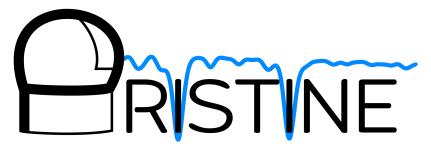


The Pristine survey

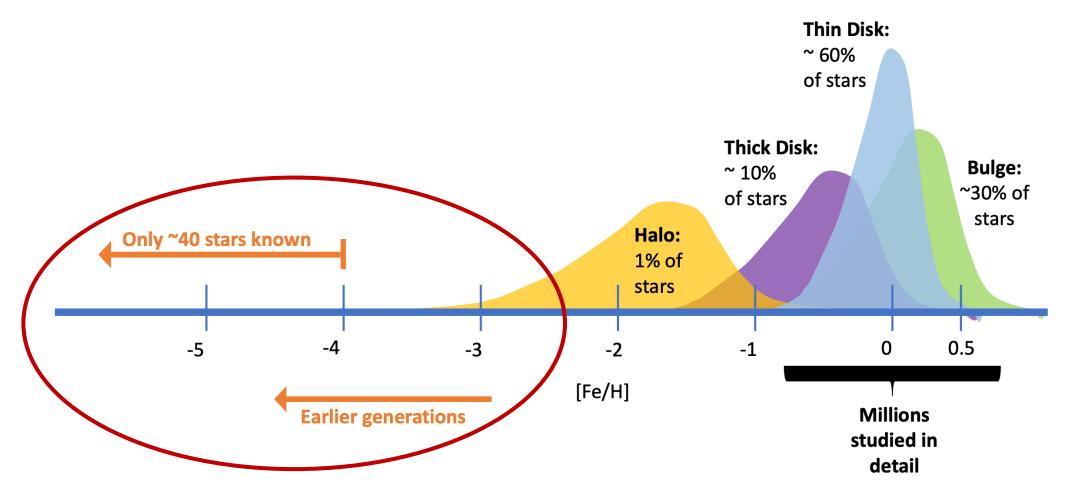
Else Starkenburg – University of Groningen



Artist impression of a highredshift galaxy star-forming. Credit: ESO

Galactic Archaeology to its limits

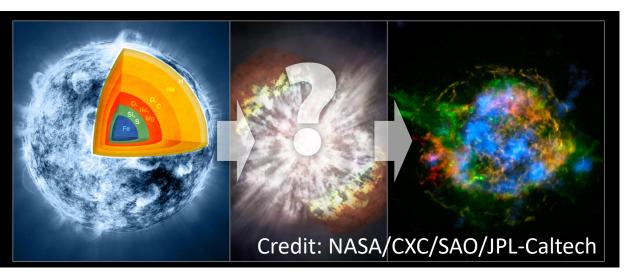
Milky Way history is very incomplete at very early times

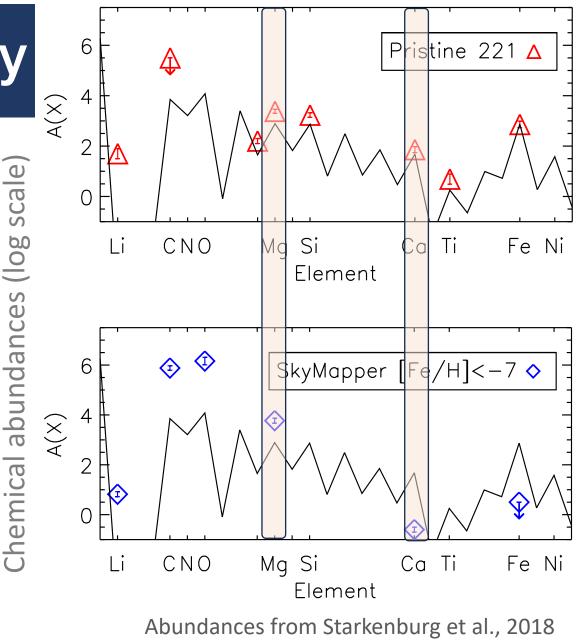


Galactic Archaeology

Early chemical evolution: Where abundances are excitingly different

 Different types of Supernovae





& Nordlander et al., 2017

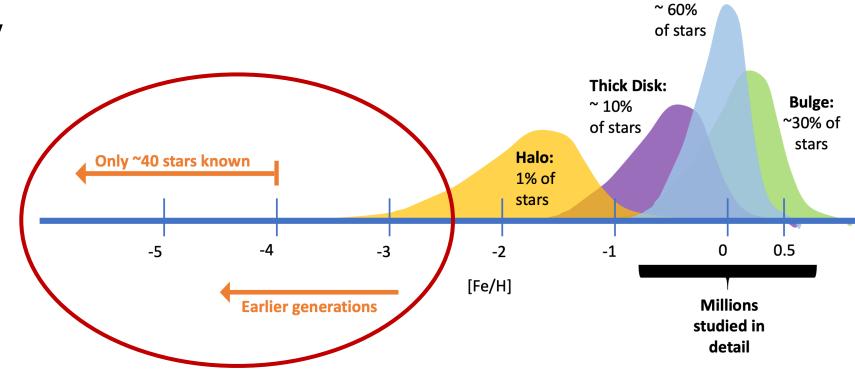
Galactic Archaeology to its limits

Milky Way history is very incomplete at very early times

Step 1:

