The *RISE* of r-process in dwarf galaxies

mostly about arXiv: 2308.13702, but also arXiv: 2404.10067

IAU Symposium 2024 – Paraty November 2024

Guilherme Limberg – KICP postdoc fellow at UChicago

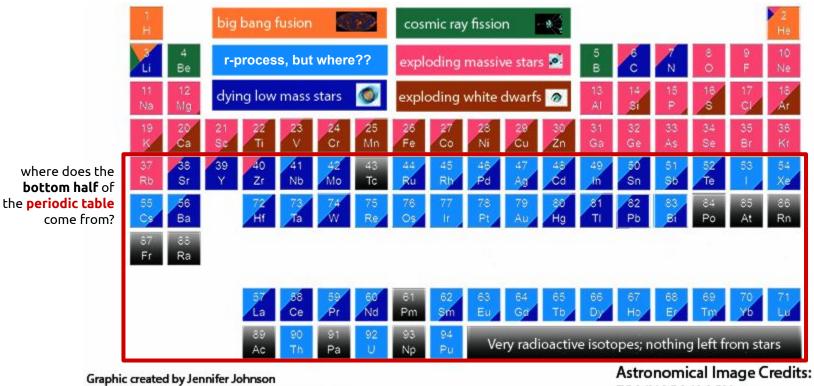
f guilherme.limberg@usp.br



@guilimberg



All "metals" are formed in processes associated with stars. But what about the **heaviest elements**?



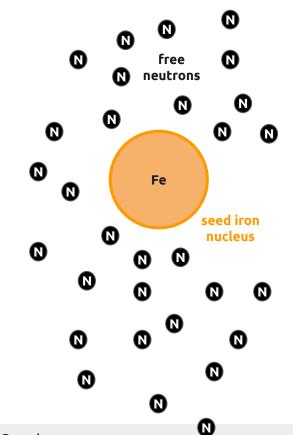
http://www.astronomy.ohio-state.edu/~jaj/nucleo/

