

JWST Unveils IR Stellar Content of NGC 6822 with NIRCam and MIRI

Conor Nally, Olivia Jones + Annette Ferguson + WINGS Team



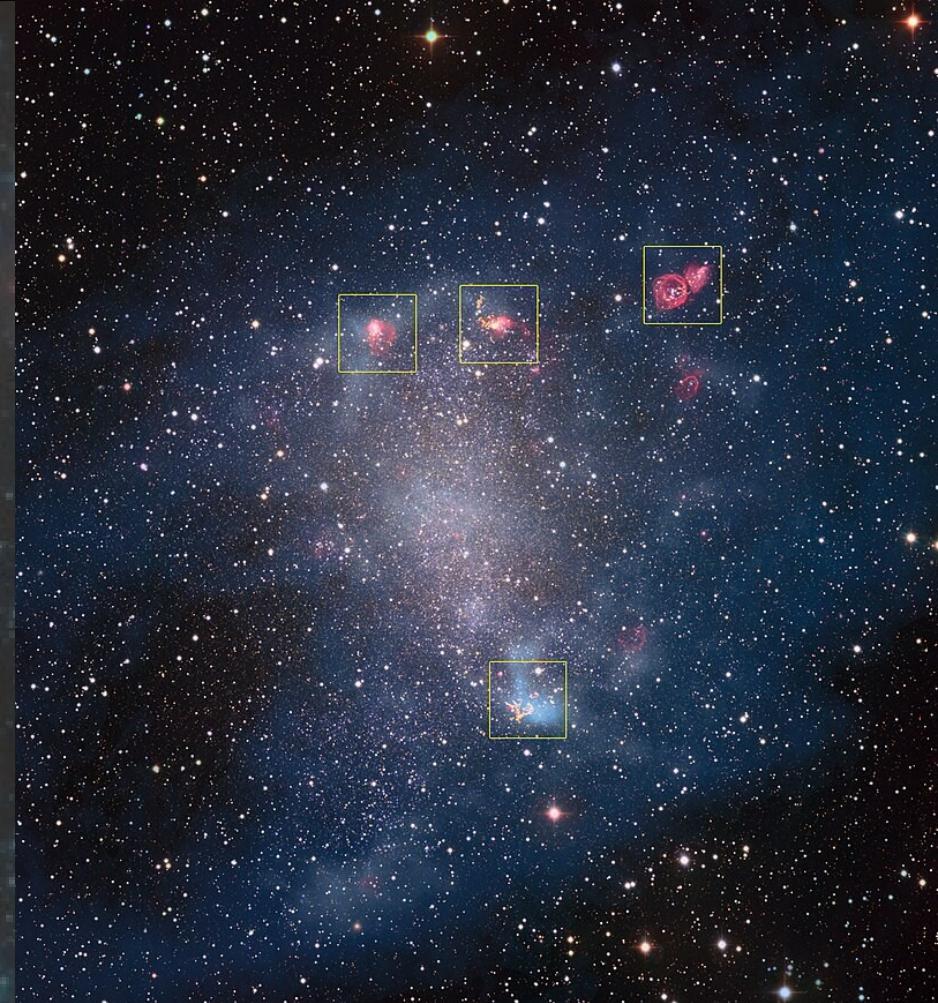
University
of Edinburgh

IAU S395
Paraty 2024



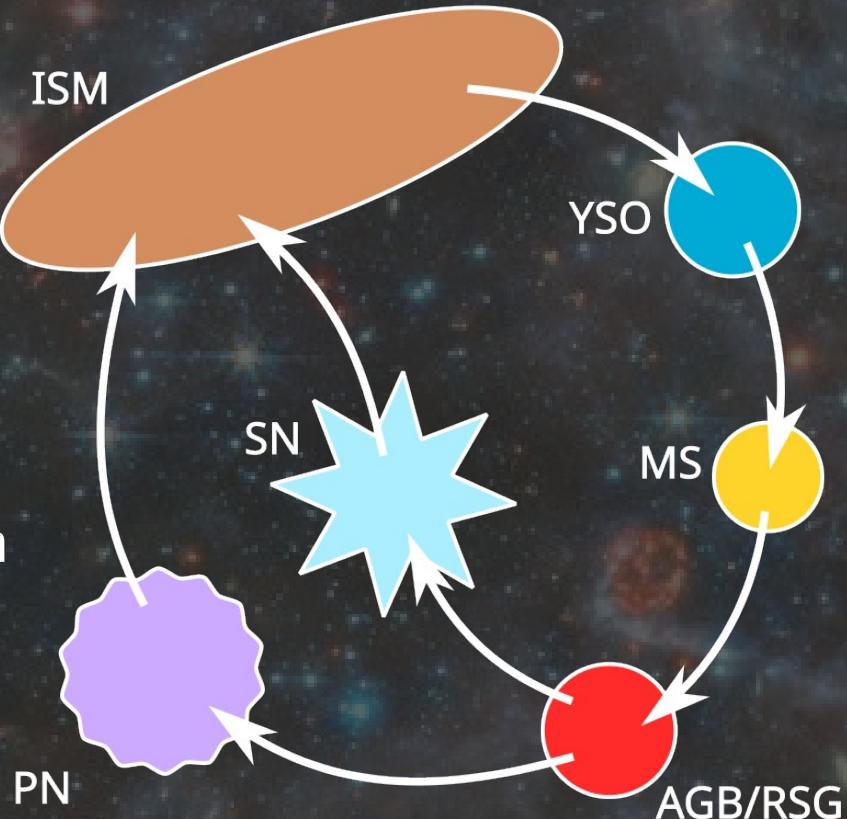
NGC 6822

- ★ Irregular dwarf galaxy
- ★ Nearby (~500kpc)
- ★ Metal poor (~ $0.2 Z_{\odot}$)
- ★ Tidally isolated (disputed)
- ★ Home to several massive star forming regions (Hubble + Spitzer regions)
- ★ Rich population diversity offer a good opportunity to study dust in early universe



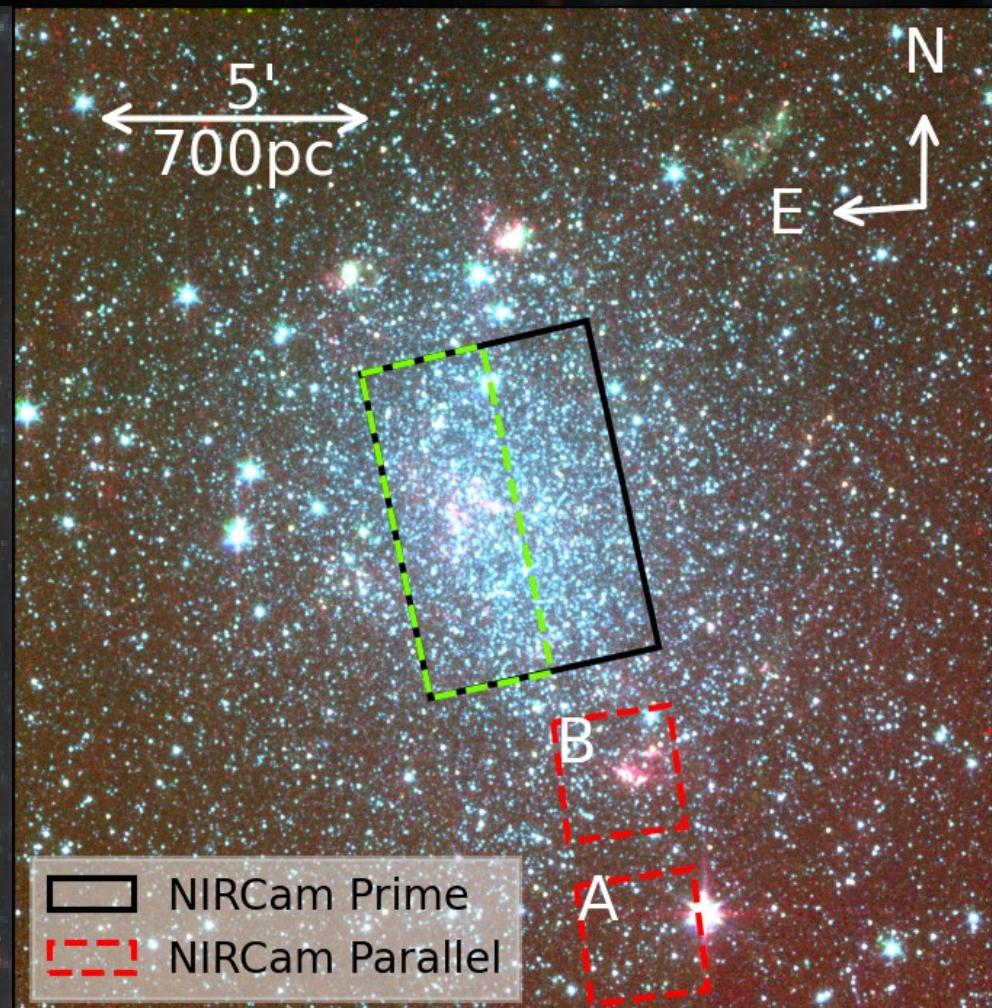
Dust Life Cycle

- ★ Interstellar medium is the dust bank
- ★ Dust locked up in newly formed stars
- ★ Matter enriched during stellar evolution
- ★ AGBs inject dust back into the ISM
- ★ SN inject dust but also vaporise some of it again
- ★ **The key stages of this cycle are bright in the infrared**



