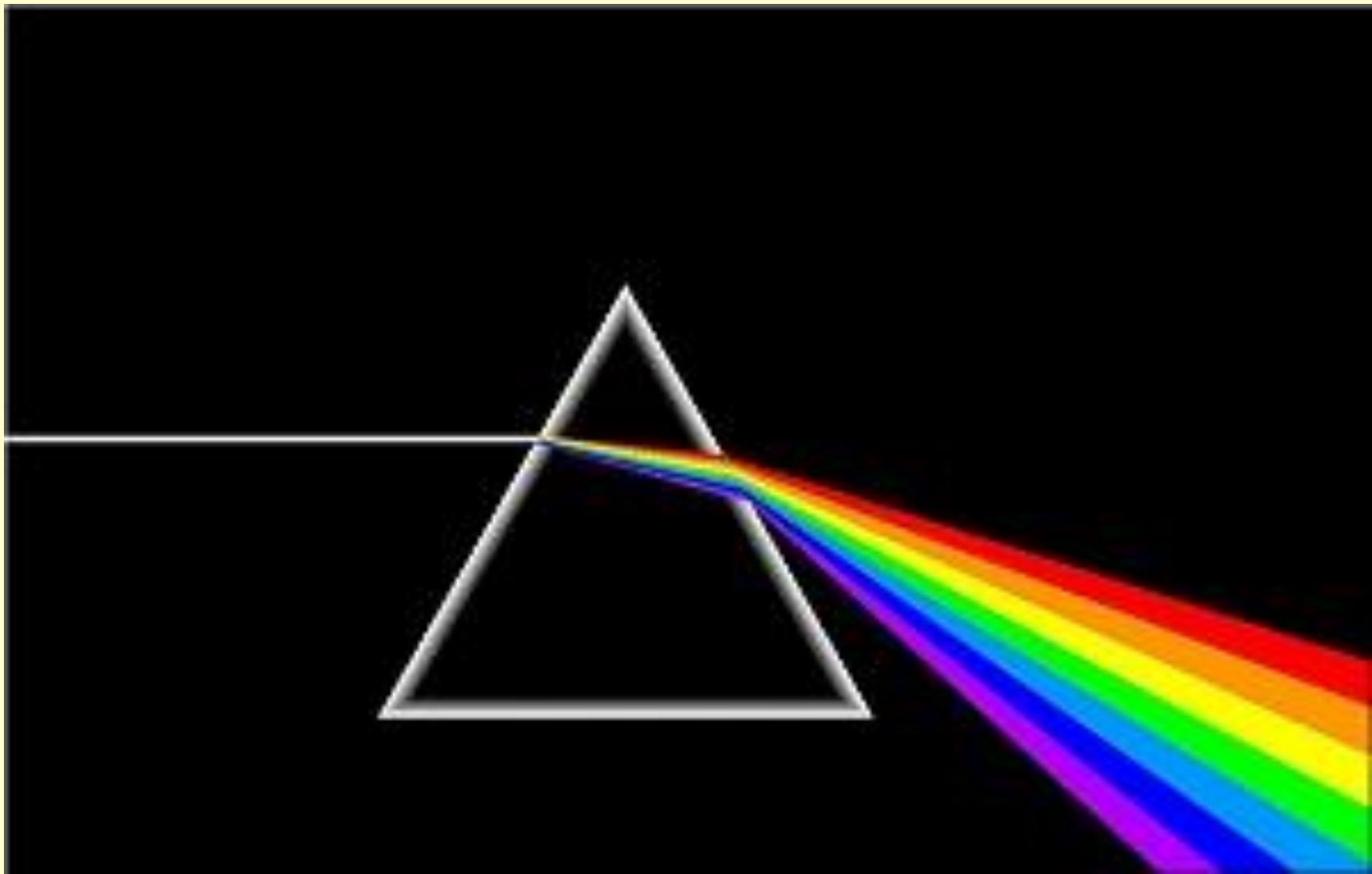


Movimento do Sol relativo ao centro do S.Solar

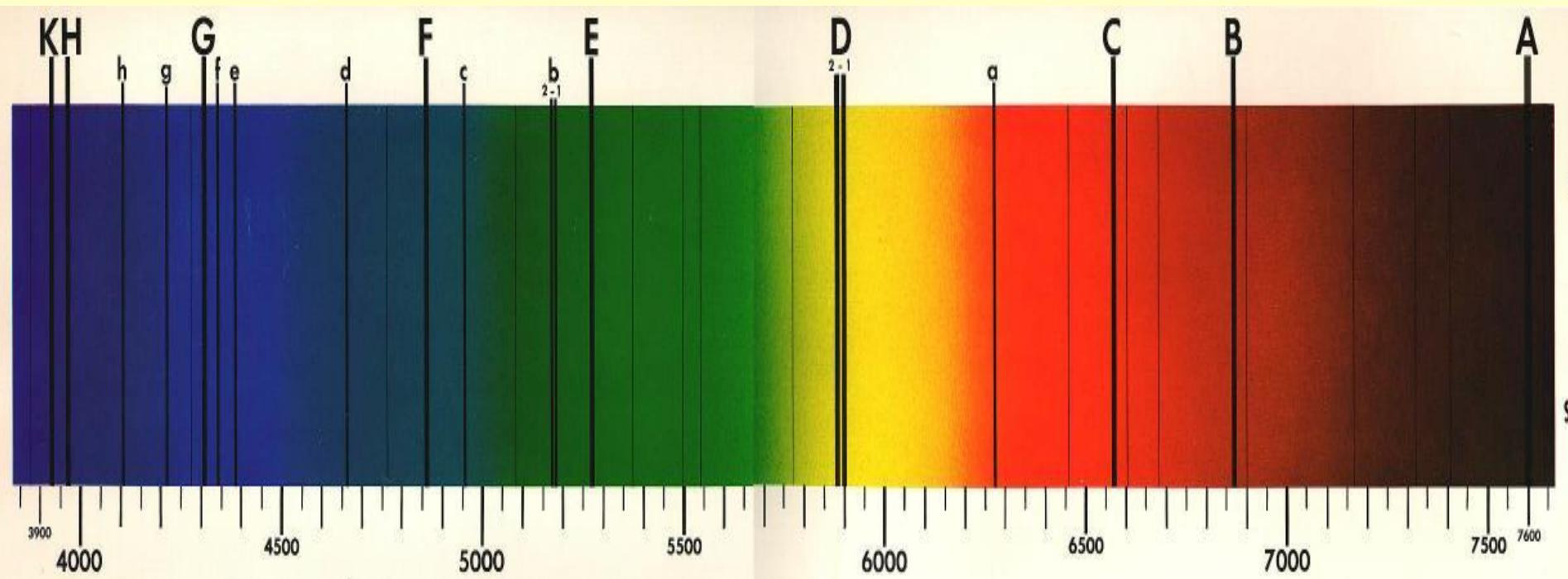


Sirius A+B

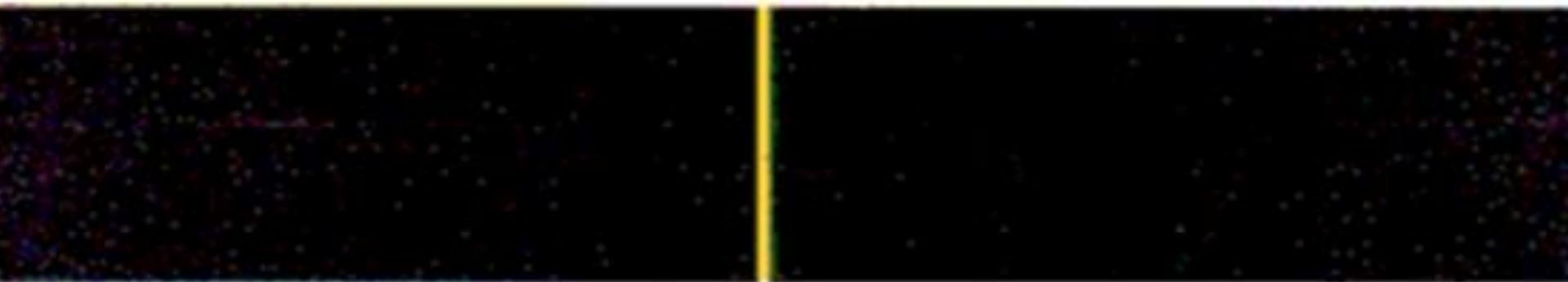


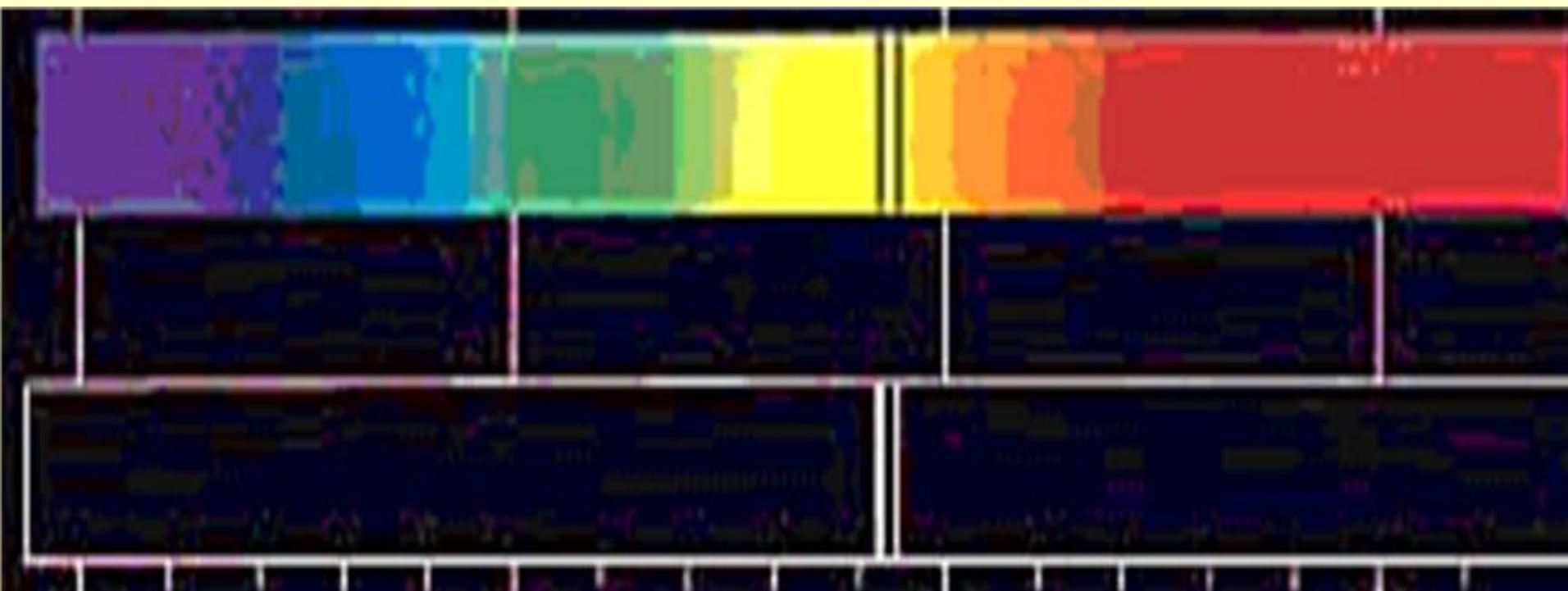


Espectro do Sol (Fraunhofer)