ESA/NASA/AASNova

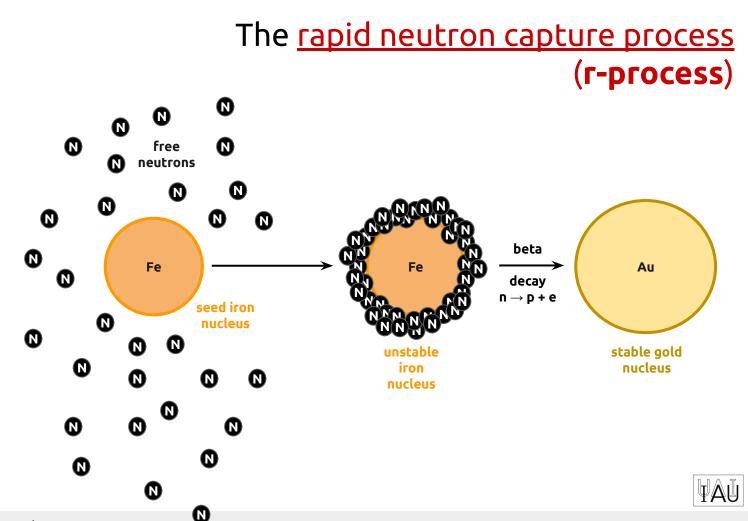


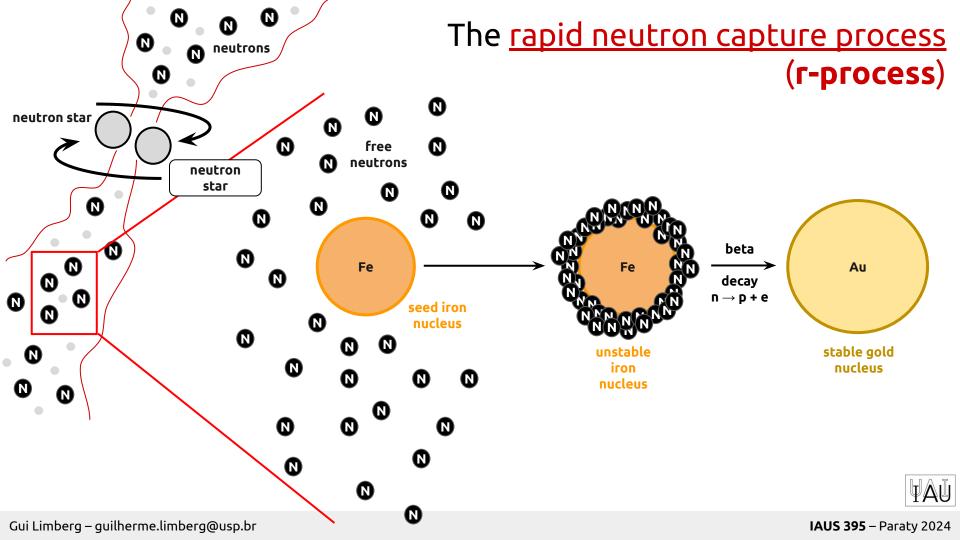
The rapid neutron capture process

(r-process)

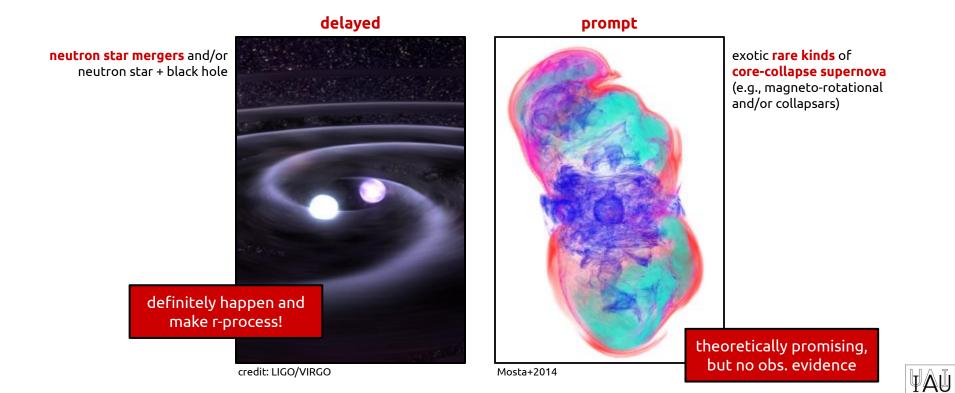








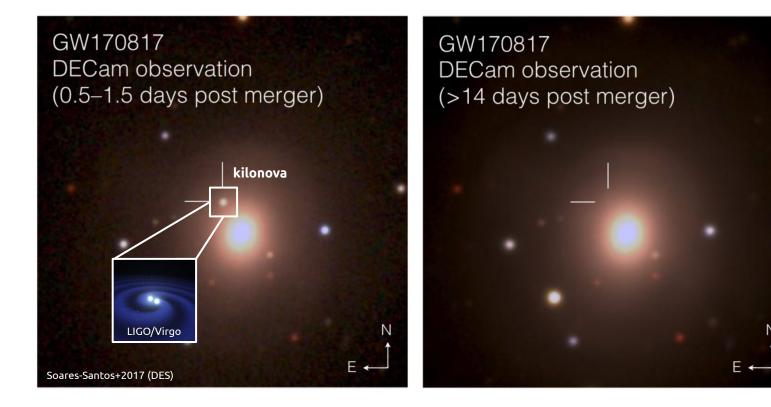
Viable **r-process** site(s) are **rare** and **prolific**!



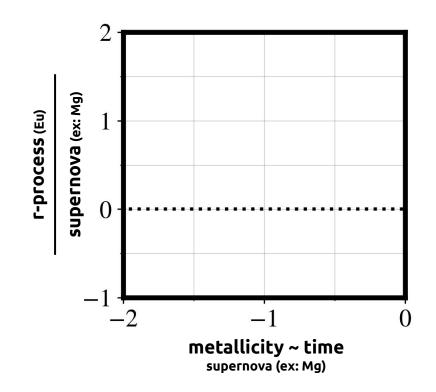
Gui Limberg – guilherme.limberg@usp.br

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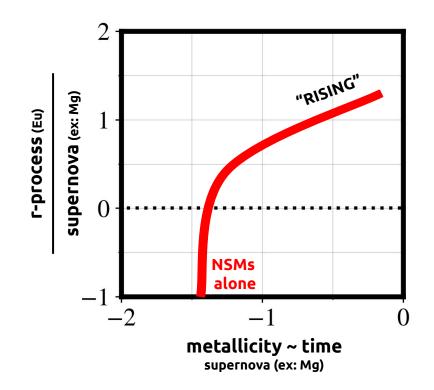
Multimessenger observations of GW170817 confirmed neutron star mergers as a site for r-process