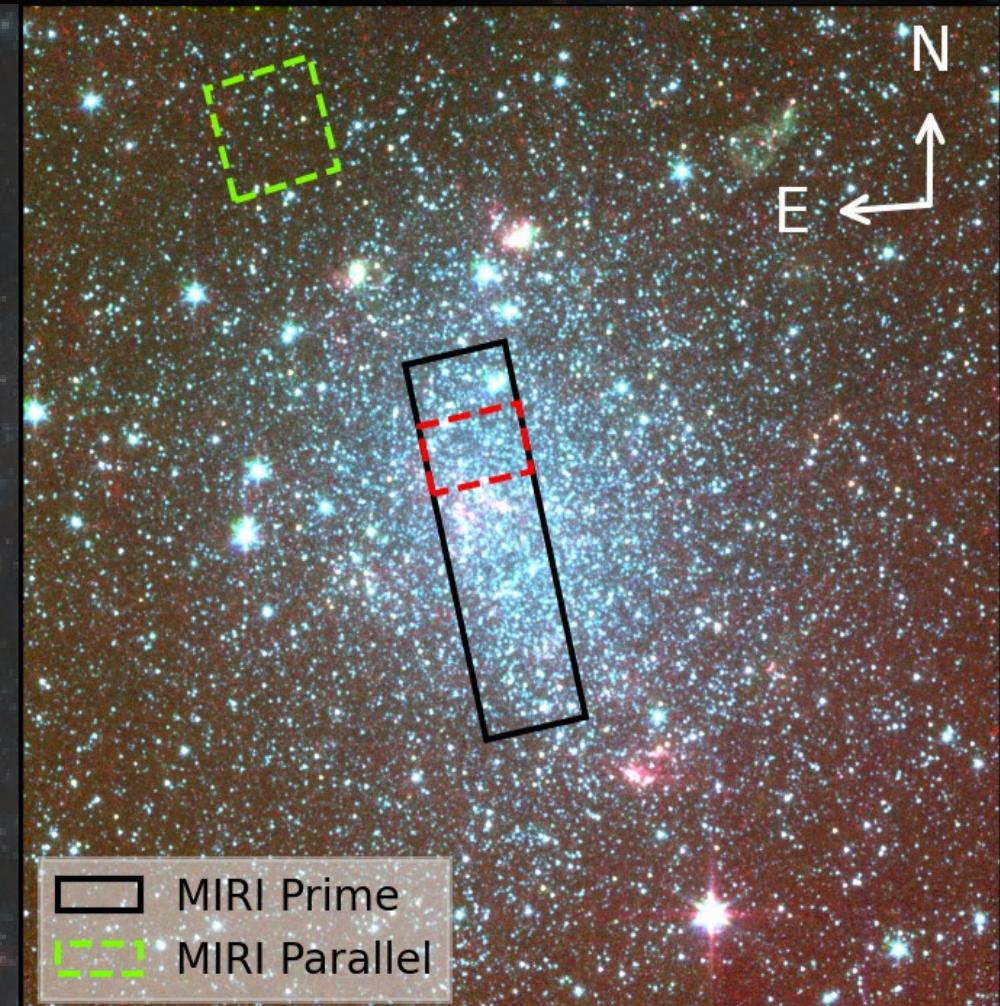
JWST Imaging

- ★ 20 hour JWST imaging program
- ★ 4 NIRCam filters (4 pointings)
F115W, F200W, F356W, F444W
- ★ 4 MIRI filters (6 pointings)
F770W, F1000W, F1500W, F2100W
- ★ 2 NIRCam Parallel field pointings
- ★ 1 MIRI Parallel field pointings



JWST Imaging

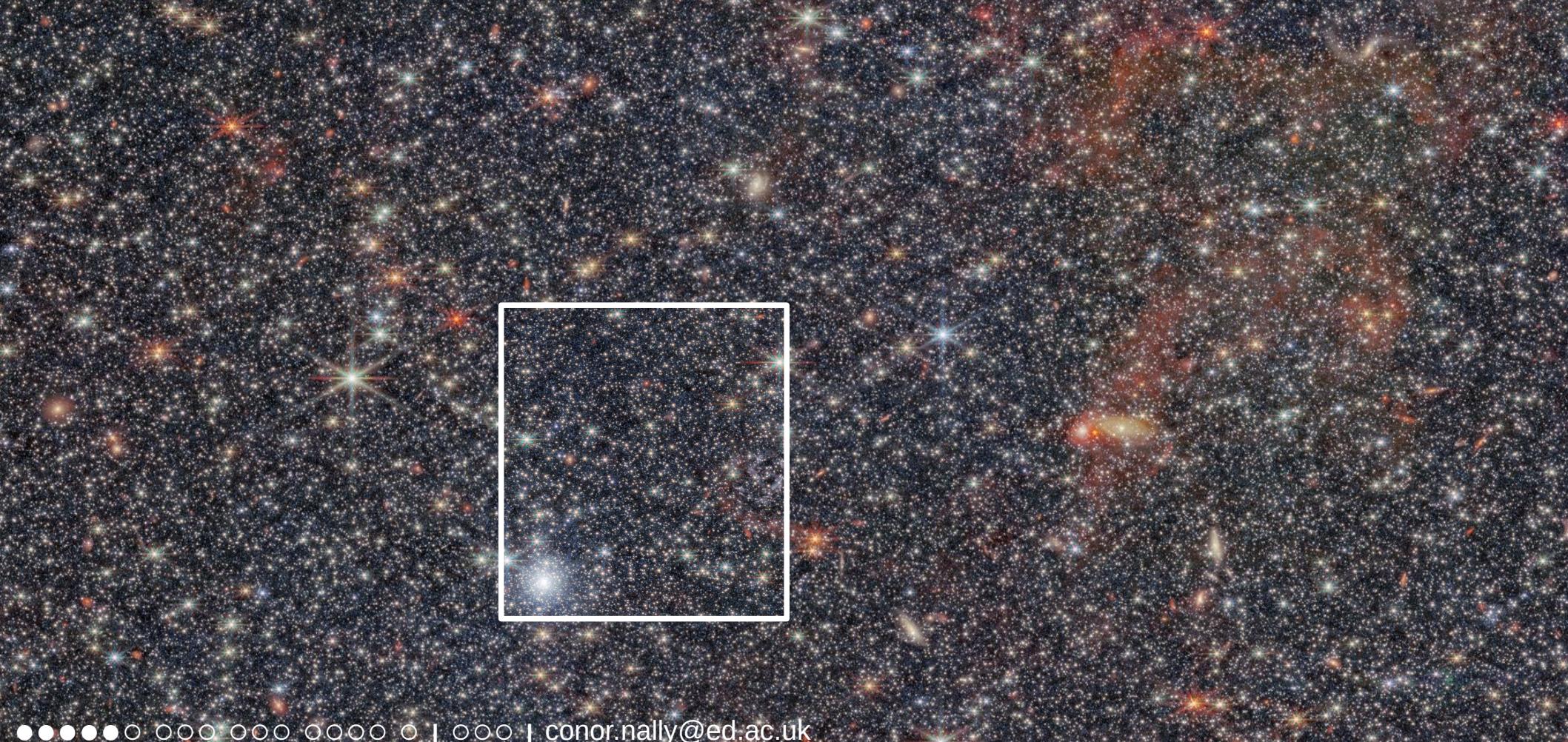
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JWST Imaging - NIRCam