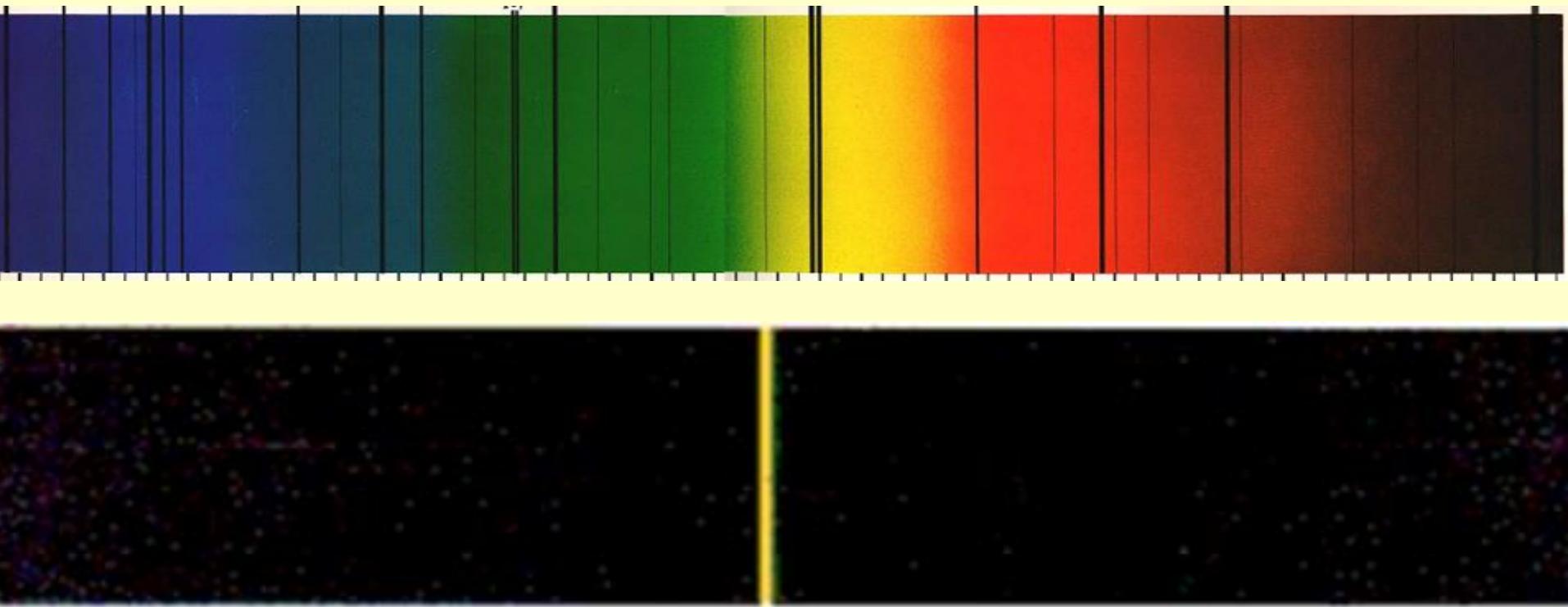


Lâmpada de Sódio





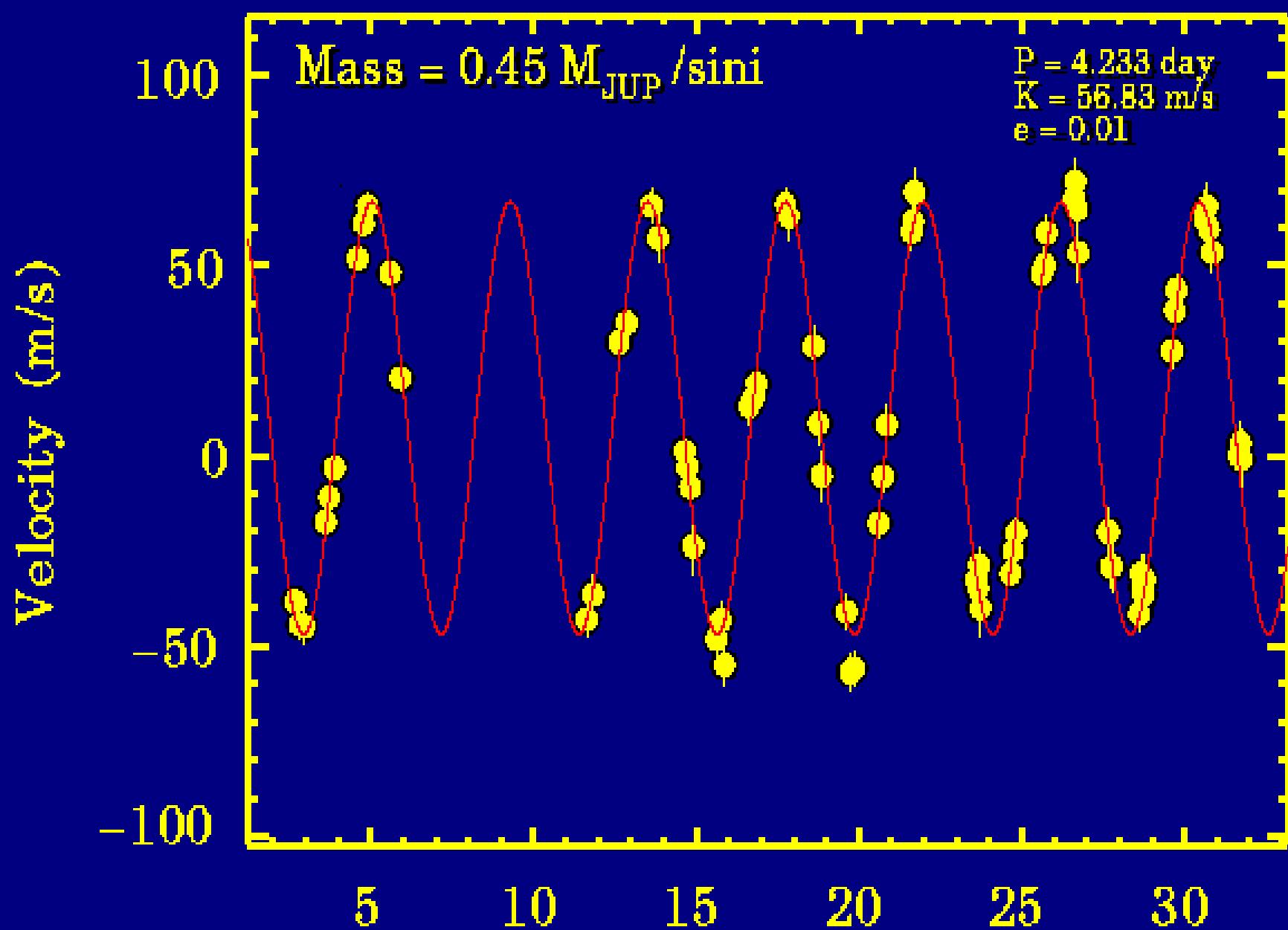
Efeito Doppler



Exemplos: Sol-Júpiter ~ 13 m/s
Sol-Terra ~ 9 cm/s

51 Pegasi

Marcy & Butler





Harps at ESO (Chile)



The Periodic Table of Exoplanets

Over 3800 Exoplanets



Terrans Group

Miniterrans (Mercury Size)	Subterrans (Mars Size)
$10^{-5} - 0.1 M_E$ or $0.03 - 0.4 R_E$	$0.1 - 0.5 M_E$ or $0.4 - 0.8 R_E$

Terrans (Earth Size)	Superterrans (Super-Earths & Mini-Neptunes)
$0.5 - 5 M_E$ or $0.8 - 1.5 R_E$	$5 - 10 M_E$ or $1.5 - 2.5 R_E$

Giants Group

Neptunians (Neptune Size)	Jovians (Jupiter Size)
$10 - 50 M_E$ or $2.5 - 6.0 R_E$	$> 50 M_E$ or $> 6 R_E$

Hot Zone



Potentially Habitable



Cold Zone



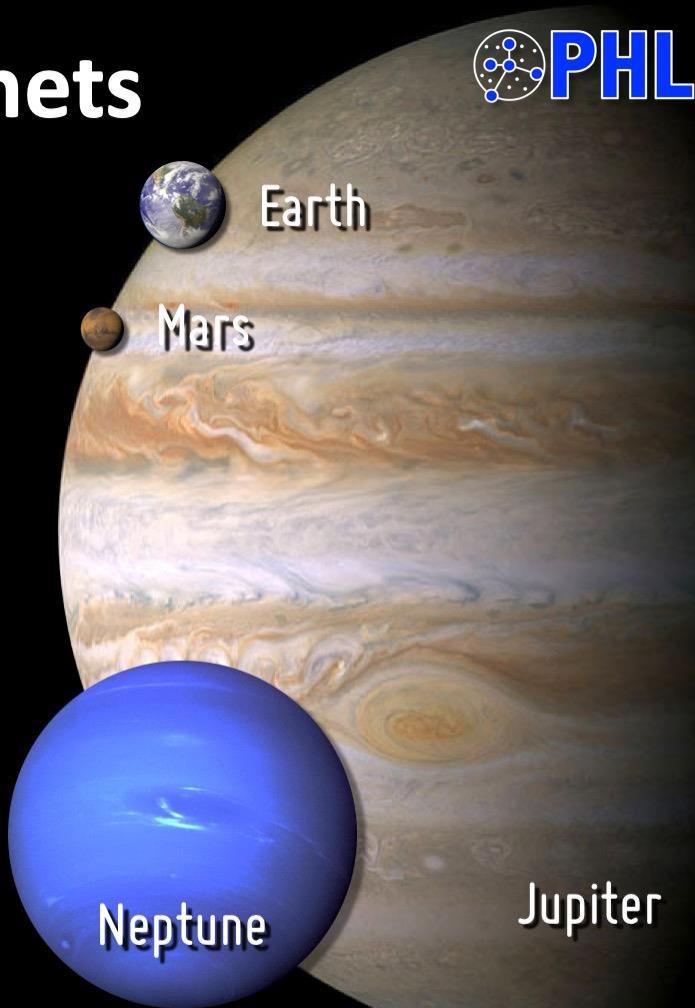
M_E = Earth Mass, R_E = Earth Radius

CREDIT: PHL @ UPR Arecibo (phl.upr.edu) Jul 2018

(2018)

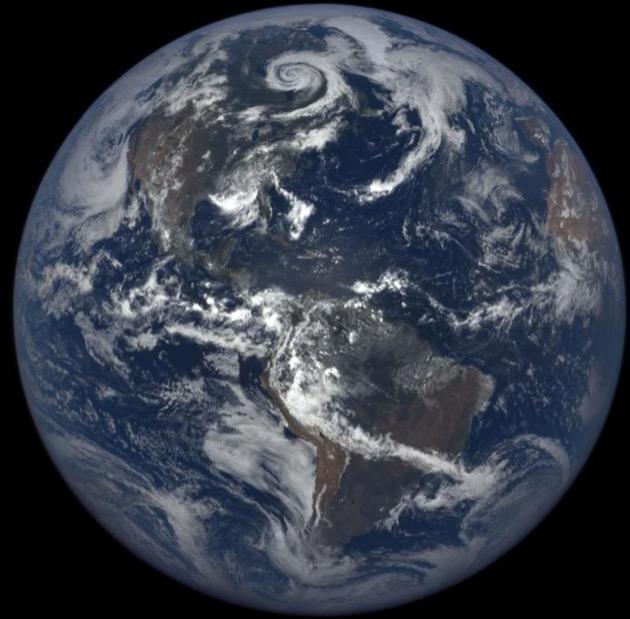
Potentially Habitable Exoplanets

Ranked by Distance from Earth (light years)



Artistic representations. Earth, Mars, Jupiter, and Neptune for scale. Distance from Earth is between brackets.

CREDIT: PHL @ UPR Arecibo (phl.upr.edu) Feb 1, 2019



Earth



Proxima b
(artistic representation)