Finding these rare stars

- Serendipitously in very large spectroscopic surveys
- Targeted, with narrowband, like:
 - The Pristine survey
 - Gaia XP spectra



Thin Disk:

Going through "the trash"

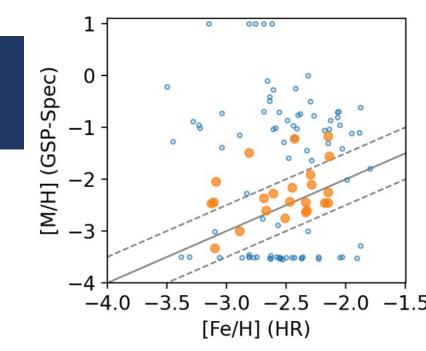
(of the spectroscopic surveys)

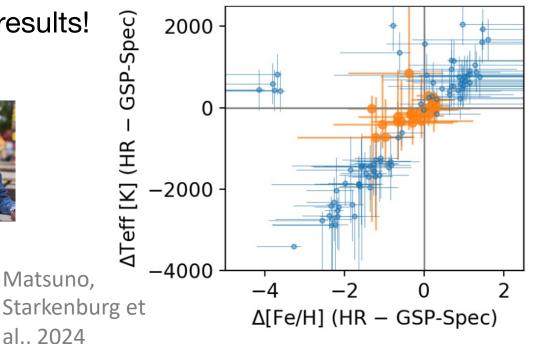
- Spectra of extremely metal-poor stars can be:
 - Mistakingly classified as hot stars
 - Mistakingly classified as noisy
 - "No lines to fit" no solution
- Many thanks to all the careful flags in the survey results!
 - These are typically bright, interesting targets
- LAMOST
 - Arentsen et al., 2023
- Gaia RVS
 - Matsuno et al., 2024; Viswanathan et al., 2024
- APOGEE
 - Montelius et al., in prep.



Matsuno,

al., 2024

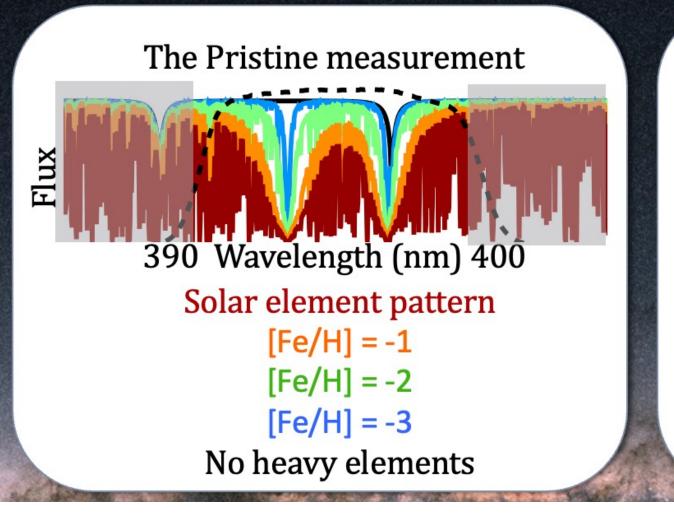




The Pristine survey



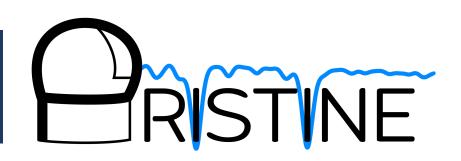
Starkenburg, Martin et al., 2017



Relative brightness in Pristine (compared to Gaia data)

brighter

Who is Pristine?



Pls: Else Starkenburg & Nicolas Martin

PhD students: Manuel Bayer, Spencer Bialek, Amanda Byström, Emma Dodd, Isaure Gonzalez Rivera, Jaclyn Jensen, Aroa del Mar Matas Pinto, Martin Montelius, Samuel Rusterucci, Akshara Viswanathan, Sara Vitali

Postdocs & staff:

David Aguado, Carlos Allende Prieto, Anke Ardern-Arentsen, Piercarlo Bonifacio, Elisabetta Caffau, Raymond Carlberg, Patrick Côté, Raphaël Errani, Sebastien Fabbro, Emma Fernández-Alvar, Morgan Fouesneau, Patrick Francois, Jonay González Hernández, Felipe Gran, Stephen Gwyn, Vanessa Hill, Rodrigo Ibata, Pascale Jablonka, Georges Kordopatis, Carmela Lardo, Linda Lombardo, Nicolas Longeard, Romain Lucchesi, Khyati Malhan, Lyudmila Mashonkina, Tadafumi Matsuno, Alan McConnachie, Camila Navarrete, Julio Navarro, Ruben Sanchez-Janssen, Mathias Schultheis, Federico Sestito, Salvatore Taibi, Guillaume Thomas, Eline Tolstoy, Kim Venn, Karina Voggel, Kris Youakim, Zhen Yuan

Pristine started in 2015 taking data with a CaHK filter on the Canada-France-Hawaii-Telescope

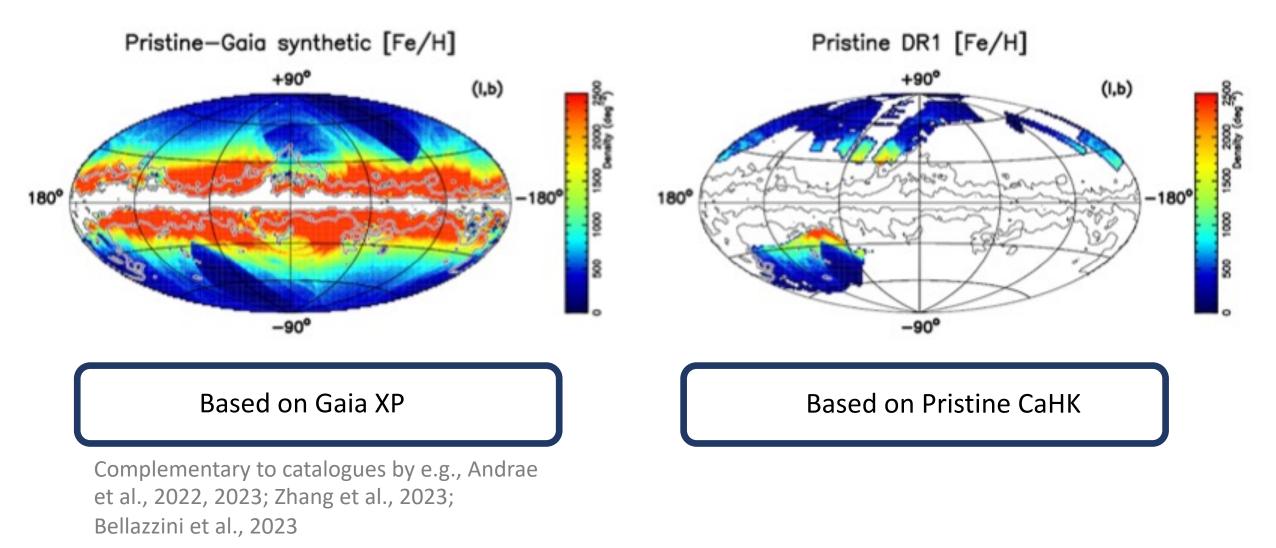
- to date 40+ papers

(Building on a long history of CaHK, see also talks of Ting Li, Wako Aoki, Anirudh Chiti)

Pristine Data Release 1



Martin, Starkenburg et al., 2024

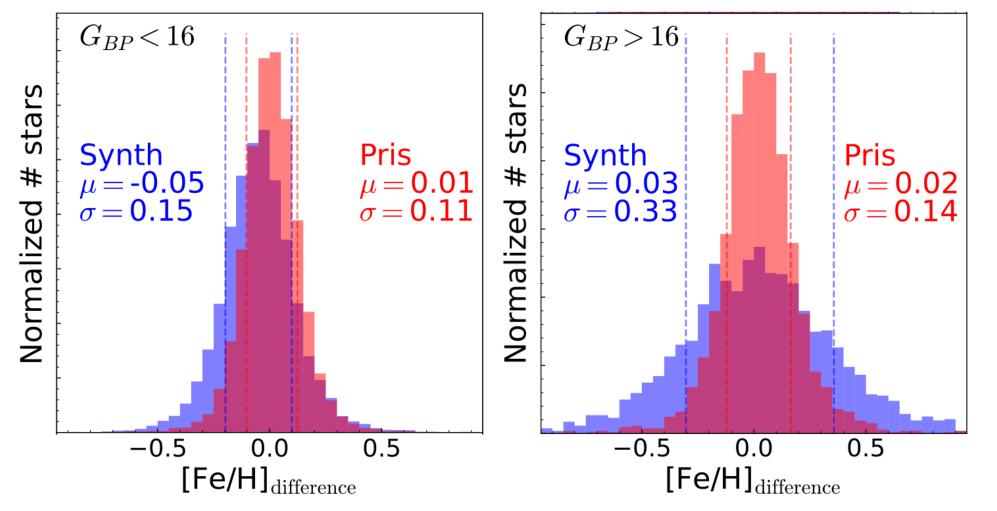


Pristine Data Release 1



Martin, Starkenburg et al., 2024

• Difference in quality at the fainter end ($G_{BP} > 16$)