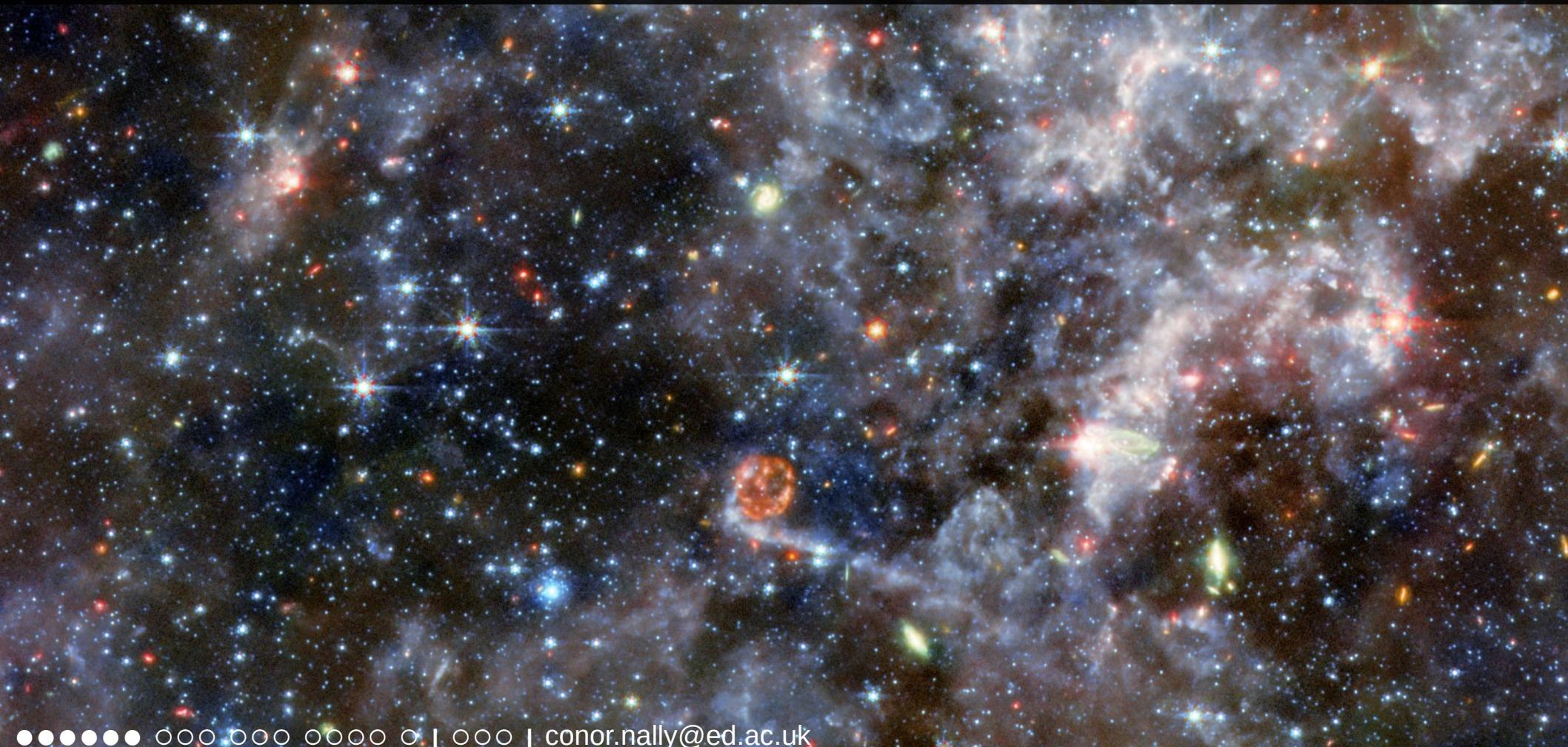
JWST Imaging - NIRCam



JWST Imaging - NIRCam



JWST Imaging - MIRI



Photometry - StarbugII

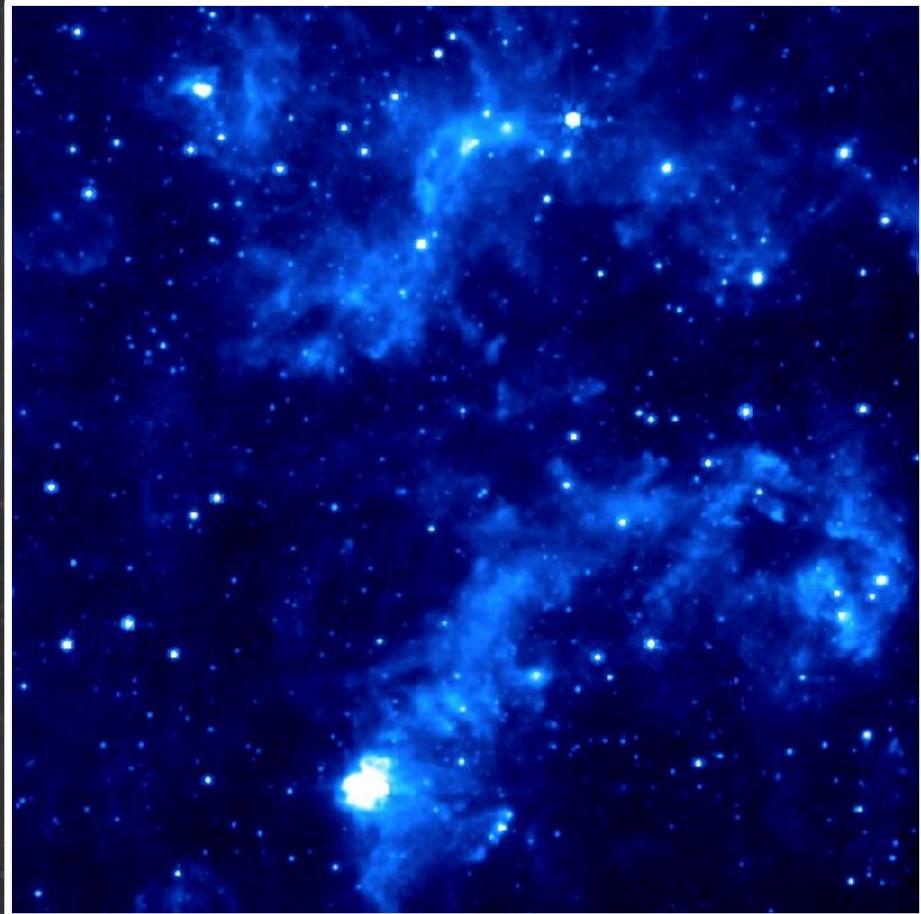
- ★ Source crowding
- ★ Dusty diffuse emission contamination
- ★ Large range in PSF FWHM ($0.023''$ – $0.8''$)
- ★ Background source contamination
- ★ Large catalogues
- ★ Open source era



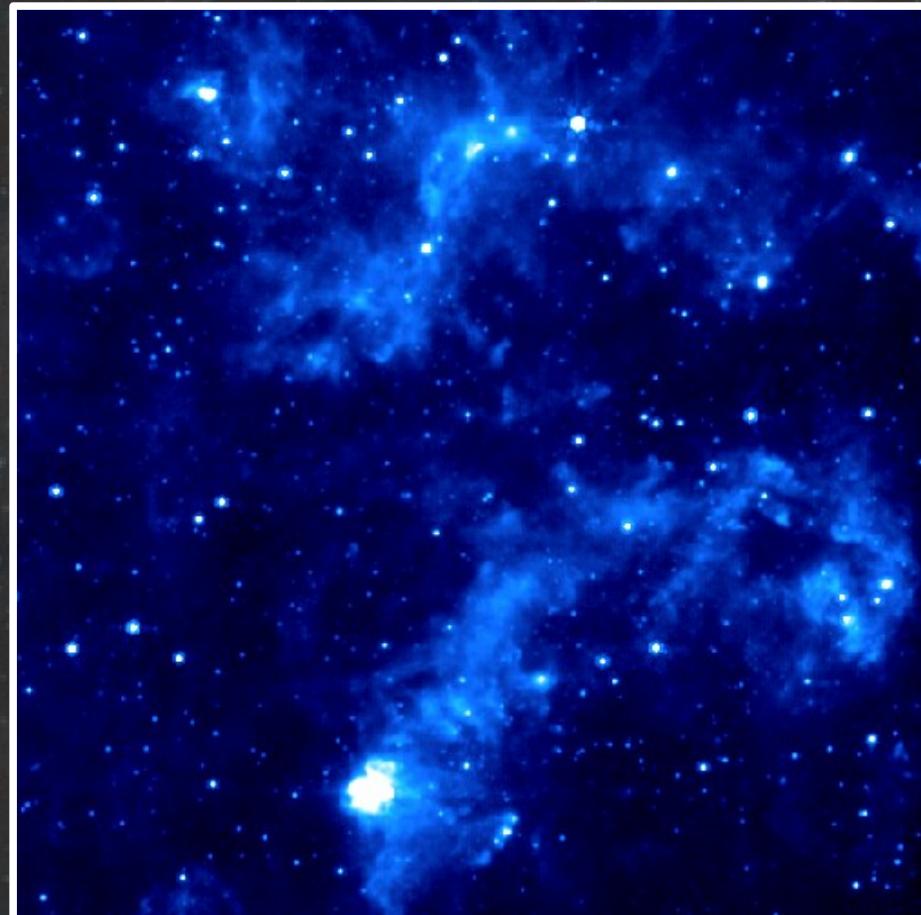
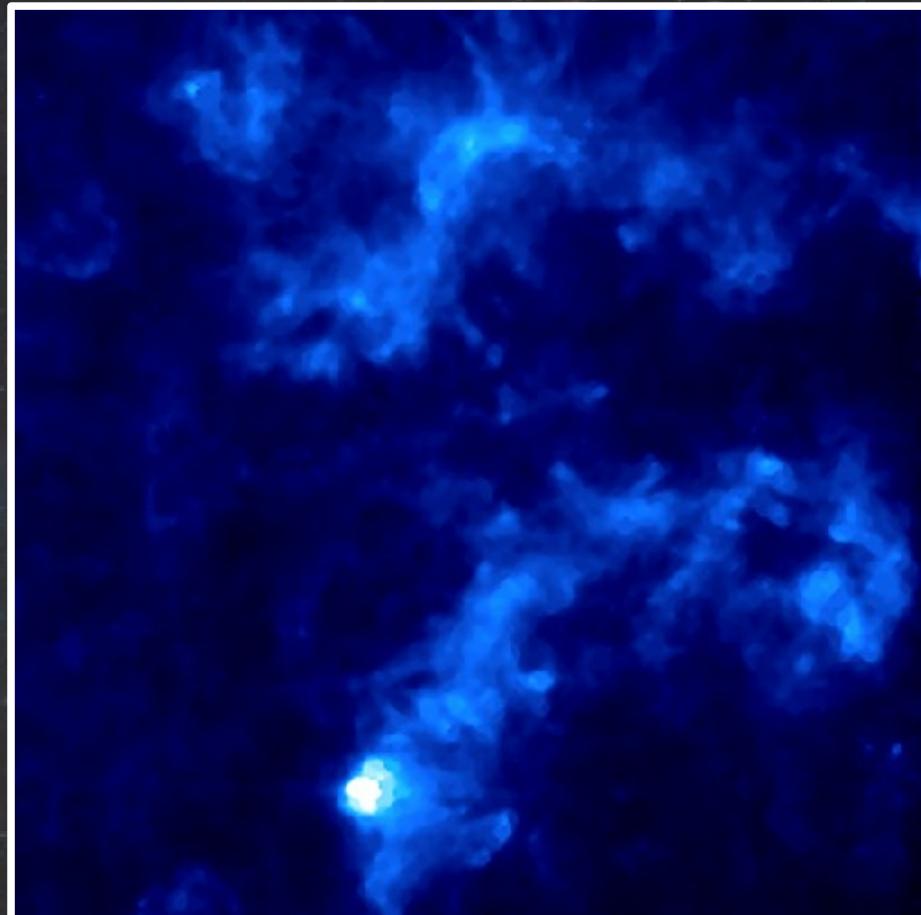
STARBUGII [ascl:2309.012]
\$~ pip install starbug2