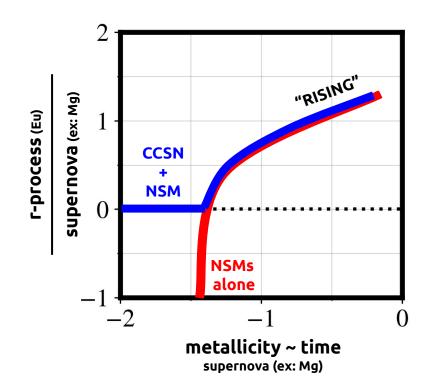




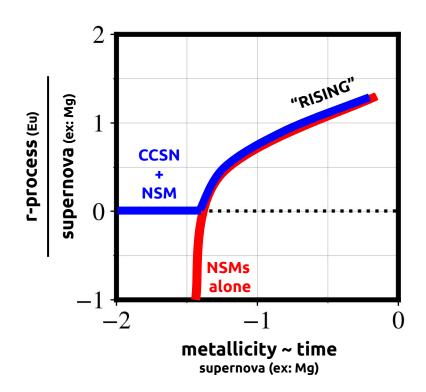












Why don't we just do it?

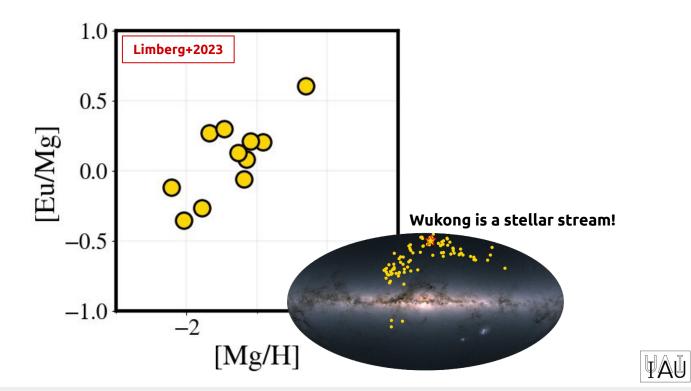
- high-res. spec.
- metal-poor
- real member
- cool (*T*eff < 5000K)
- bright (*g* < 18)



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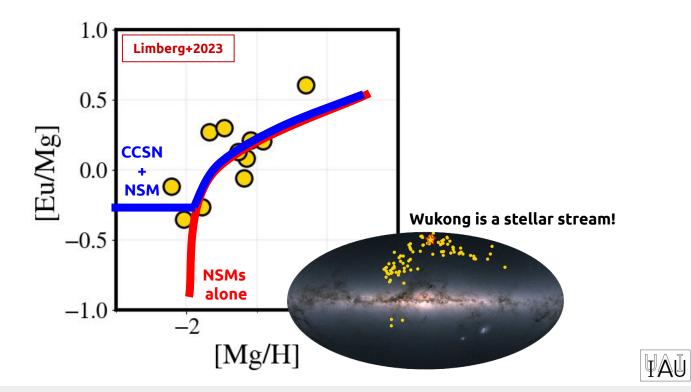
The first dwarf galaxy with **r-process** enrichment clearly from **delayed sources**

(presumably neutron star mergers)



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Takeaway messages and summary

- (G)galactic Archaeology to unveil the origin of chemical elements
 - stars preserve the compositions of their natal gas clouds
- Sources of r-process must be rare and prolific
 - Delayed and prompt play their role, but what are they?
 - Delayed = neutron star mergers (GW170817) ✓
 - Prompt = core collapse or fast neutron star mergers?
- Massive disrupted dwarfs: Wukong