Credit: PHL @ UPR Arecibo, NASA EPIC Team

Three Potentially Habitable Worlds Around Gliese 667C



Gliese 667C c

Gliese 667C f

Gliese 667C e

Earth

CREDIT: PHL @ UPR-Brownsville, NASA

TRAPPIST-1 System

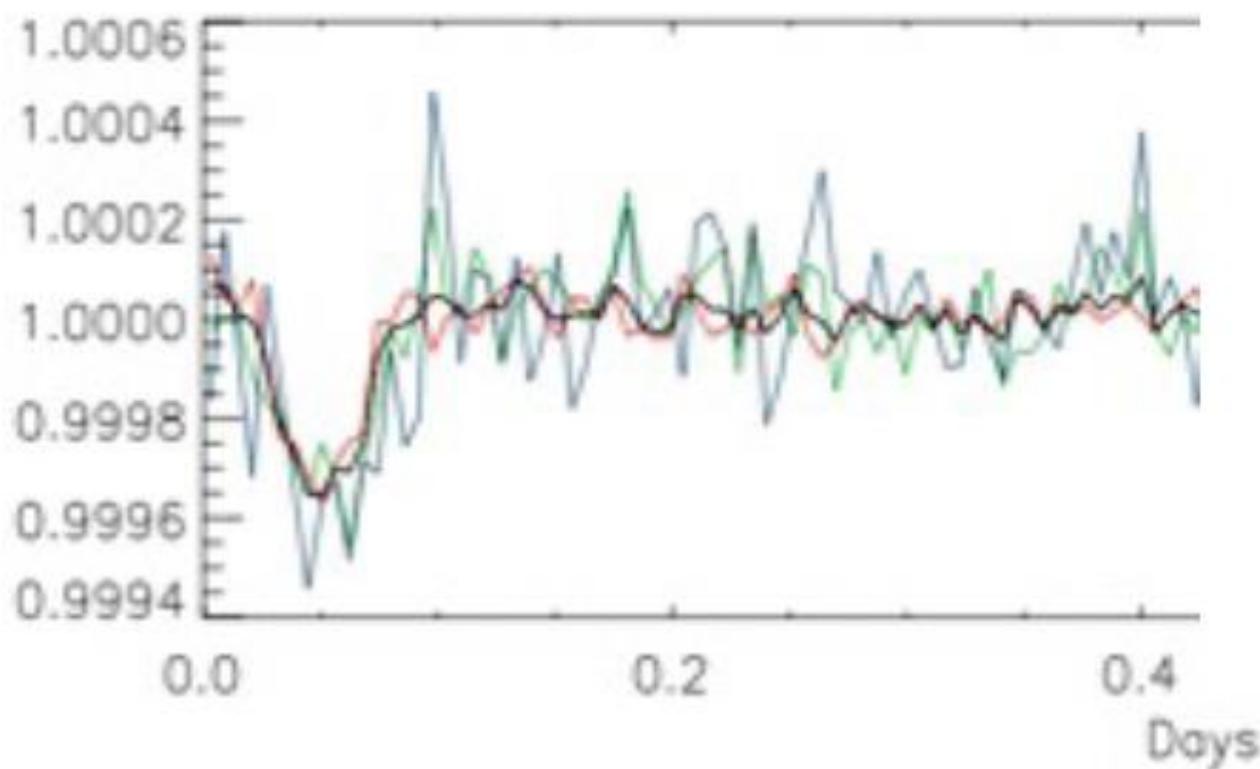
500 hours of Spitzer data have revealed extensive details about these planets, validating and expanding greatly upon the initial discovery by the TRAPPIST telescope in Chile