Photometry - Source Detection

- ★ Three fold ensemble of detection methods
- ★ Targetting faint, compact and dust embedded stars
- ★ An order of magnitude greater detection rate than JWST pipeline
- ★ Detection of first extra galactic brown dwarfs (Zeidler et al. 2024)
- ★ Array of tunable parameters to reduce contaminants

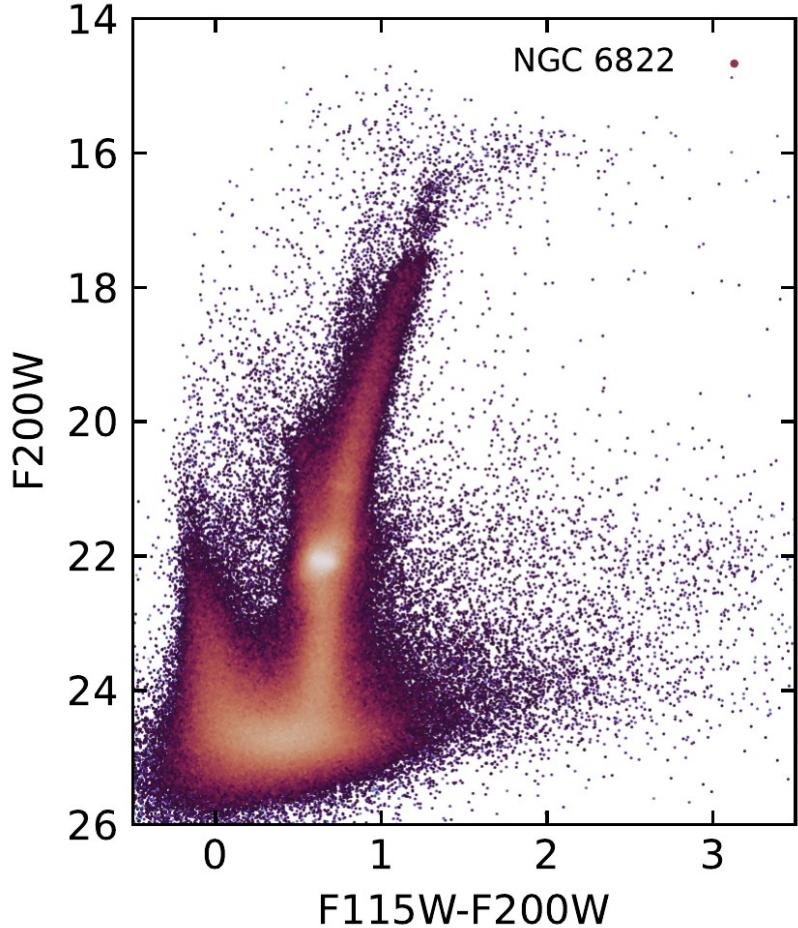


Photometry - Diffuse Emission



NIRCam Catalogue

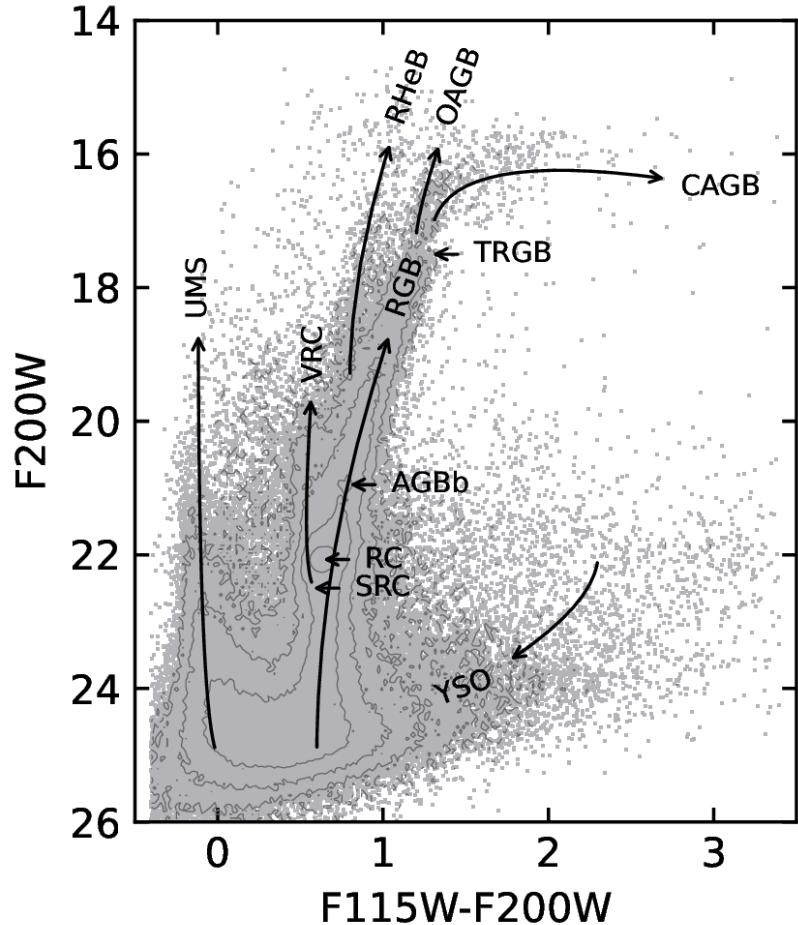
- ★ ~ 1,000,000 detected sources
- ★ ~ 3 magnitudes below the Red Clump
- ★ ~ 7 magnitudes deeper than comparable IR studies before
- ★ Wide array of spectral types found
- ★ Emergence of the elusive AGB bump



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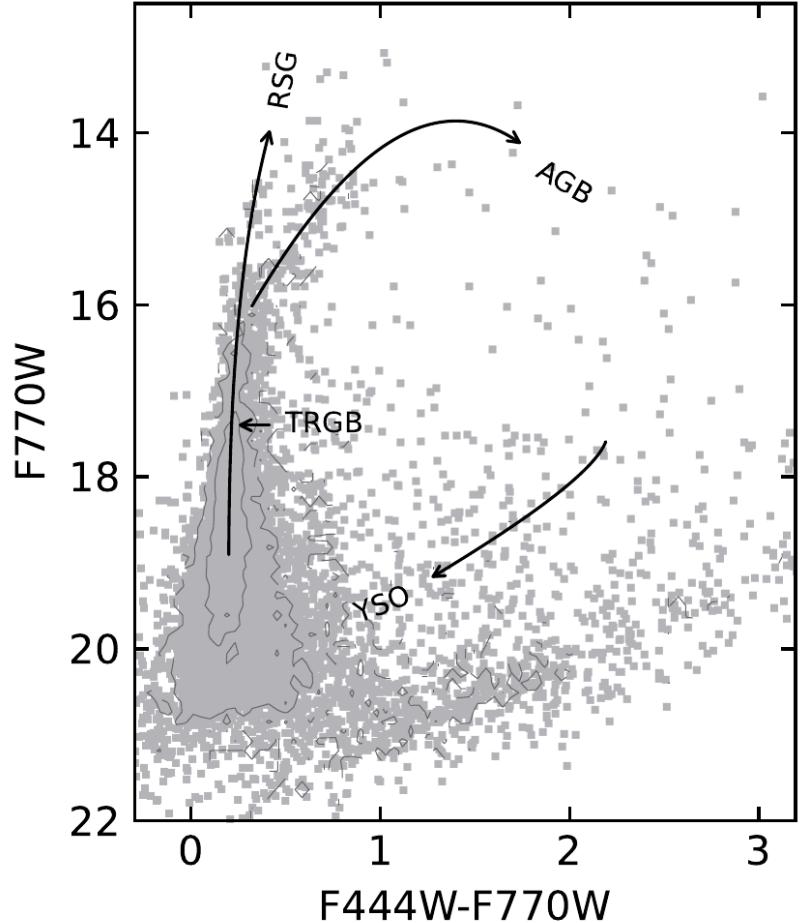
(Nally et al. 2024)



MIRI Catalogue

- ★ ~ 22,000 detected sources
- ★ We match the catalogues across the full wavelength range (1 – 21 microns)
- ★ We identify ~2000 AGB candidates
- ★ Detection of 90 deeply embedded YSOs (Lenkić et al. 2024)

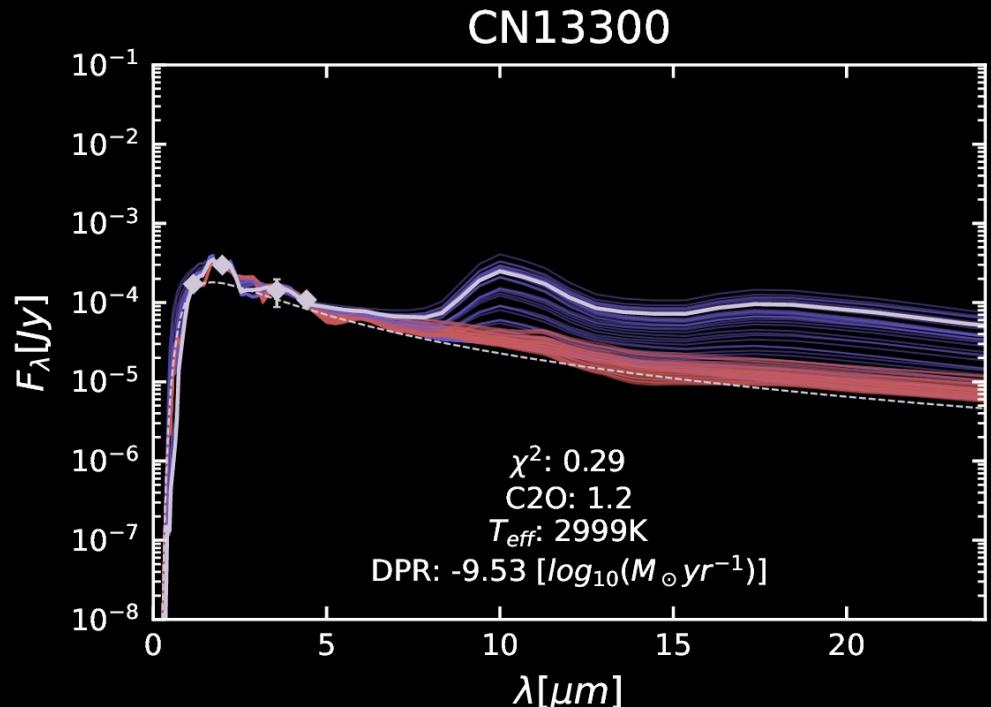
(Nally et al. 2024)