Low-metallicity stars

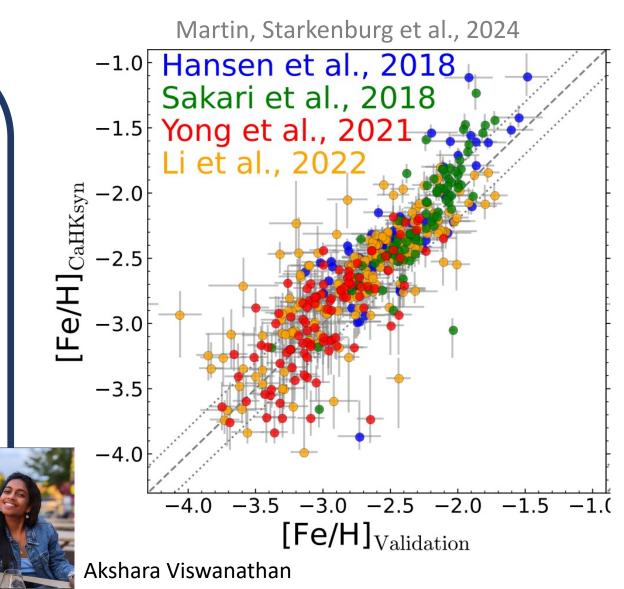


Works well until extremely lowmetallicity regime

• Dataset (+ Pristine data) released Martin, Starkenburg et al., 2024

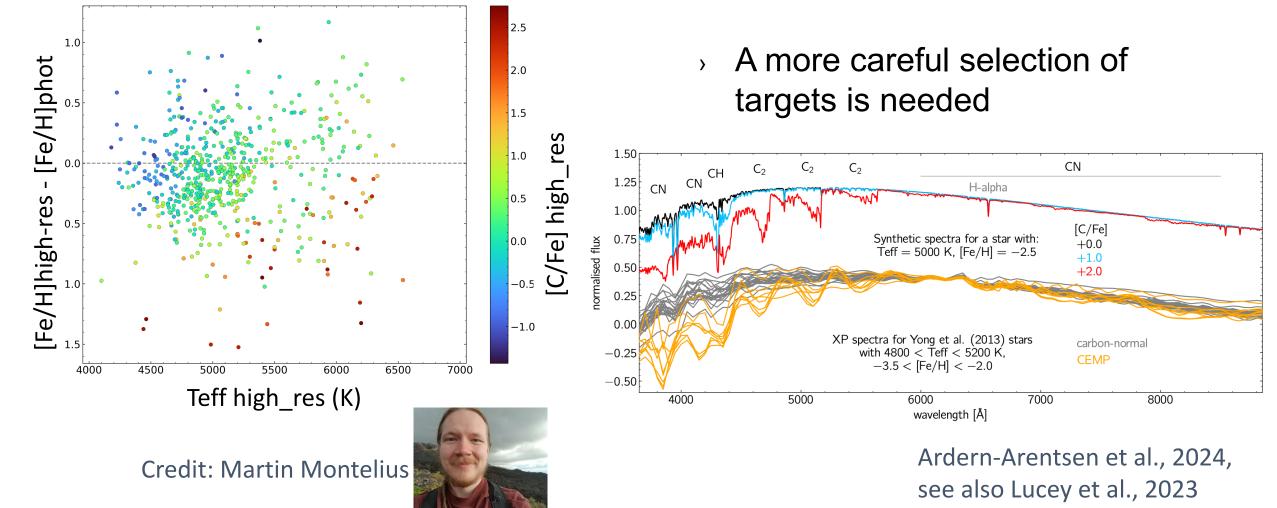
Our own follow-up

- Great success rates
 - 77% and 38% in finding stars with [Fe/H]<-2.5 and -3.0 (Viswanathan et al., 2024)



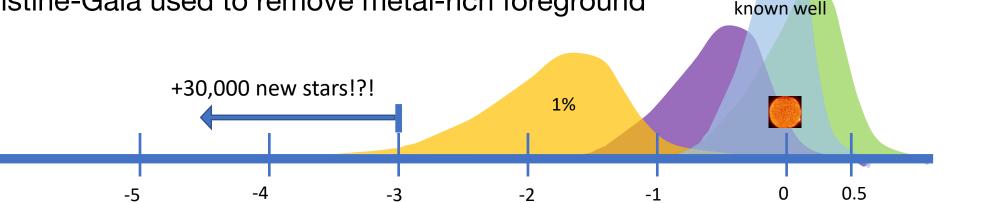
What are we missing?