FIM

CoRoT-7b

Measured



Corot - 7B

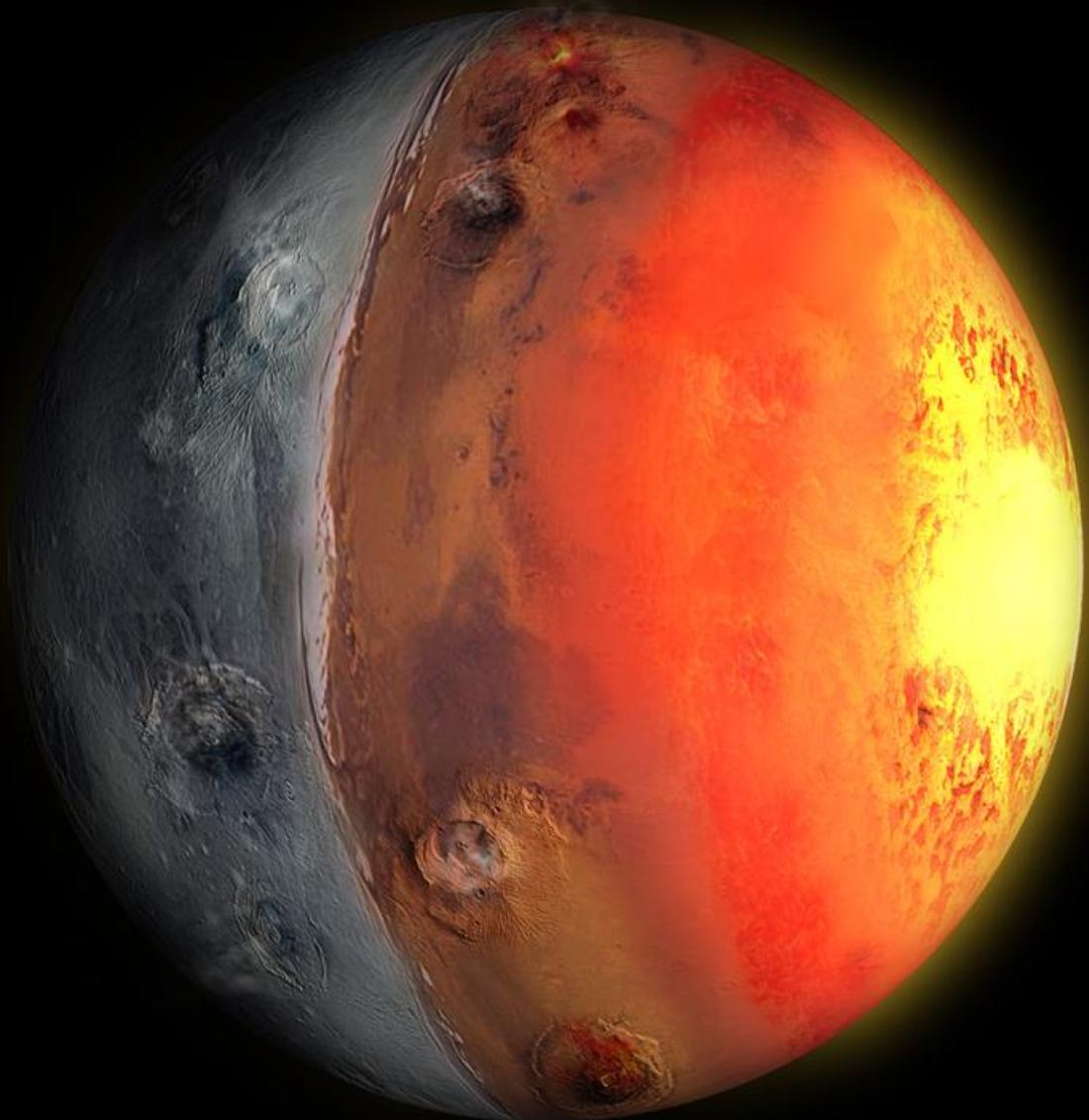
$m = 5T$

$P = 0.85 \text{ d}$

$a = 2.5 \text{ milhoes km}$

$R = 11000 \text{ km}$