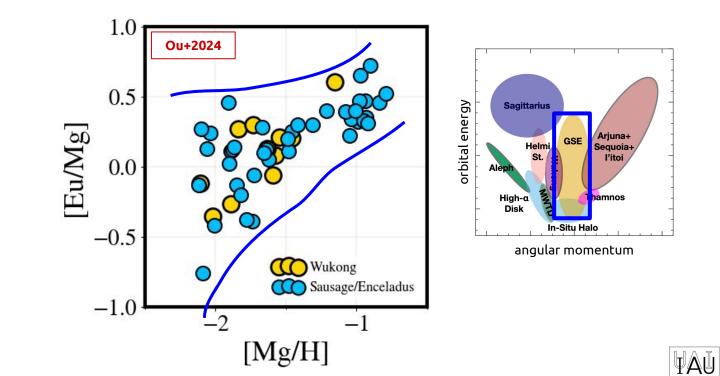
guilherme.limberg@usp.br

@guilimberg

- Delayed source(s) dominate r-process chemical evolution
- The future: tackle this with other disrupted dwarfs of various masses
 We need to go very metal poor to disentangle neutron star mergers from core-collapse supernovae

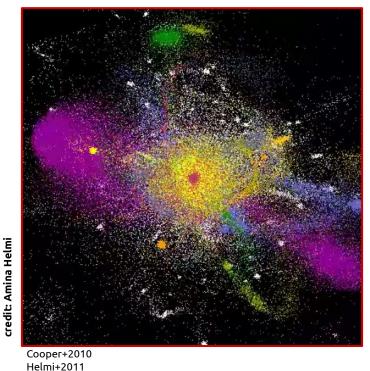


The first dwarf galaxy with **r-process** enrichment clearly from **delayed sources**



Disrupted dwarfs to the rescue? The **Milky Way**'s **halo** contains the debris of many accreted dwarfs

simulated halo* from Aquarius project

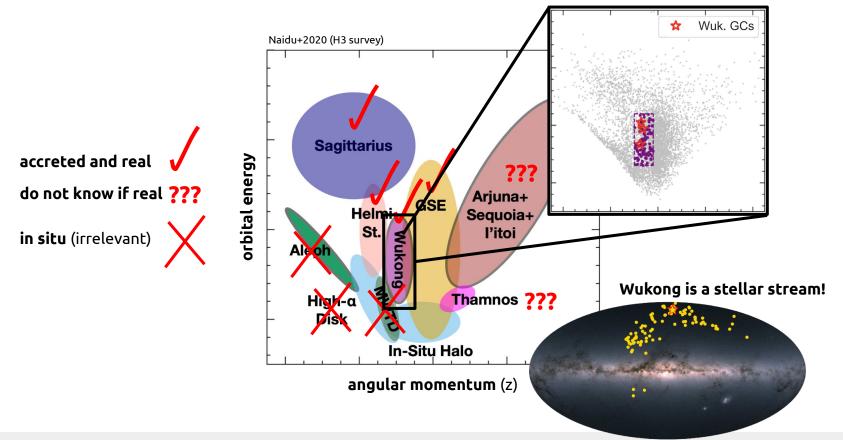




Gui Limberg – guilherme.limberg@usp.br

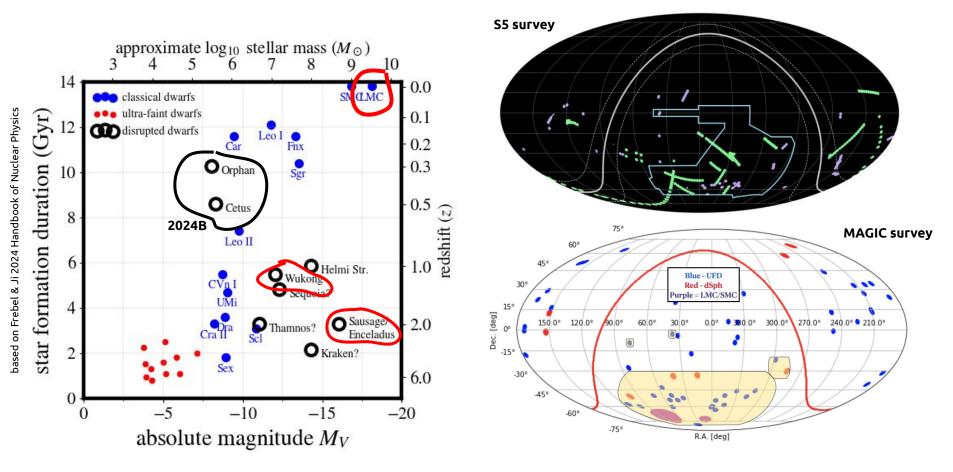
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Gaia mission provides the astrometric information needed