We know we are missing the most extreme carbon-rich and cool stars



Big leap straight ahead

- Pristine & WEAVE agreement:
 - Spectroscopy for extremely metal-poor star candidates
 - Homogeneous study > 30,000 stars
 - Increase samples > 10-fold
- Pristine & 4MOST follow-up the inner Galaxy
 - Pristine-Gaia used to remove metal-rich foreground





millions

Metal-poor Targets

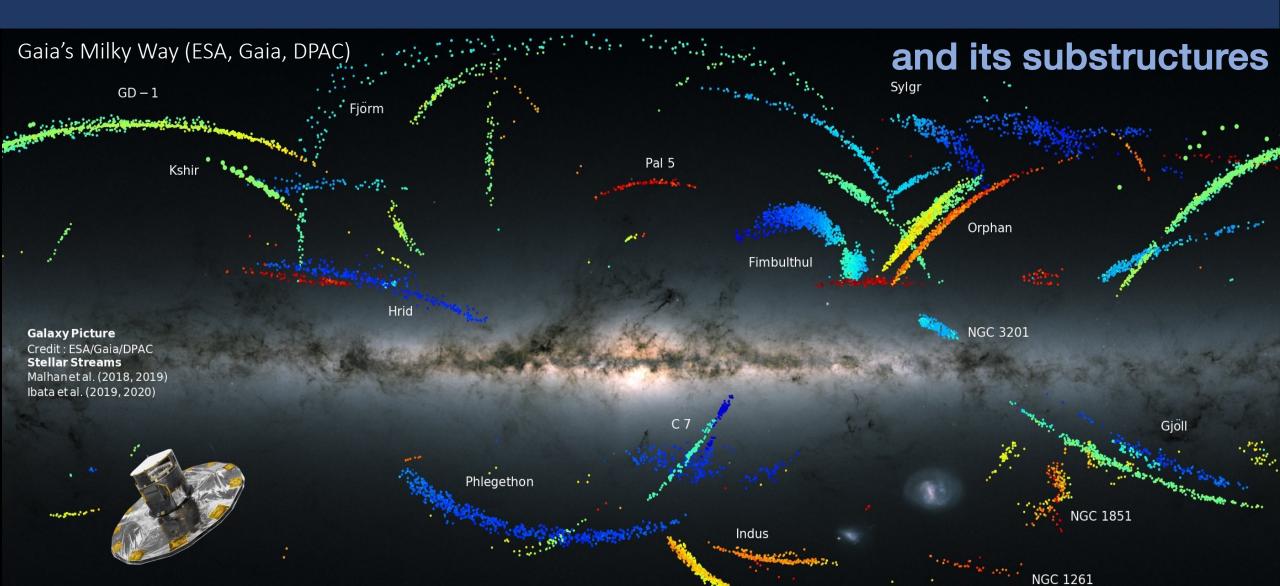
The chemical history of the early Galaxy

The very early generations of stars across Galactic components



Very metal-poor stars on disky orbits (with Pristine): e.g., Sestito et al., 2020; Fernández-Alvar et al. 2021; González Rivera de La Vernhe et al. 2024

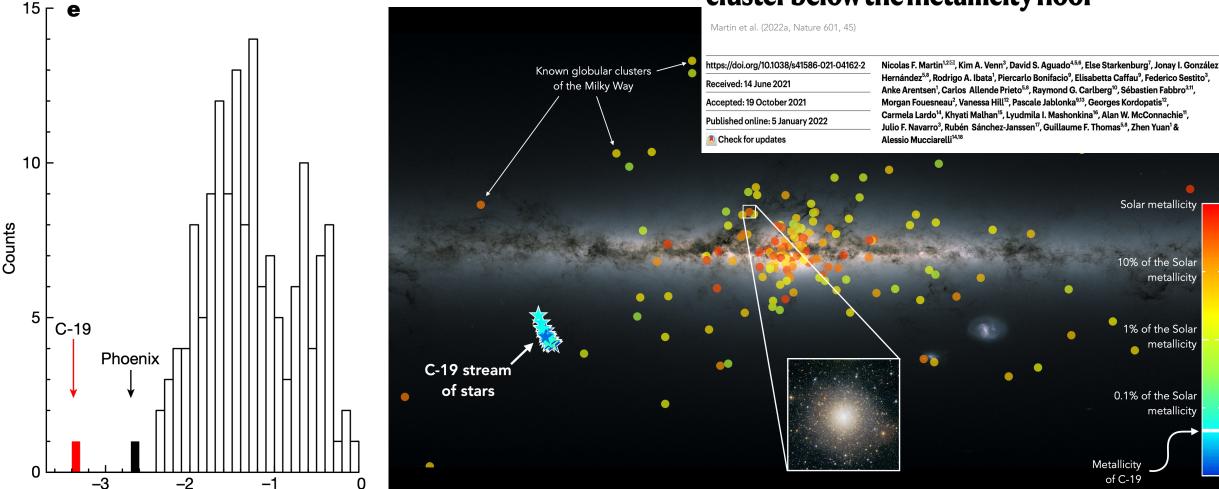
The chemical history of the early Galaxy



Will we find more of these?