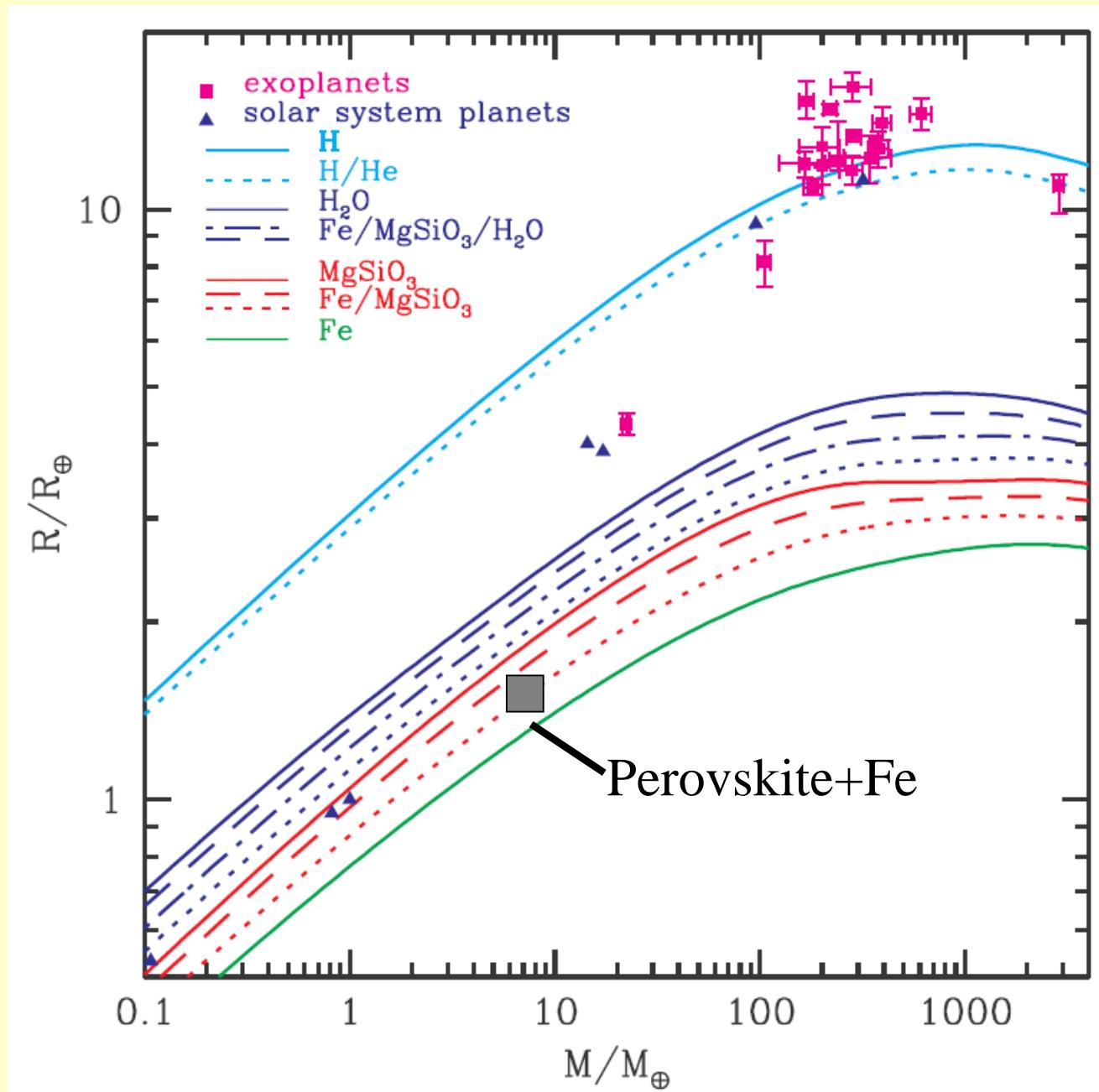
**(maré = +65/
-20 km)**

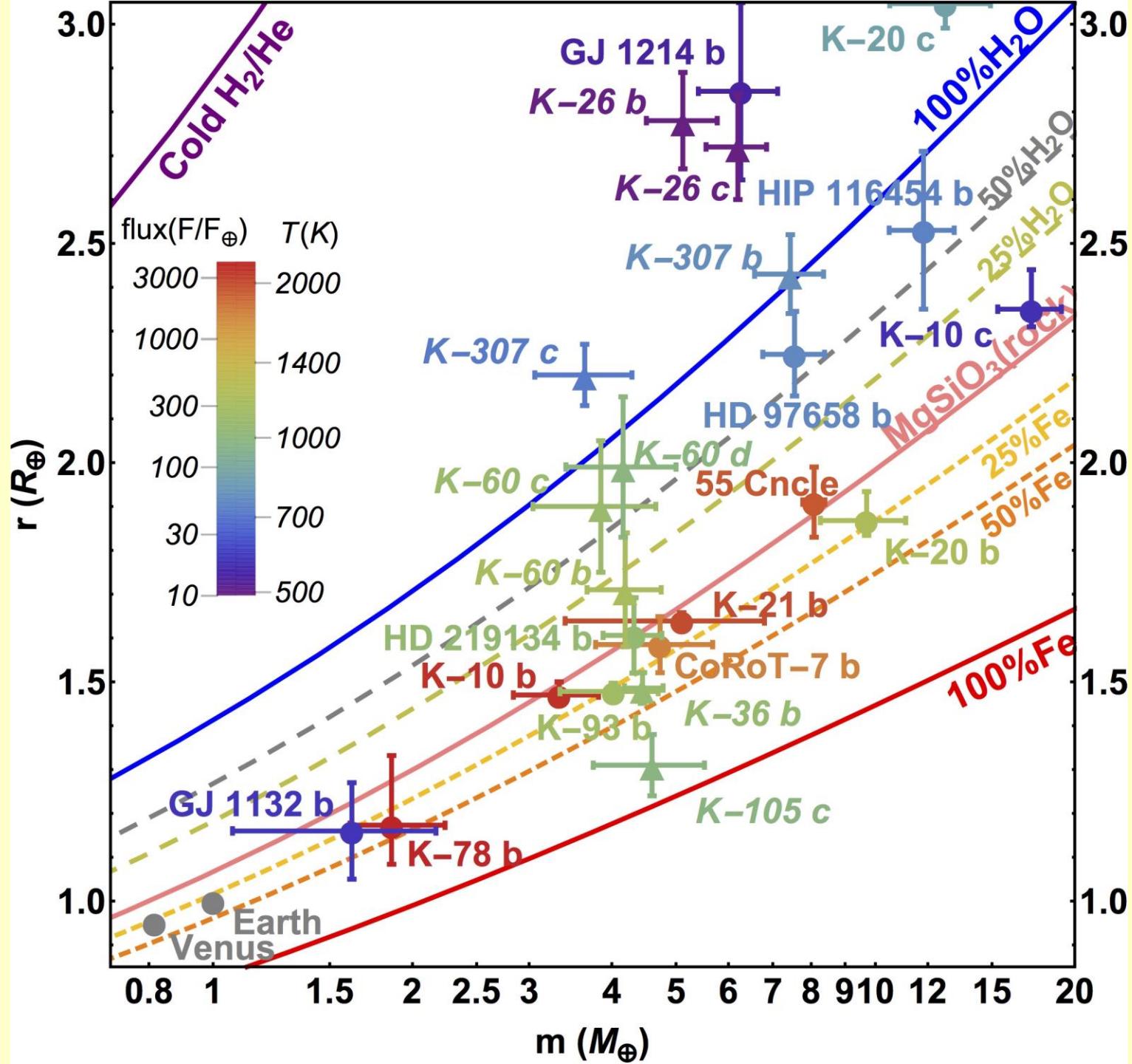


Ref: CoRoT + IAG 2010