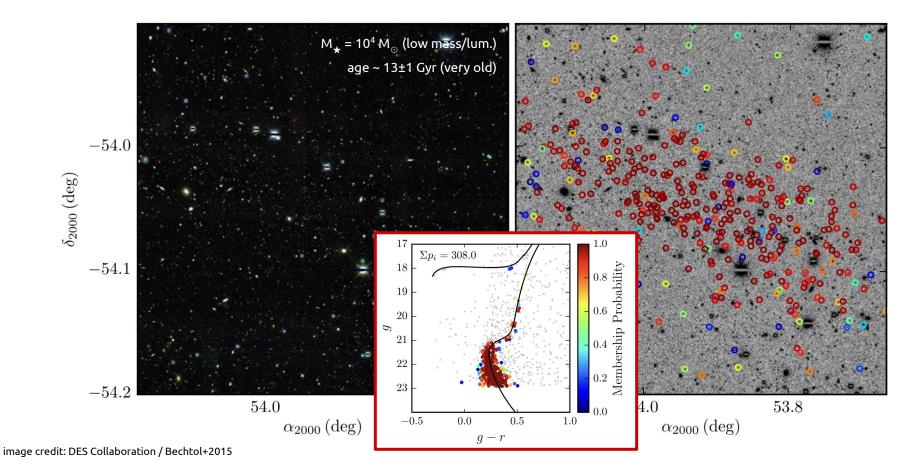


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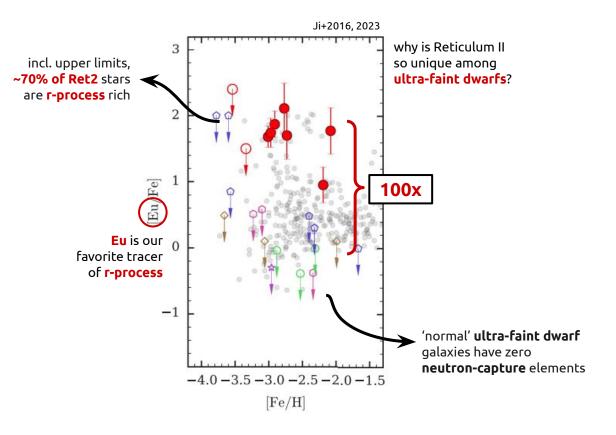
What should we do moving forward?



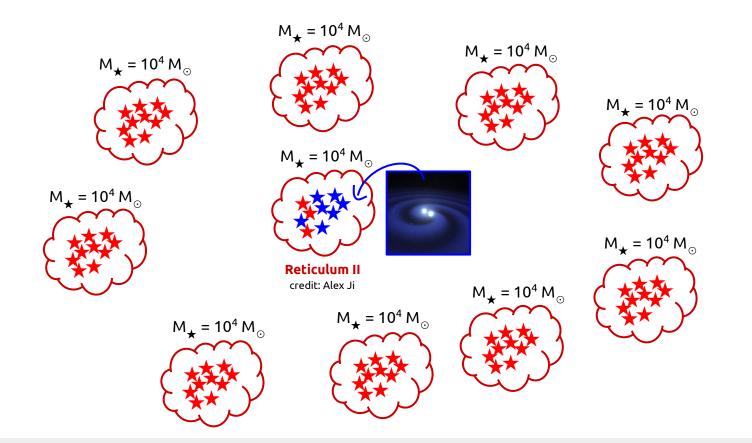
Ultra-faint dwarf **Reticulum II** was enriched in r-process by a single, rare, and prolific event...



... and **Reticulum II** is immensely **r-process** rich. Why?