SED Fitting AGB Models

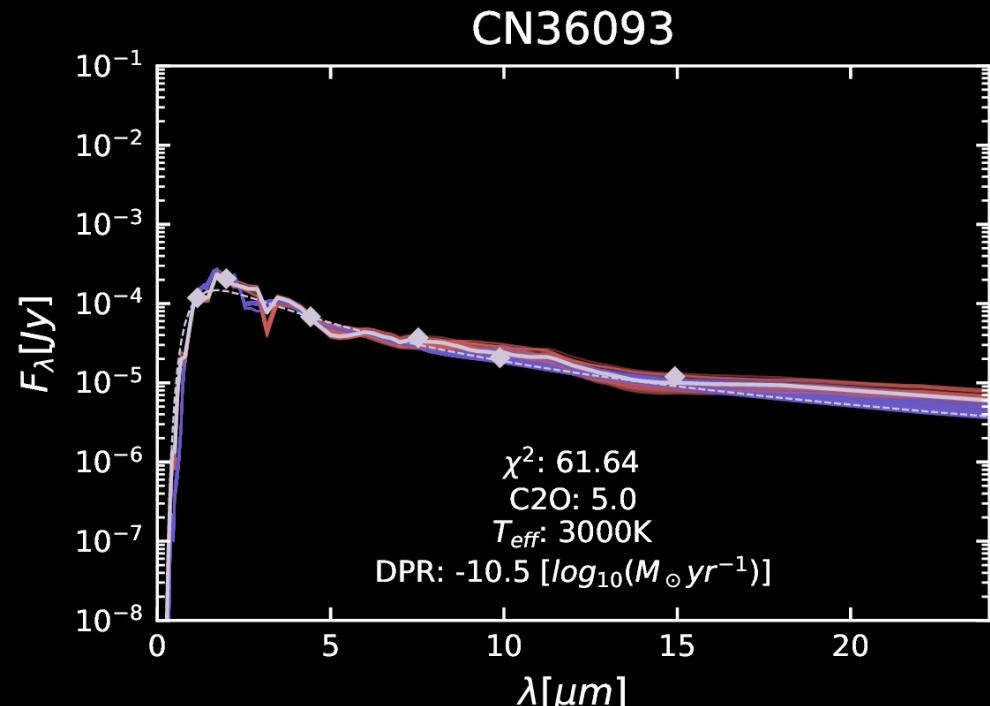
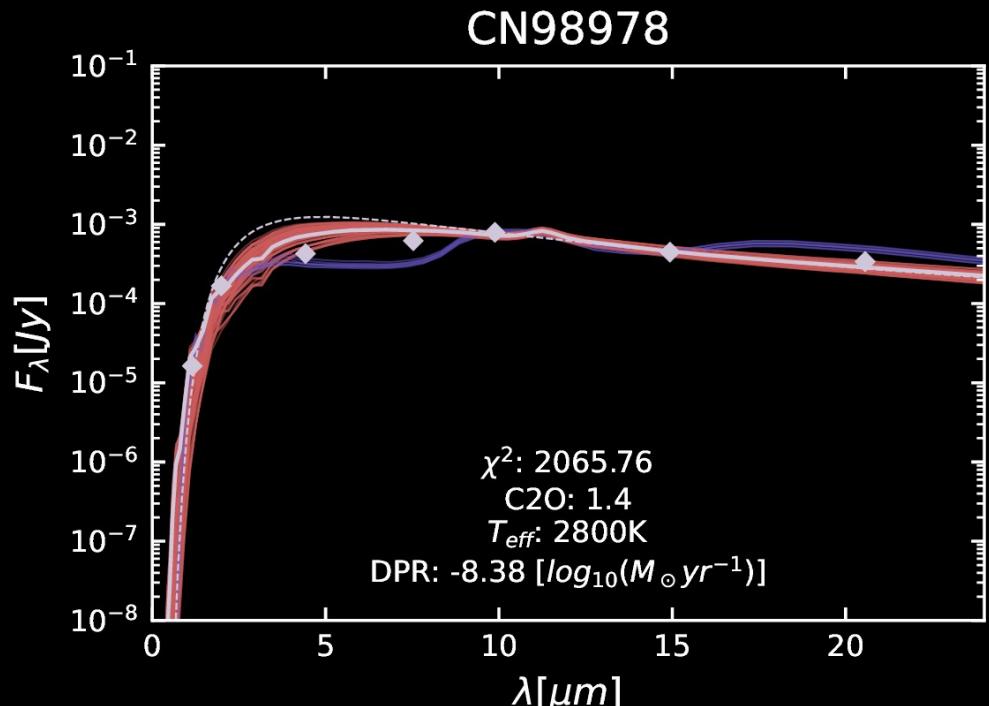
Preliminary

- ★ We fit GRAMS carbon-rich AGB (Srinivasan+2011) and oxygen-rich AGBs (Sargent+2011) to AGB SEDs
- ★ Determining:
 - Chemical class dominance (C or O)
 - Dust production rate (DPR)
 - Bolometric luminosity
 - Temperature
 - Dust shell properties
- ★ Long wavelength baseline allows us to constrain dust chemistry