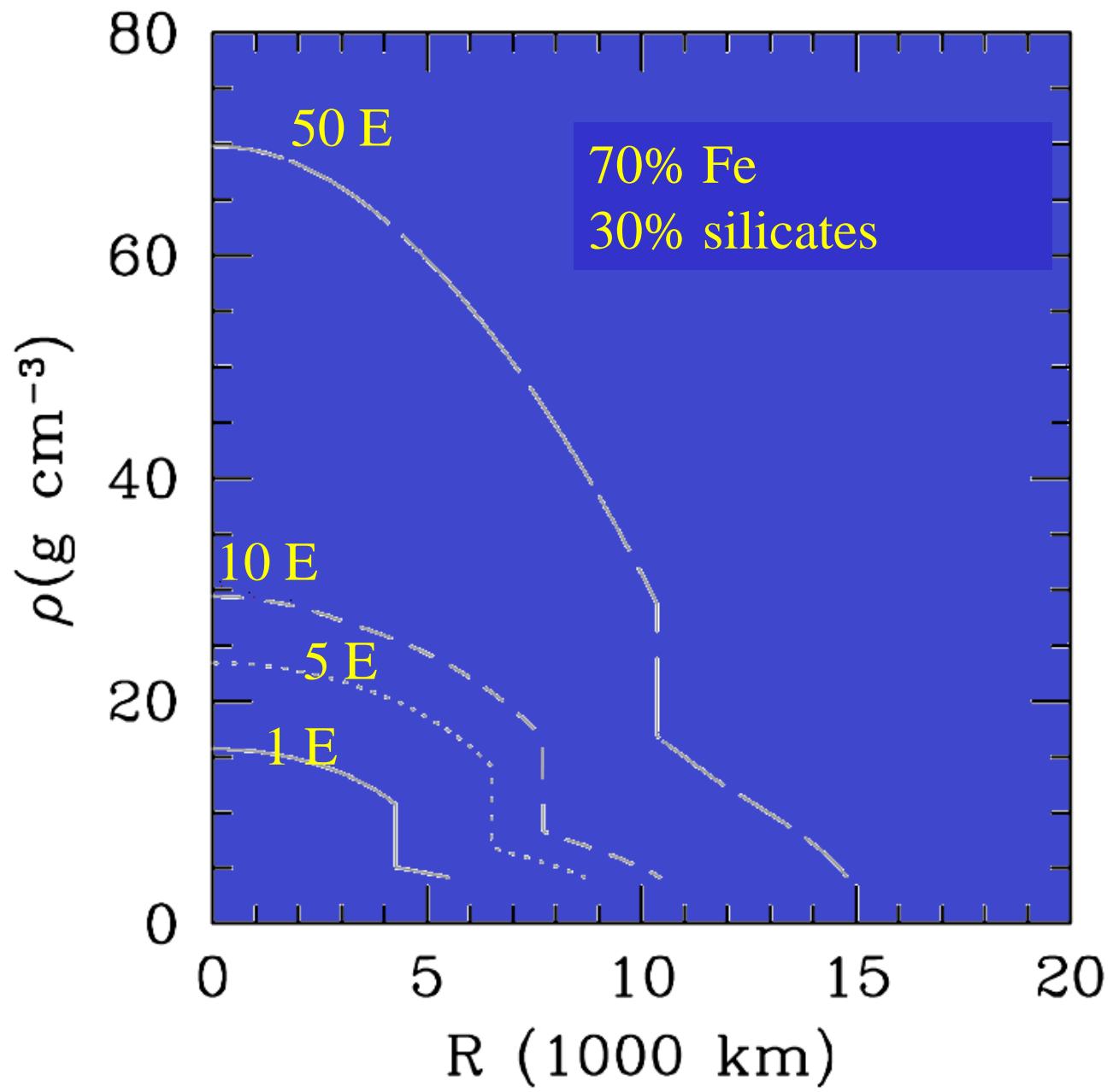
Seager
et al.
ApJ 669,
2007

ABO_3

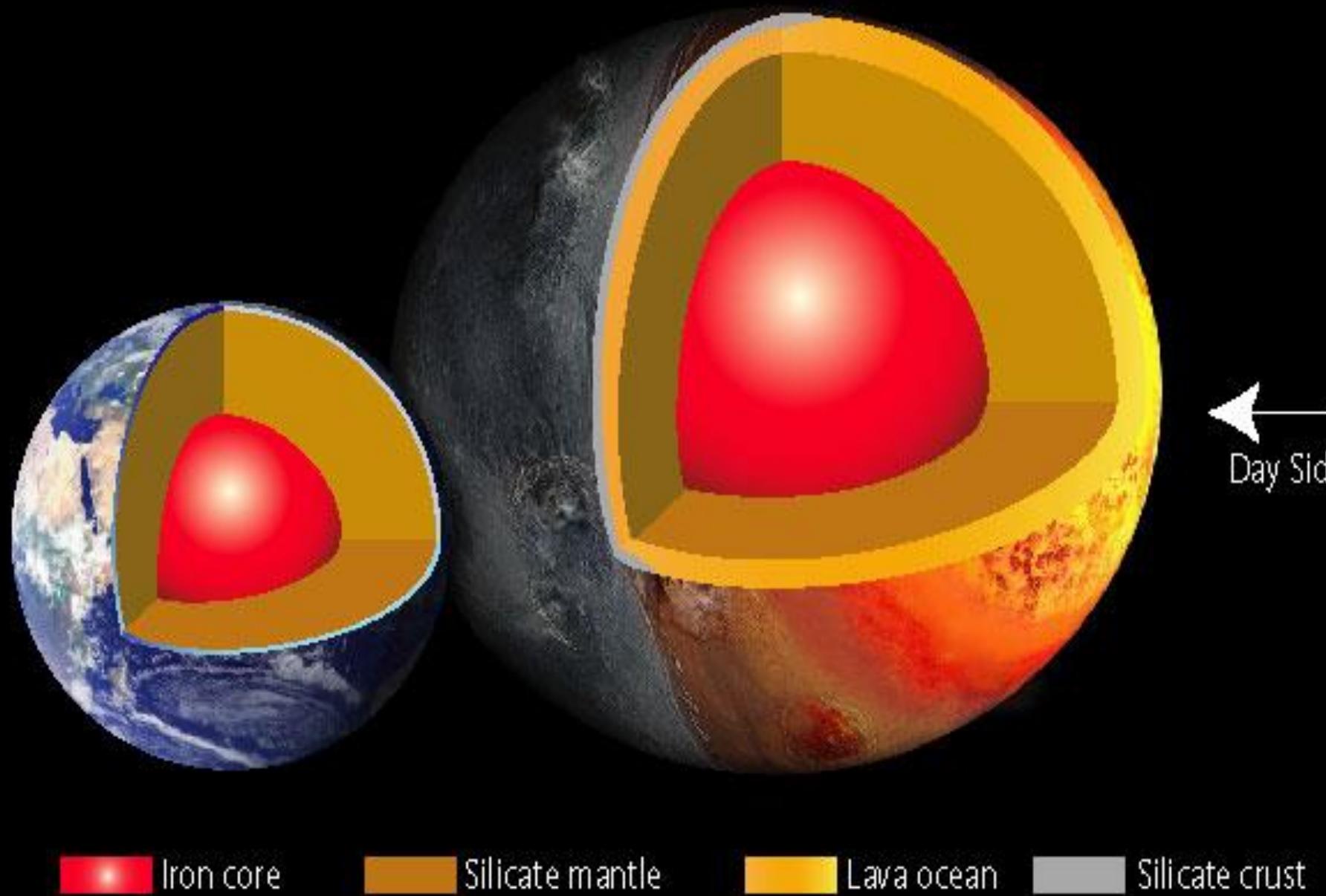


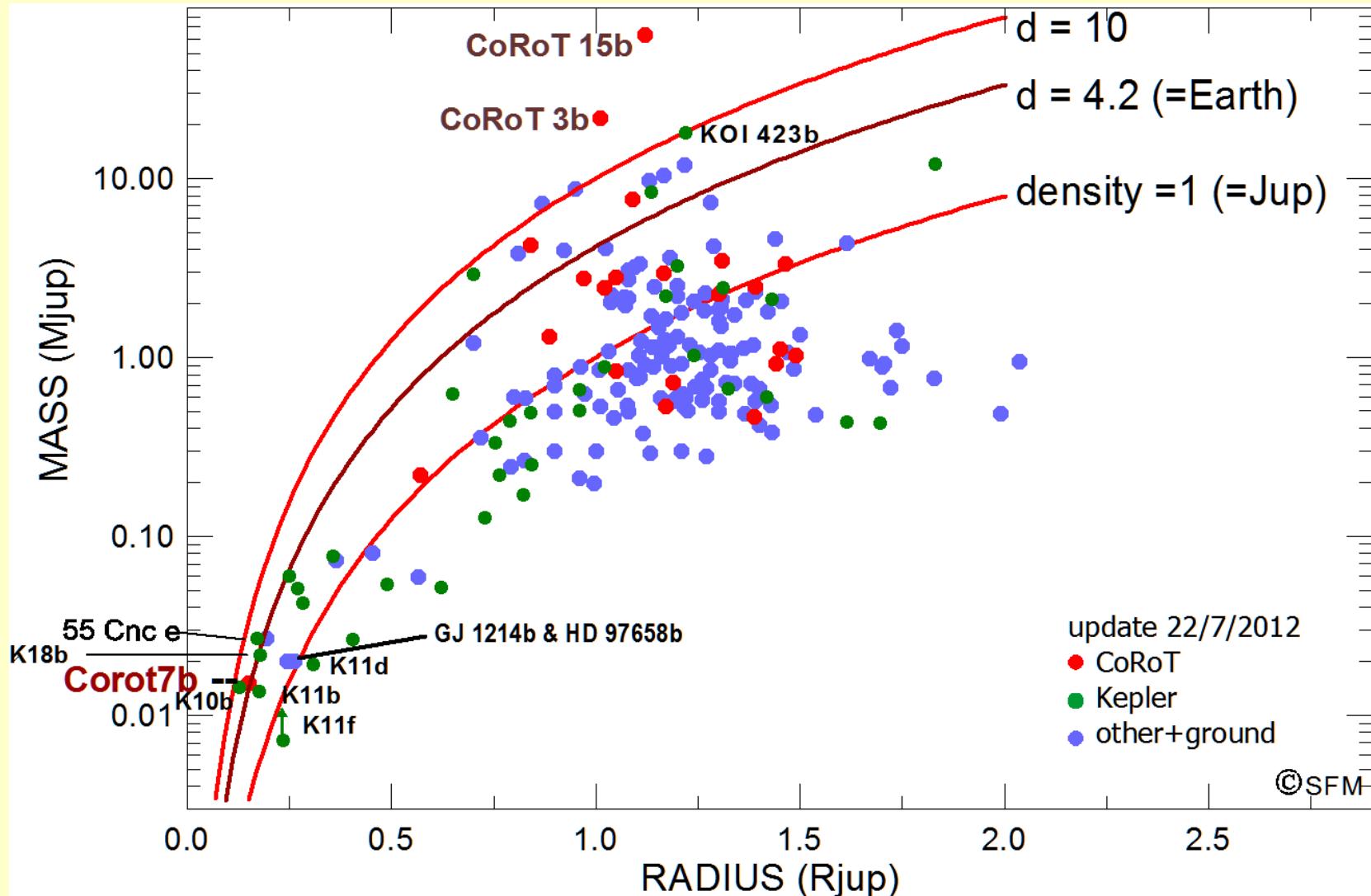


Seager
id.