[Fe/H]

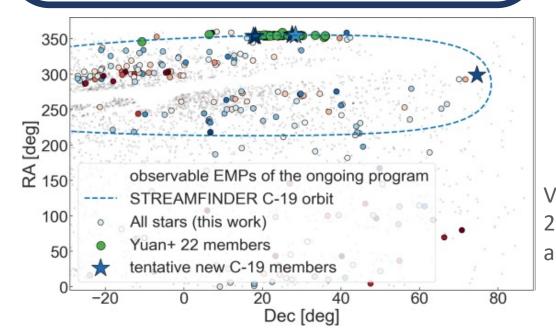
A stellar stream remnant of a globular cluster below the metallicity floor

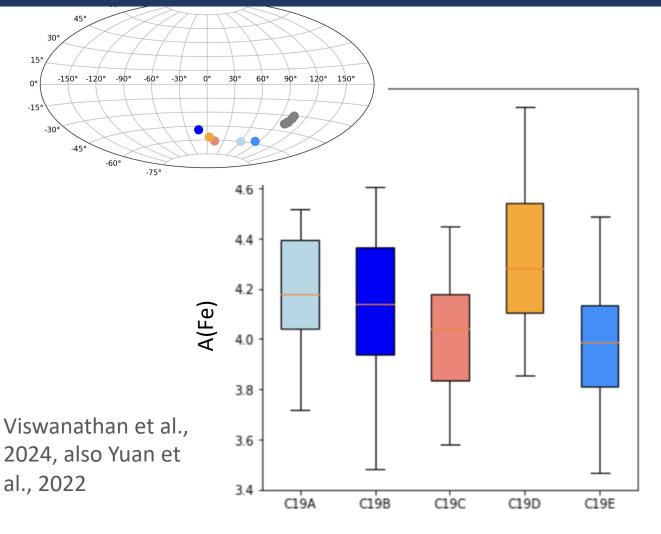


Martin, Venn, Aguado, Starkenburg et al., 2022

Will we find more of these?

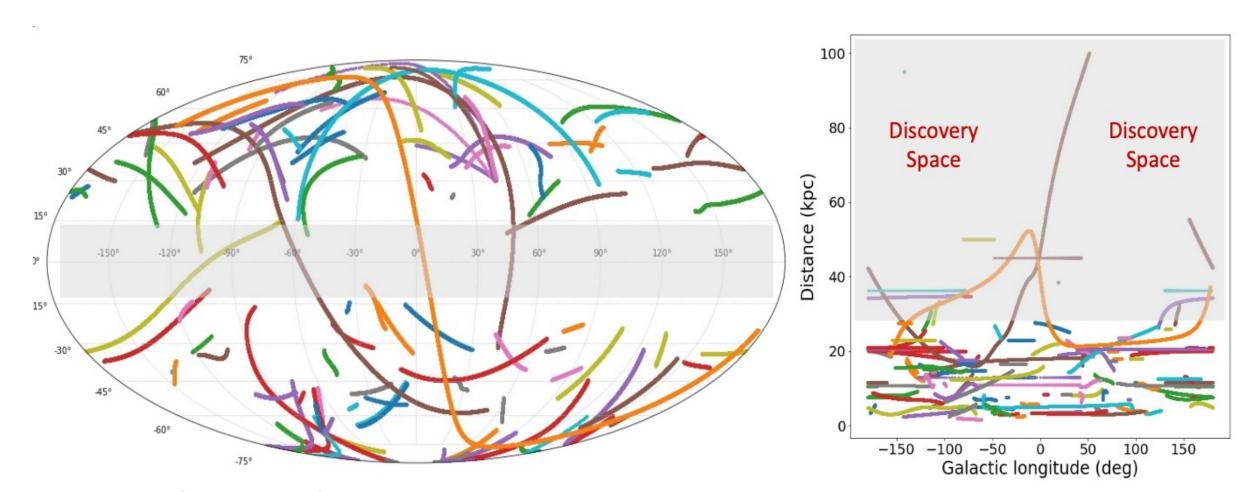
- C-19 is consistent with one metallicity
- We find more and more members further from the body of the stream





Venn et al., in prep.

How far out can we go?