SED Fitting AGB Models

Preliminary

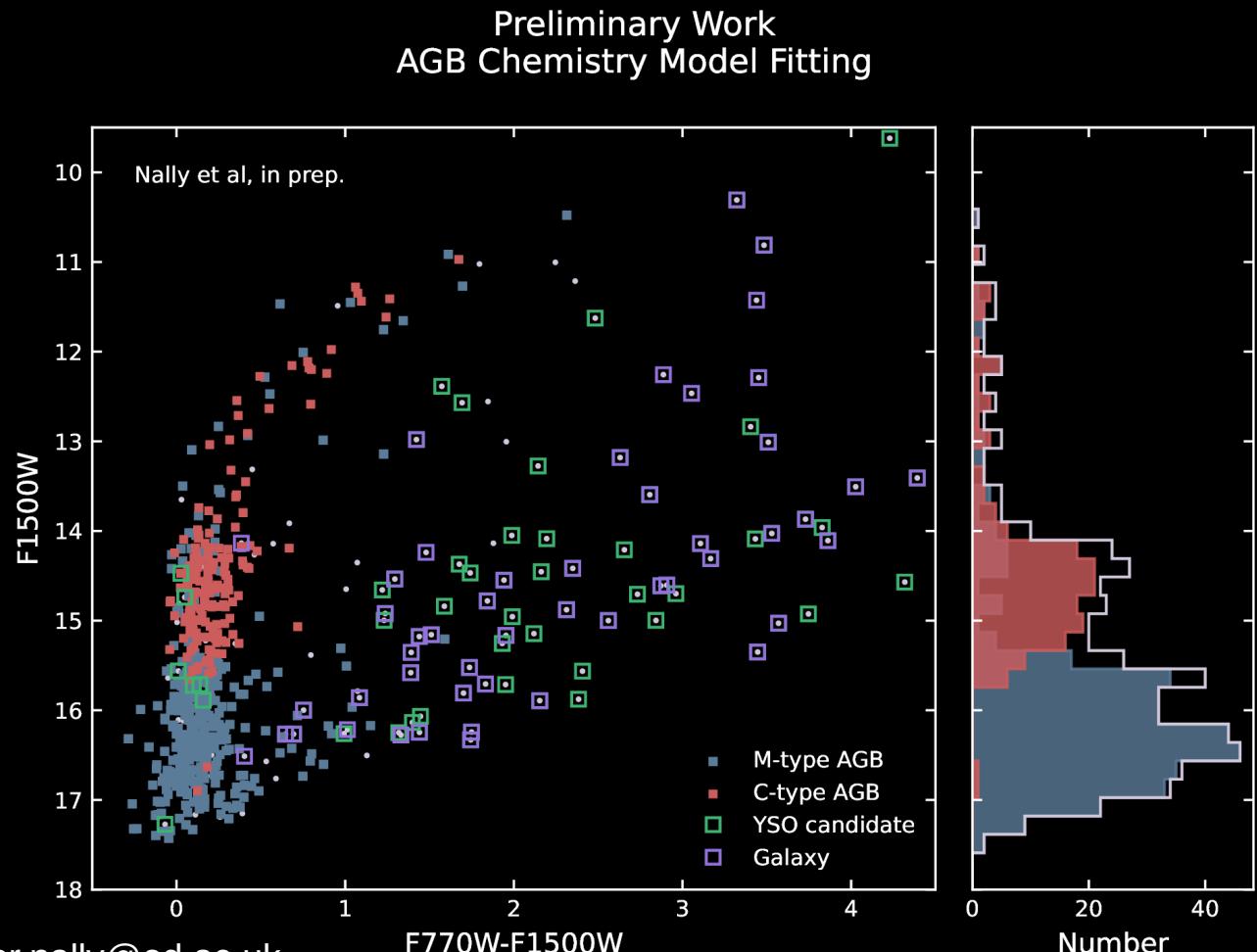


- ★ Determining: Chemical class, Teff, DPR, Lbol and dust shell properties

AGB Chemical Separation

Preliminary

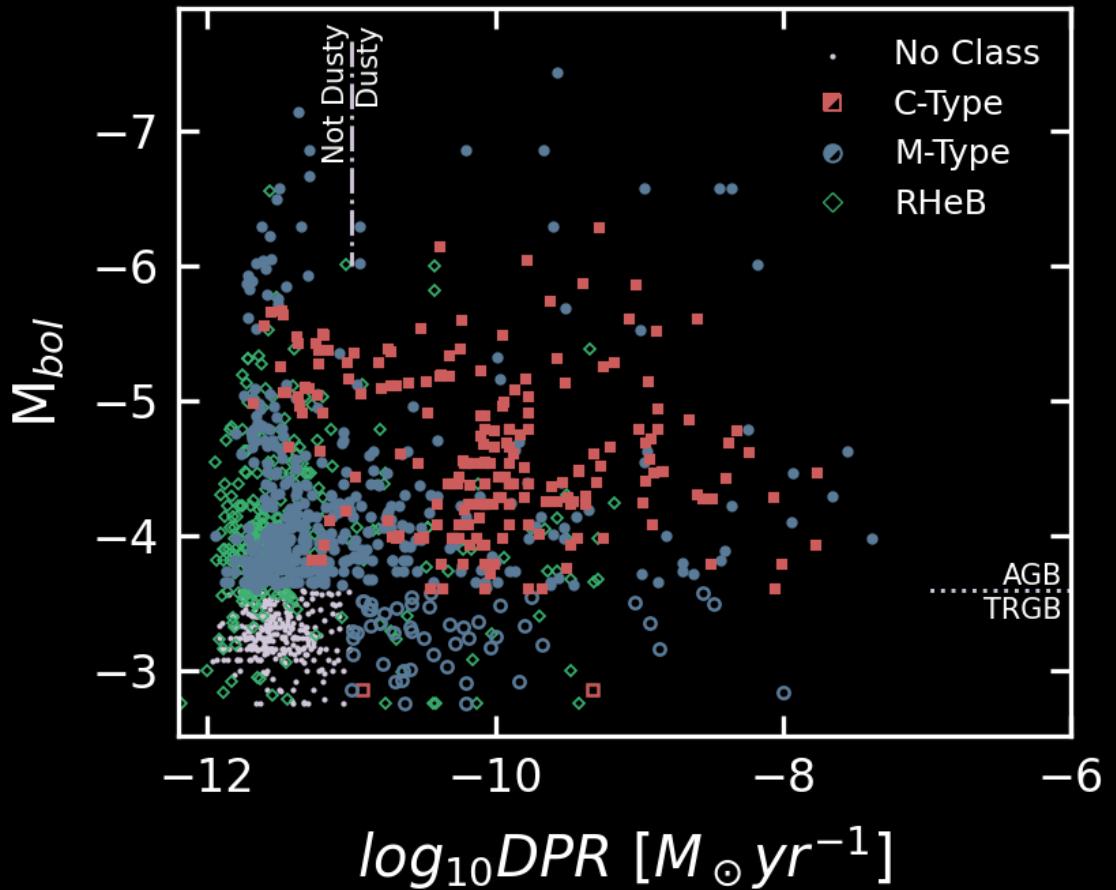
- ★ AGB chemical dominant species
 - Oxygen M-type OAGBs
 - Carbon C-type CAGBs
- ★ We can separate the AGB chemical types (almost)
- ★ We can remove galaxy and YSO contaminants



AGB Dust Production

Preliminary

- ★ Majority of the DPR comes from a few individuals
- ★ Most numerous dusty stars are C-rich AGBs
- ★ The highest mass losing individuals are O-rich AGB
- ★ We find a possible dust producing hot RHeB



Summary



- ★ STARBUGII is an open-source photometry suite designed for complex crowded source detection and PSF photometry
- ★ We present a near- to mid-IR matched catalogue of NGC 6822, recovering a wide array of stellar populations. We have only just scratched the surface of this data set
- ★ AGB chemical SED modelling shows great CMD space separation
- ★ Carbon dominated AGB dust production but high mass loss OAGB individuals

JWST WINGS

JAMES WEBB SPACE TELESCOPE
SMALL MAGELLANIC CLOUD | NGC 346

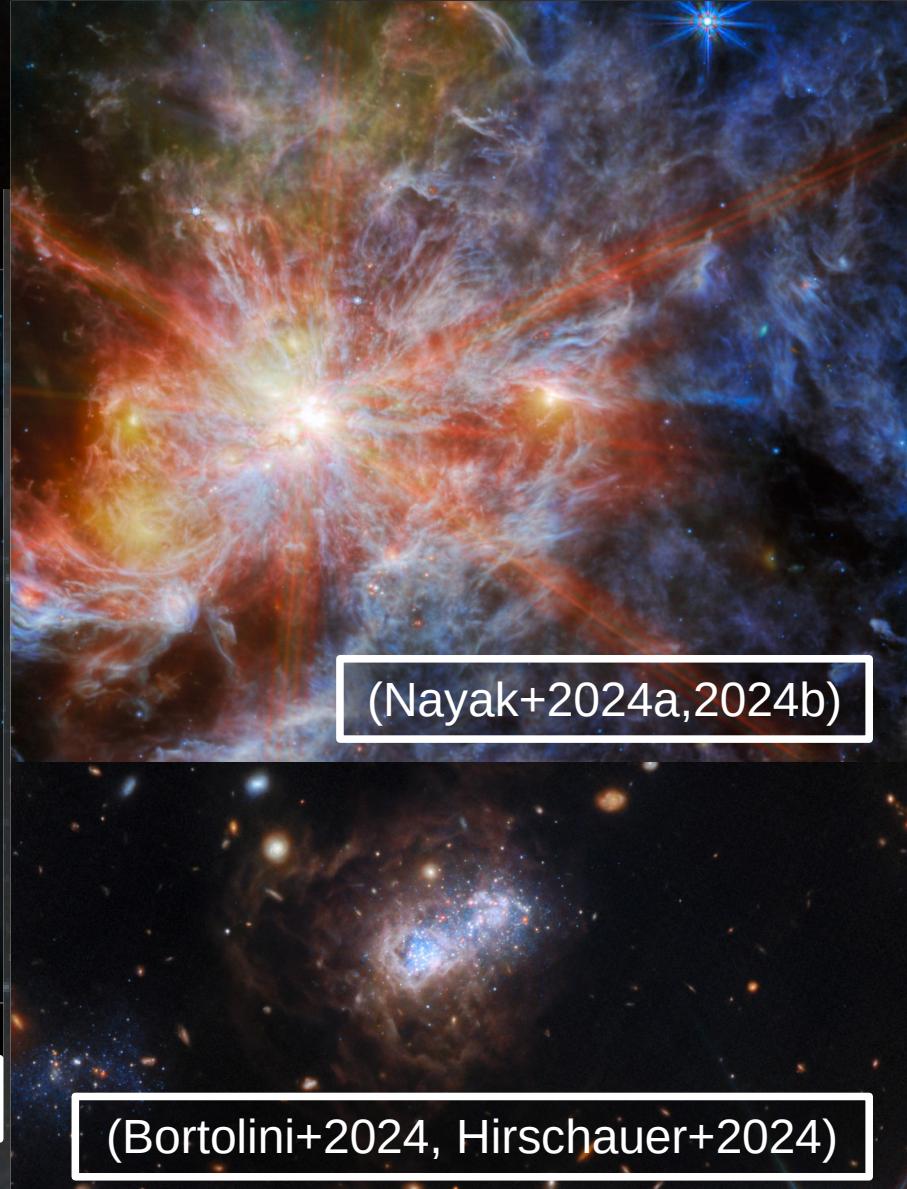


(Jones+2023)