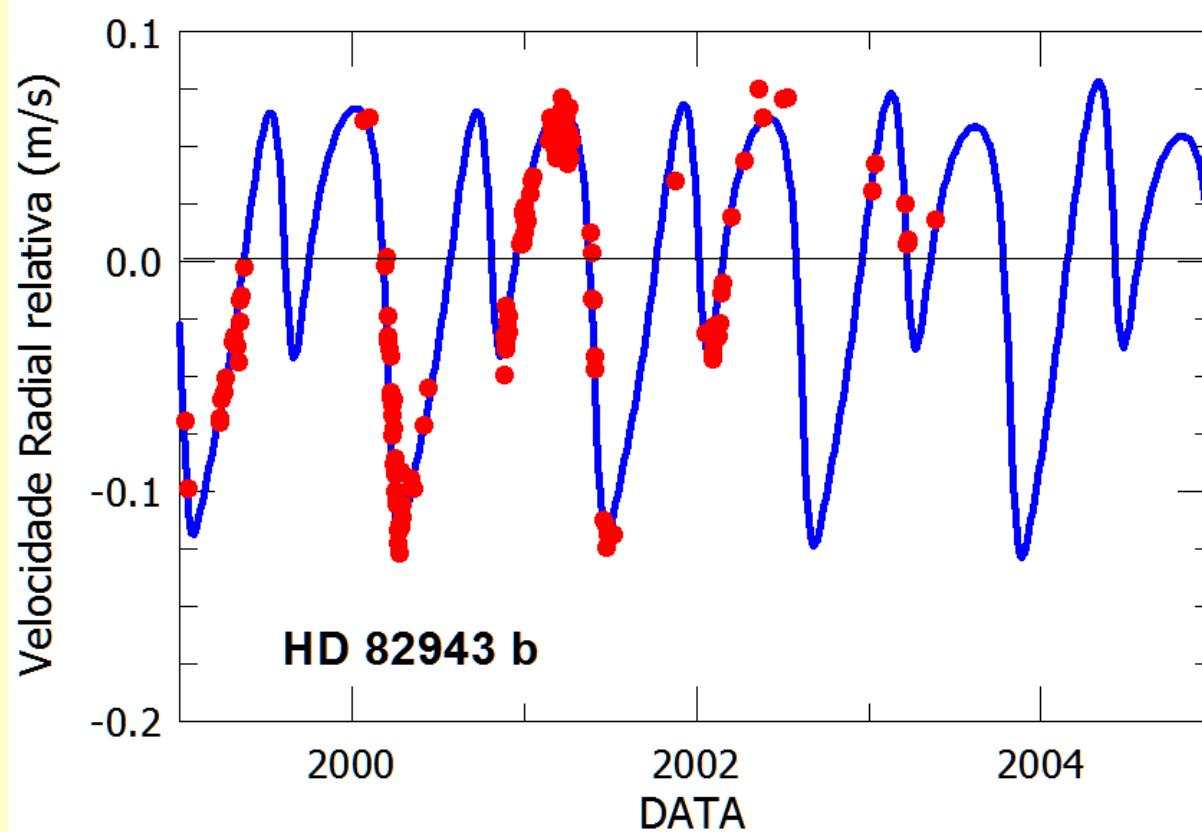
Possible internal structure (compared to Earth)



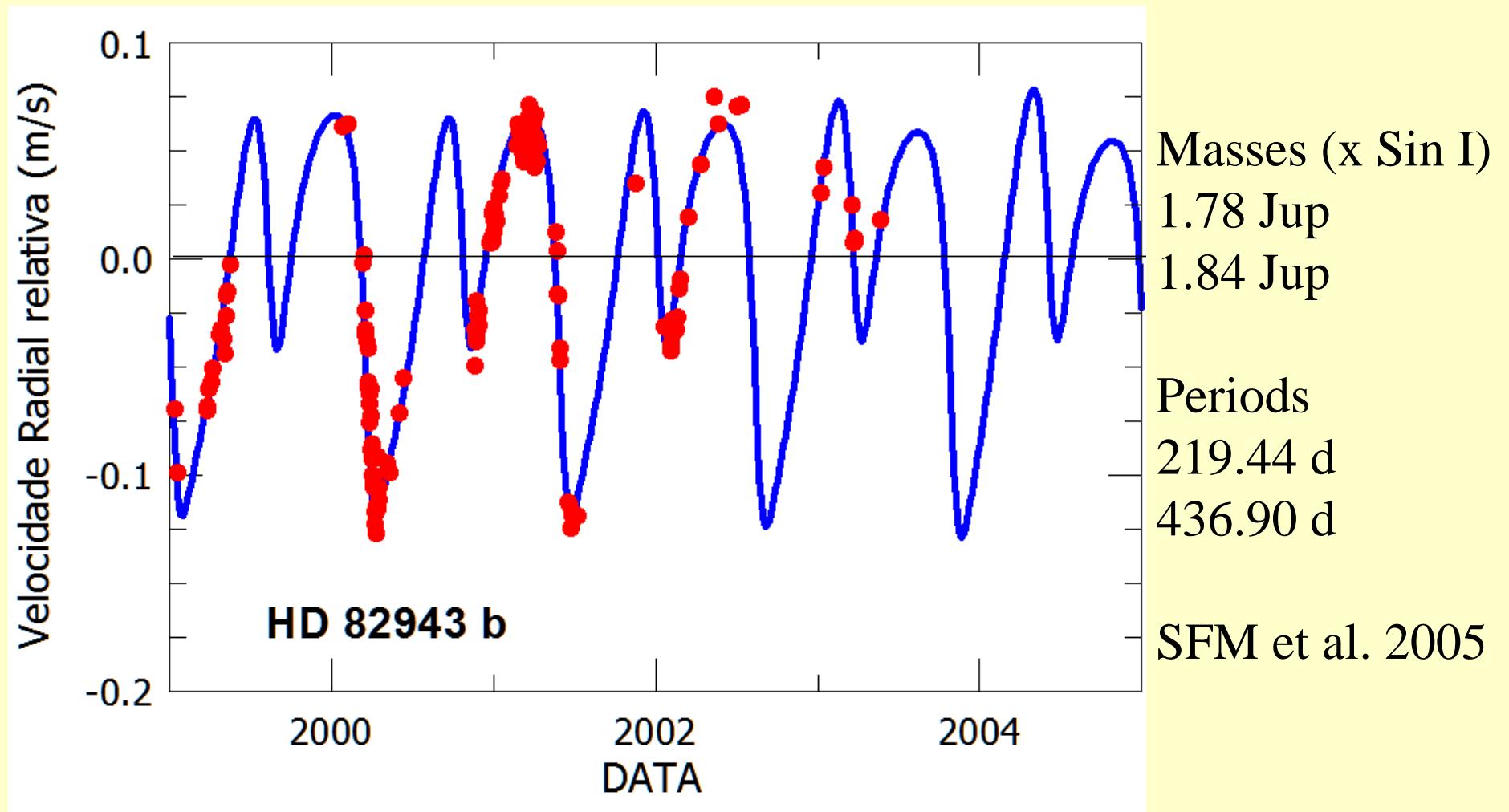


Relação Massa – Raio
Linhos de densidade ($1=Jup$)

SISTEMAS COM MÚLTIPLOS PLANETAS



Radial velocity of a star with two planets (variation)



OS PLANETAS DE UPSILON ANDROMEDA

MASSAS

[unidade: Júpiter]