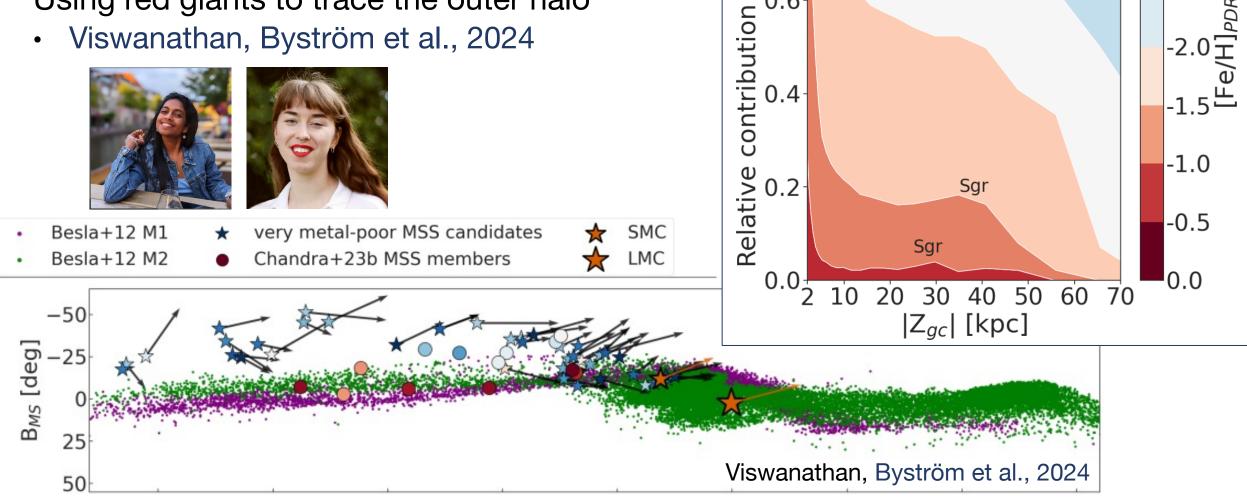


From galstreams package, Mateu, 2023

Outer Galaxy

Using red giants to trace the outer halo

Viswanathan, Byström et al., 2024



1.0

bins

of [Fe/H] |

0.6

HAC VOD 4.0

-3.5

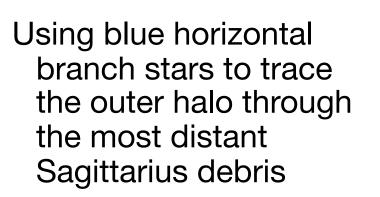
-3.0

-2.5

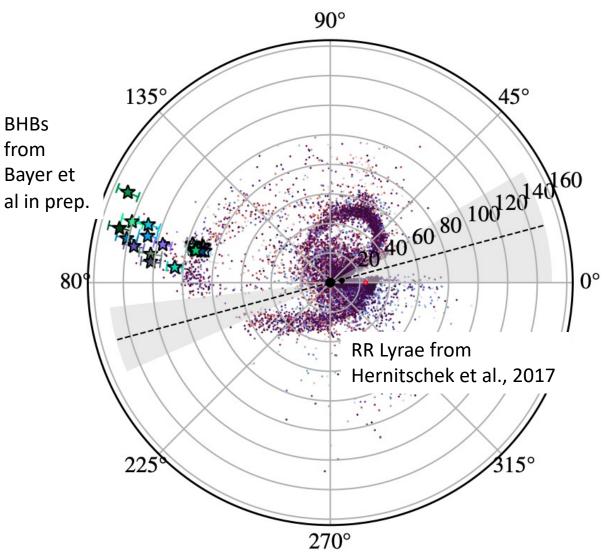
VMP stars d>50 kpc

outer-VOD

Additional serendipitous science



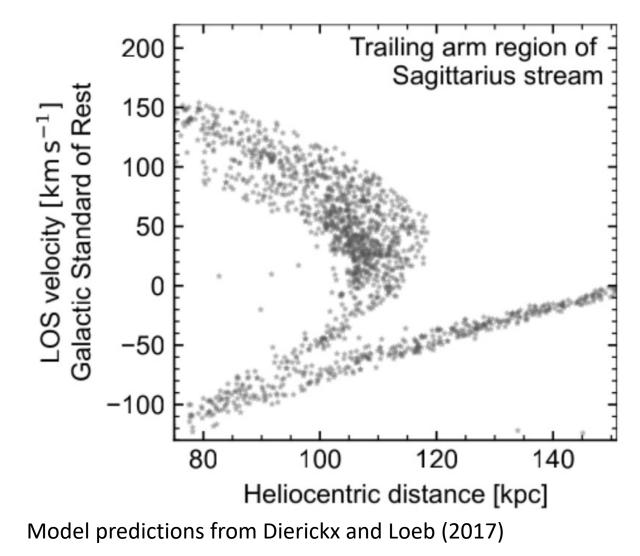
• Bayer, Starkenburg, Thomas et al., in prep.