JAMES WEBB SPACE TELESCOPE
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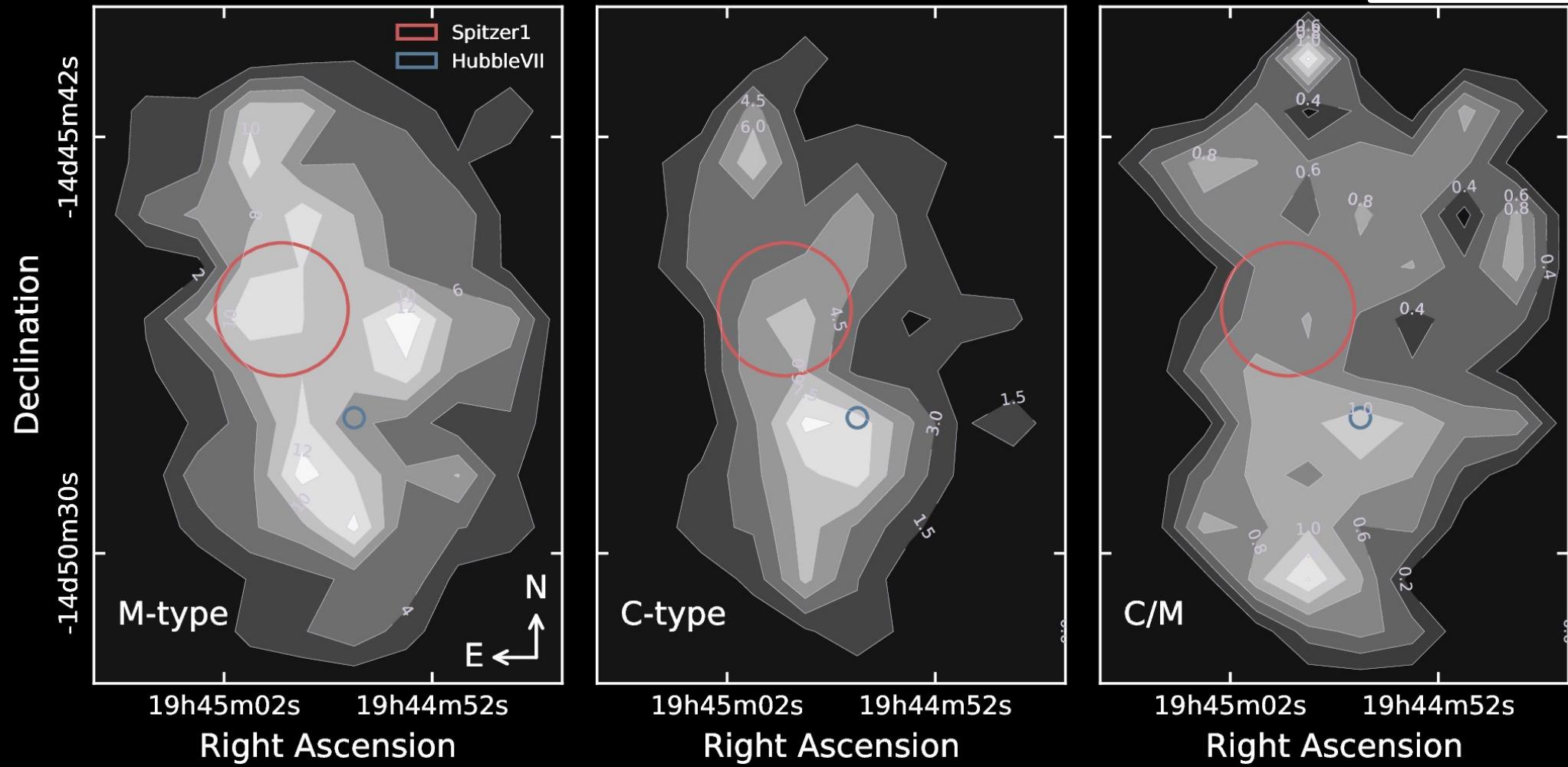
(Habel+2024)



(Nayak+2024a,2024b)

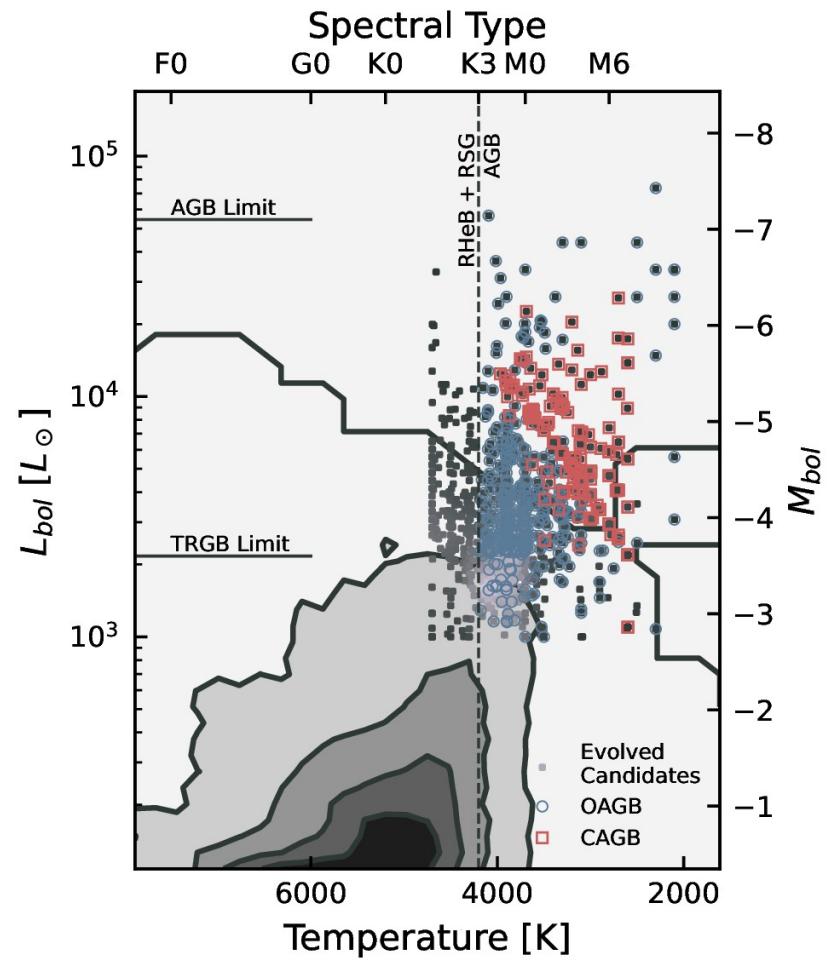
C/O Spatial Distribution

Preliminary



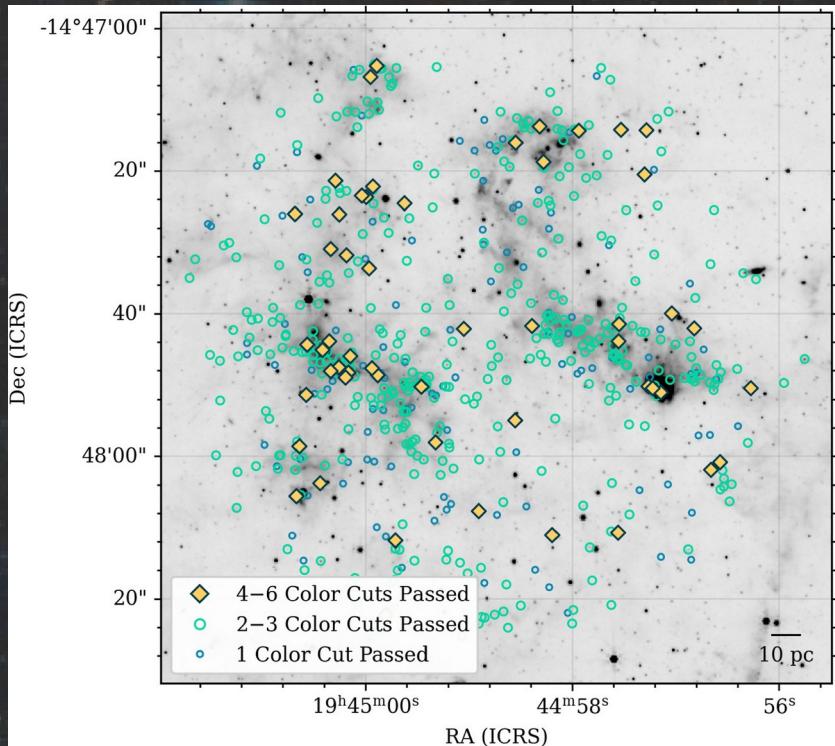
H-R Diagram

Preliminary



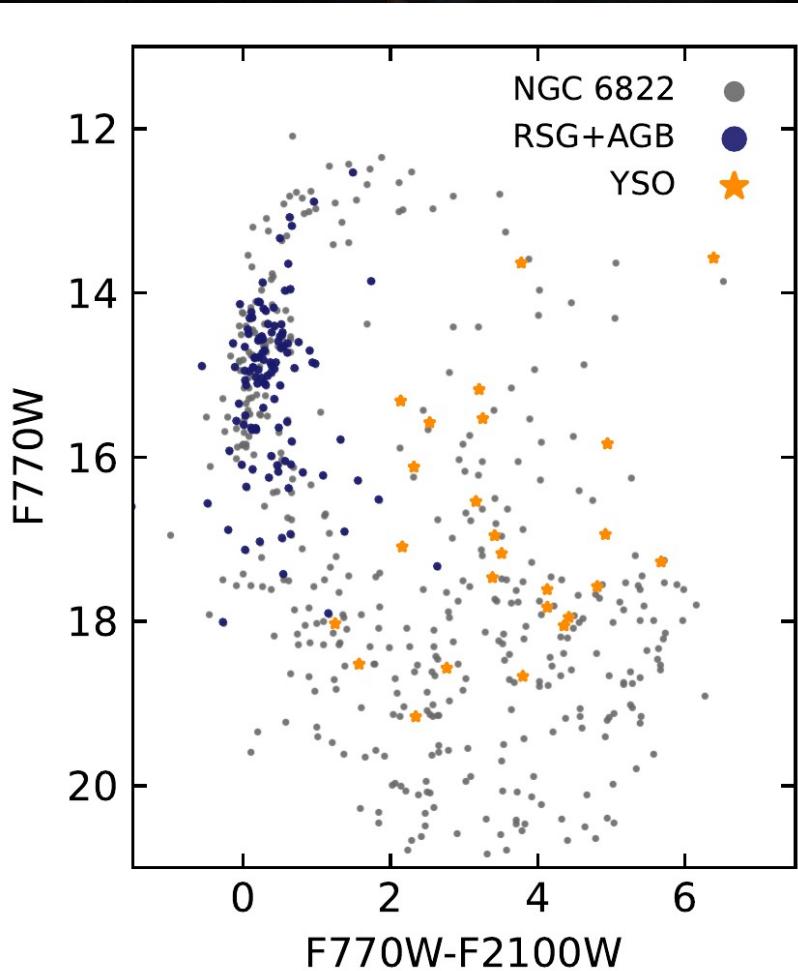
MIRI Catalogue - YSOs

- ★ 90 YSOs closely follow ISM dust structure in Spitzer massive star forming region