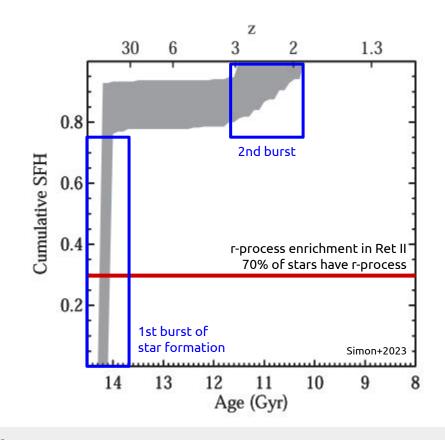


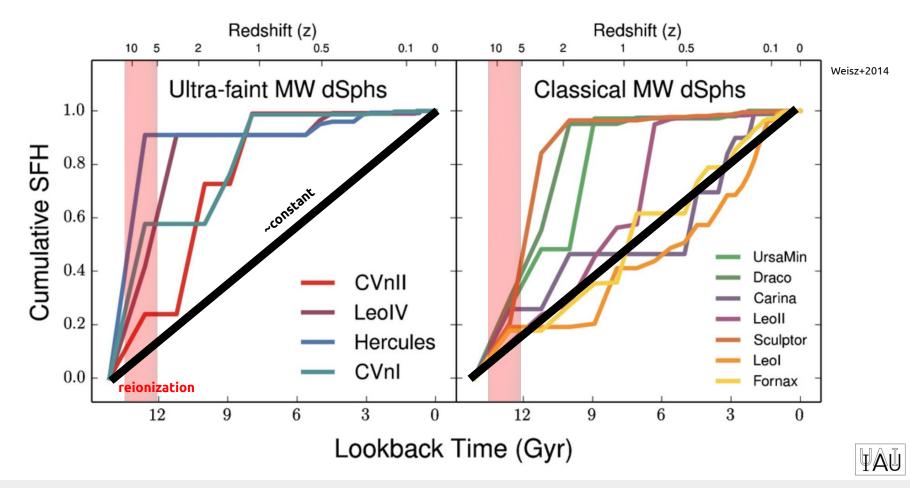
Reticulum II was perhaps just very (un)lucky

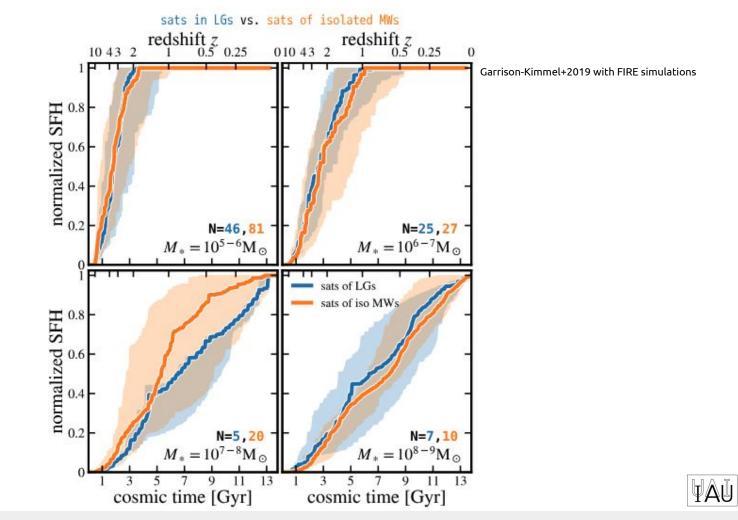




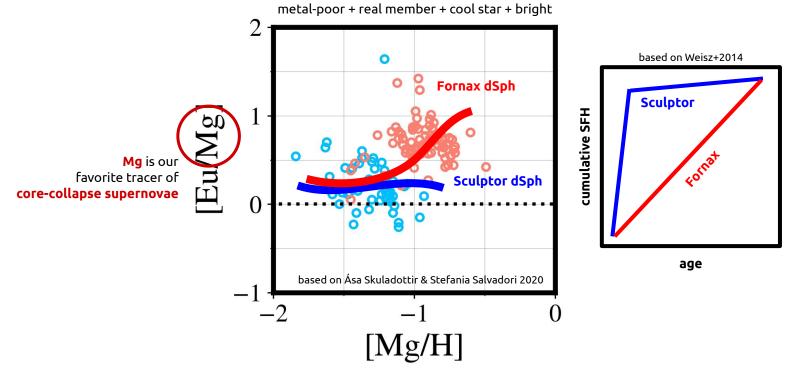
In **Reticulum II**, the r-process happened with a delay time of ≤500 Myr