Serendipitous standard candles

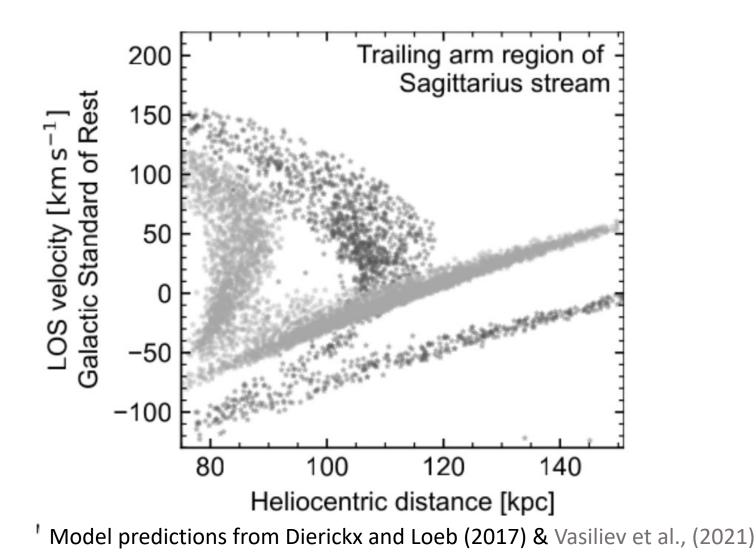
- Using blue horizontal branch stars to trace the outer halo through the most distant Sagittarius debris
- Bayer, Starkenburg, Thomas et al., in prep.





Serendipitous standard candles

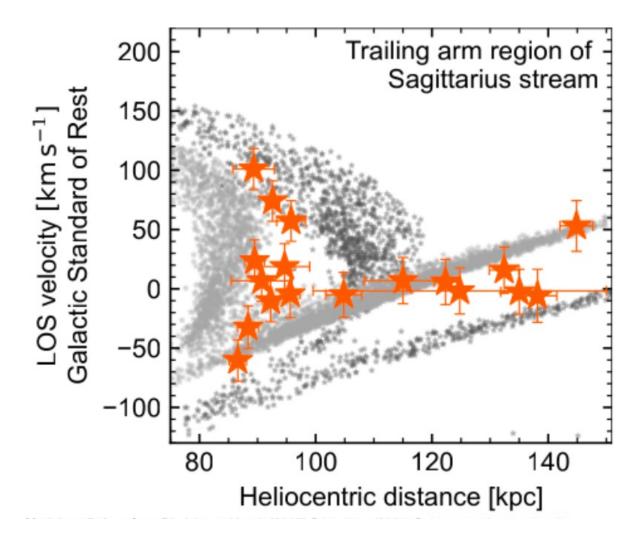
- Using blue horizontal branch stars to trace the outer halo through the most distant Sagittarius debris
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Serendipitous standard candles

- Using blue horizontal branch stars to trace the outer halo through the most distant Sagittarius debris
- Bayer, Starkenburg, Thomas et al., in prep.





Getting ready for the future... to better study the past



Finding these rare stars

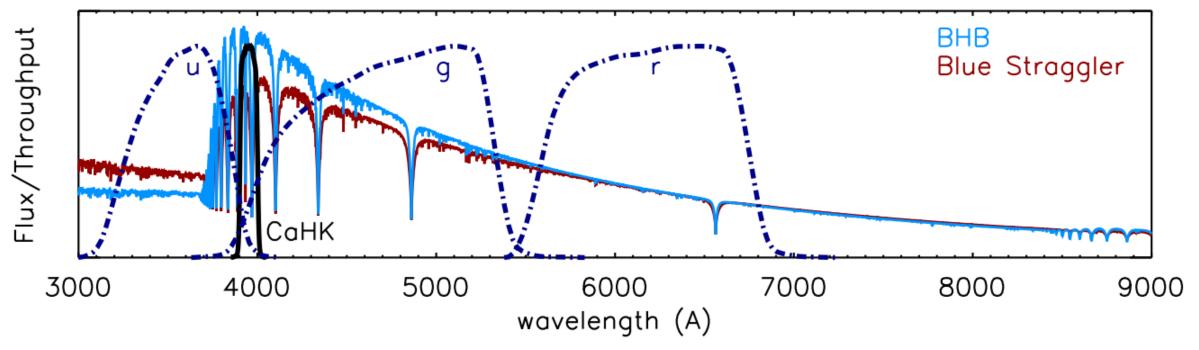
- Candidate selection
- Follow-up with WEAVE and 4MOST

Interesting physics at the metal-poor (& far) end

- Special supernovae
- Early Milky Way build-up to the outer halo
- The earliest Globular clusters?

Serendipitous science: Blue Horizontal Branch stars

- Select A-stars, but disentangle AGN, blue stragglers, main-sequence ...
 - .. and genuine blue horizontal branch stars



Starkenburg et al., 2019