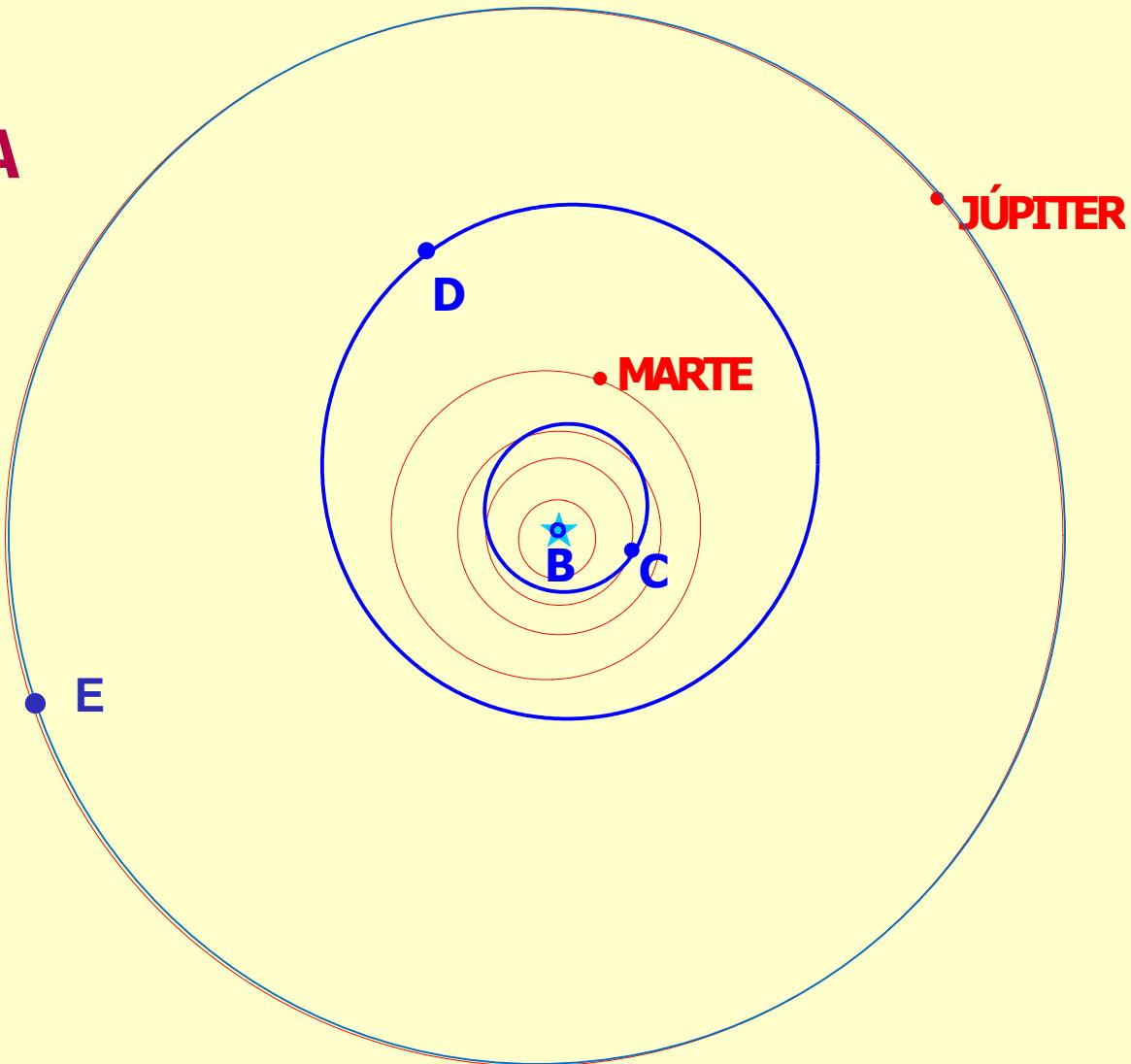
- (B) ~~0.69~~ 5.9
- (C) ~~0.98~~ 14.57
- (D) ~~3.95~~ 10.19
- (E) 1.06 ?

Periodos: 4.6 d

241 d

3a + 194 d

10a+196 d



Inclinações: 7-15°

S.Barros et al.
2014

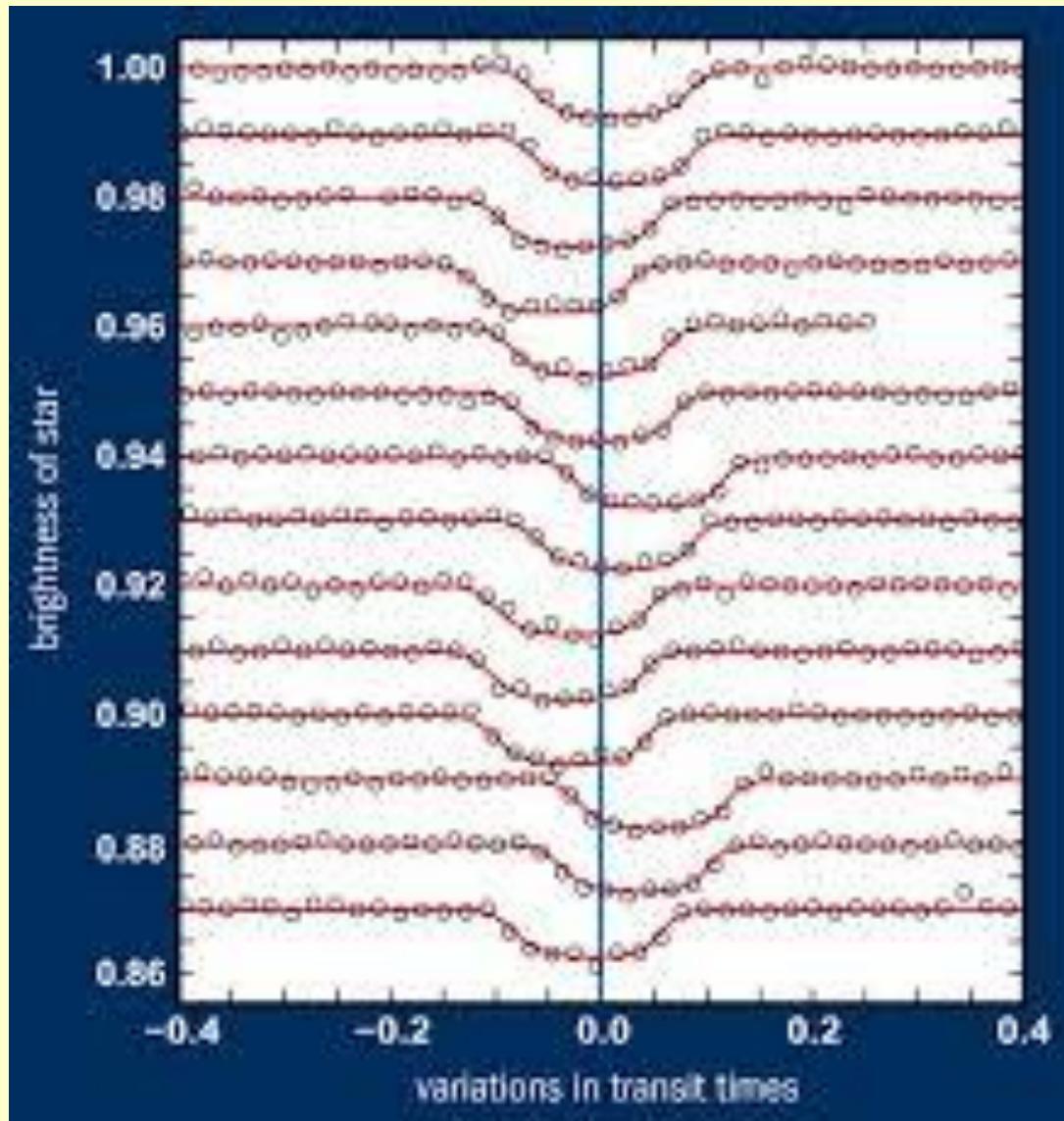
$$RV = \text{proj}_{\text{linha de visada}} \vec{V} = \mathbf{K}[\cos(v + \omega) + e \cos \omega]$$

$$\mathbf{K} = \frac{m}{m + M} \frac{na \sin i}{\sqrt{1 - e^2}}$$

Exemplos: Sol-Júpiter ~ 13 m/s
Sol-Terra ~ 9 cm/s

Não é possível separar os valores de
M, m e sin i.

TTV = Transit time variations



Super-terrás

A zona habitável

H. Rauer, DLR, 2014-2-24 (based on exoplanet.eu)