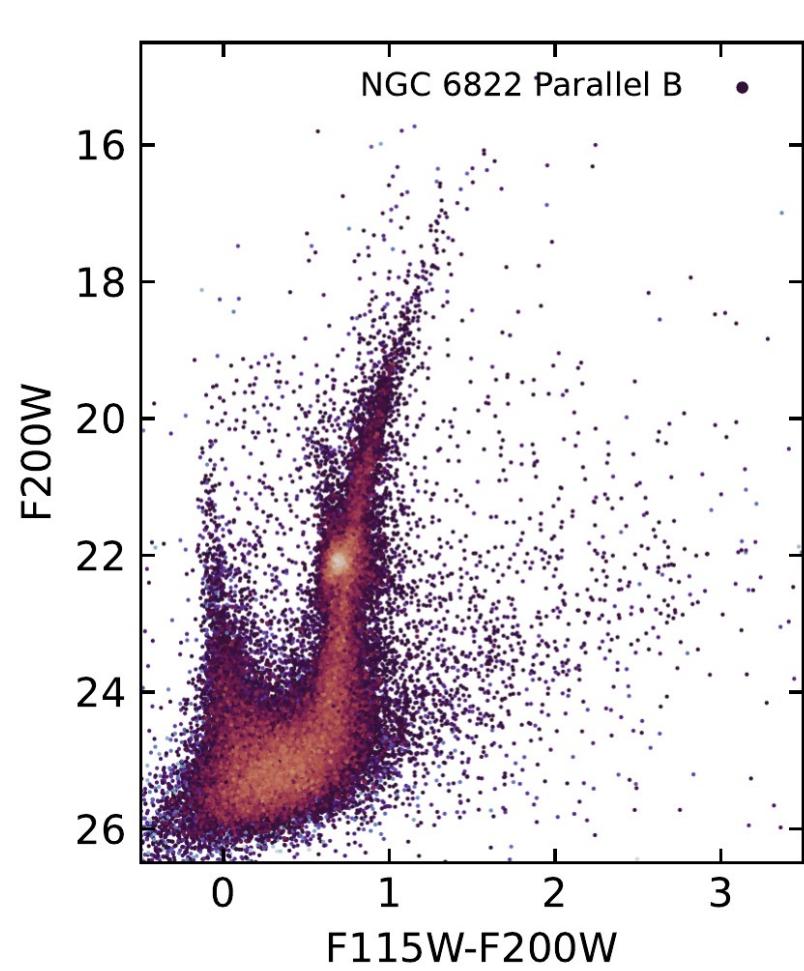
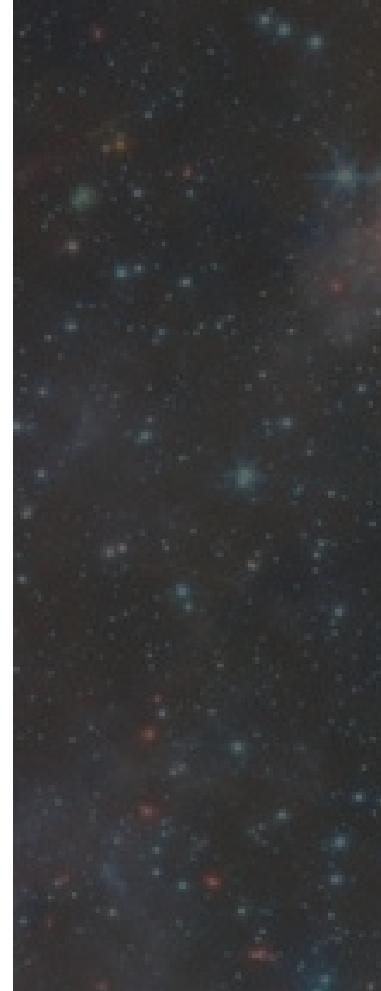
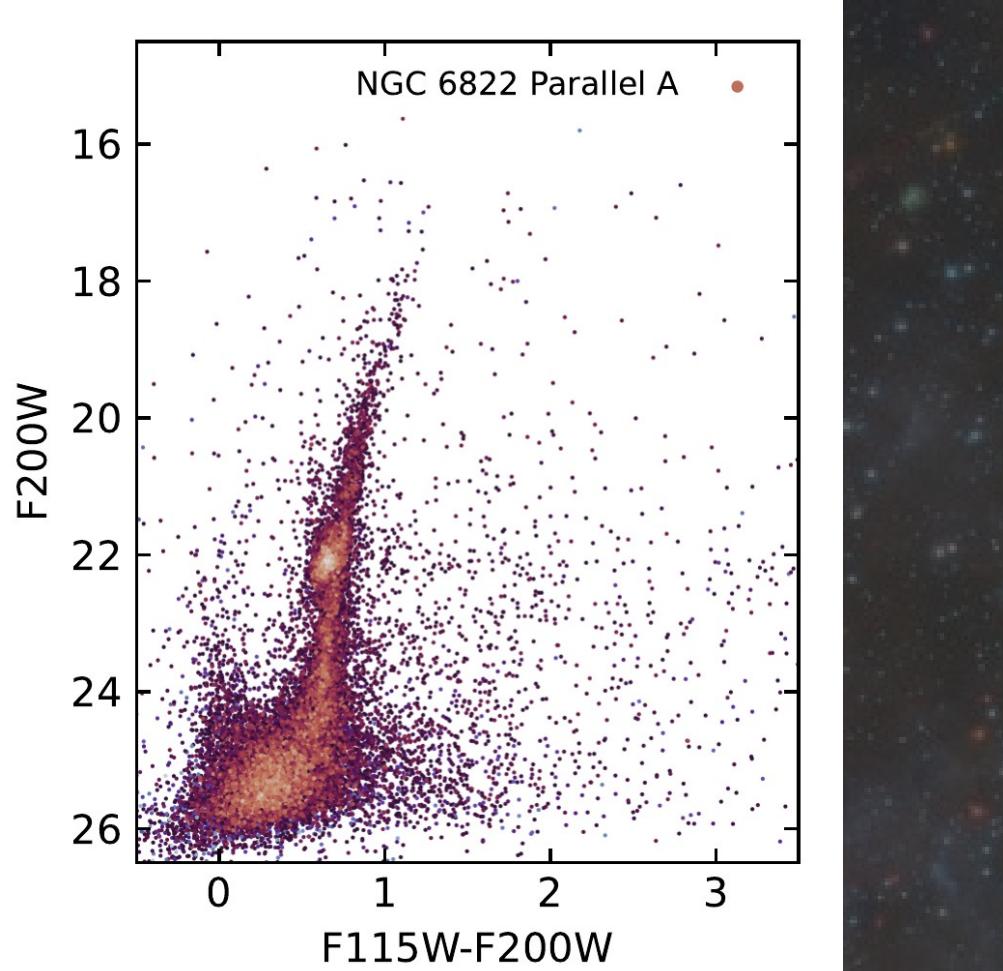


(Lenkic et al. 2024)

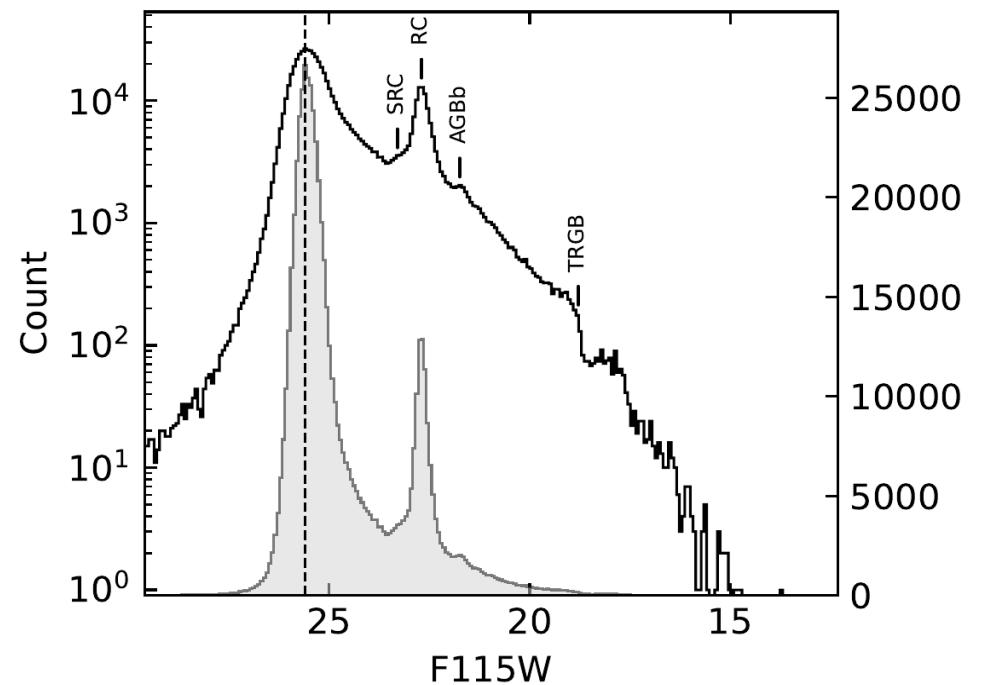
(Nally et al. 2024)



Parallel Fields



TRGB Measurement



Filter	Source	Completeness		Bright Limit	TRGB (error)
		count	[VegaMag]	[VegaMag]	[VegaMag]
<i>F115W</i>	792424	25.60		15.3	18.79(04)
<i>F200W</i>	539377	24.58		15.0	17.51(03)
<i>F356W</i>	184141	23.21		14.5	17.49(06)
<i>F444W</i>	155327	23.09		14.0	17.65(12)
<i>F770W</i>	97817	20.43		12.5	17.43(19)
<i>F1000W</i>	7098	19.41		11.5	17.40(13)
<i>F1500W</i>	1311	16.80		10.0	-
<i>F2100W</i>	794	16.86		9.0	-