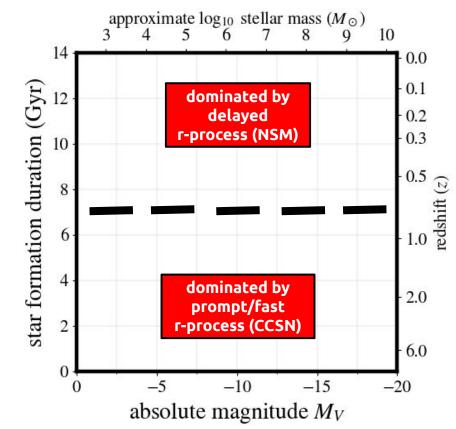
In "**massive**" **dwarf** spheroidals, **r-process** appears to be **delayed** by several Gyr



a proxy for "duration" of star formation

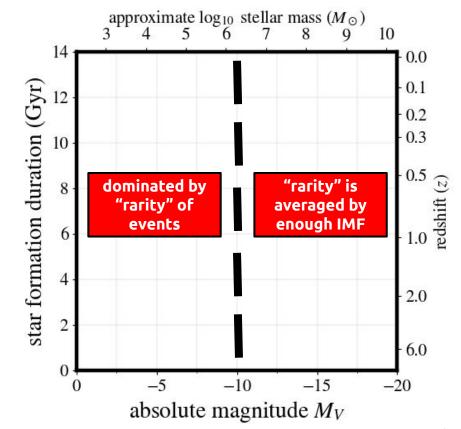


r-process enrichment in galaxies depends on star formation history and stellar mass



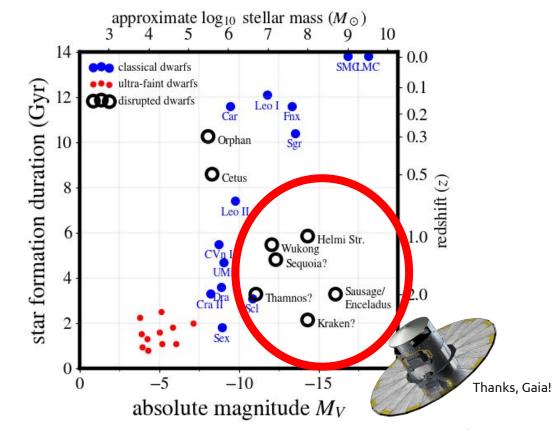
based on Frebel & Ji 2024 Handbook of Nuclear Physics

r-process enrichment in galaxies depends on star formation history and stellar mass

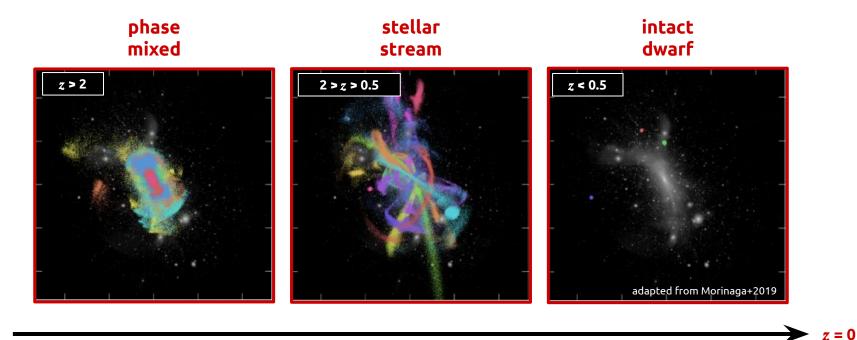


based on Frebel & Ji 2024 Handbook of Nuclear Physics

r-process enrichment in galaxies depends on star formation history and stellar mass



... but, with time, **dwarf Milky Way satellites** get more and more disrupted





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Gui Limberg – guilherme.limberg@usp.br

