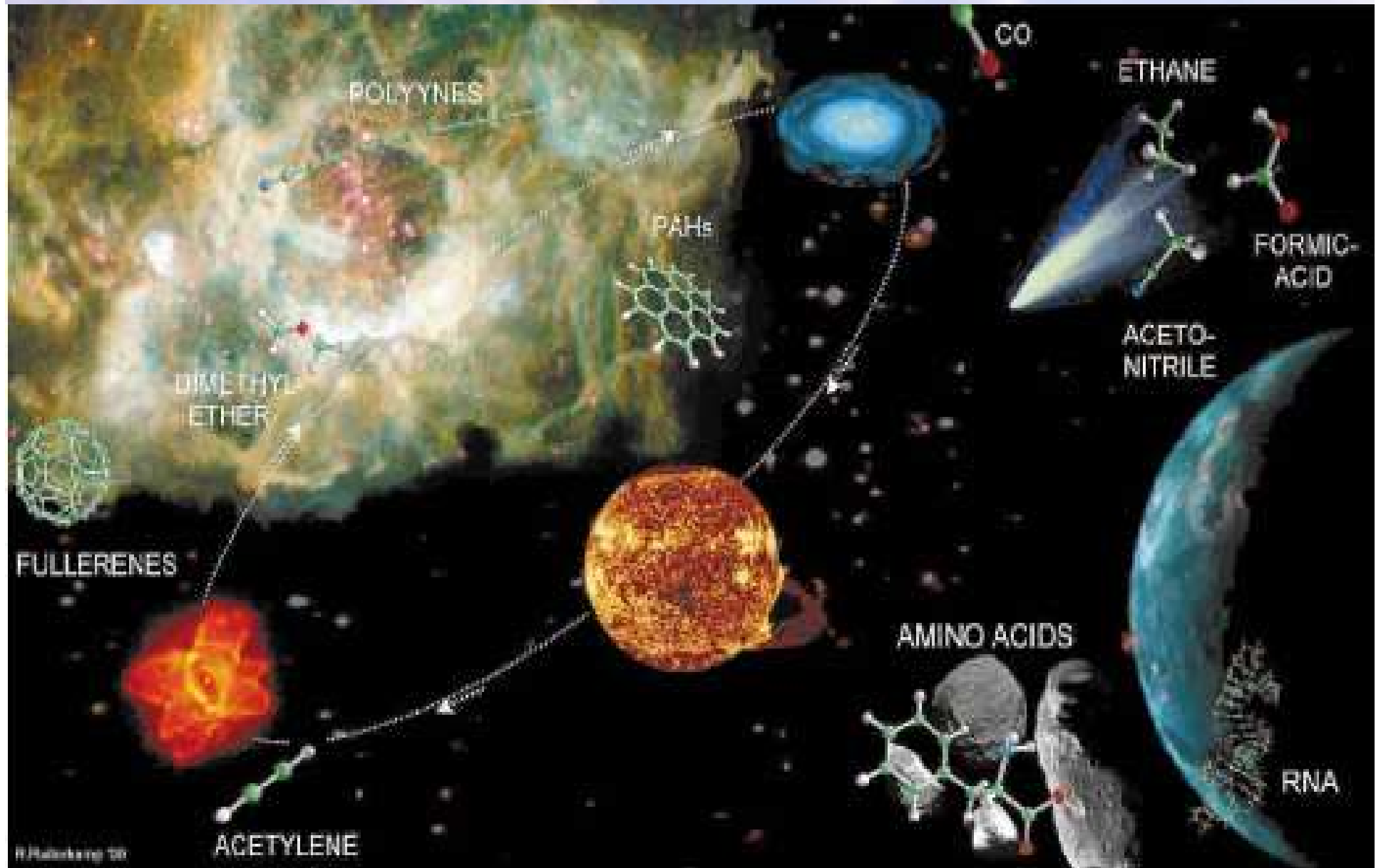


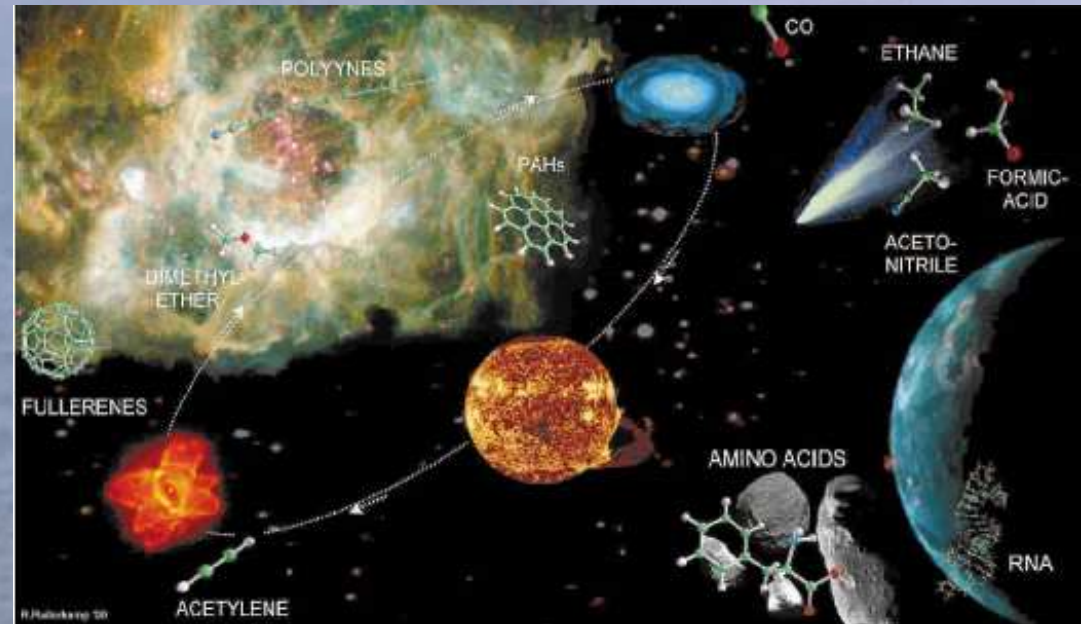
# ASTROCHEMISTRY & ASTROBIOLOGY



# Outline

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1. Astrochemistry & Meteoritic Organics
2. Extraterrestrial Delivery
3. Early Earth & Mars



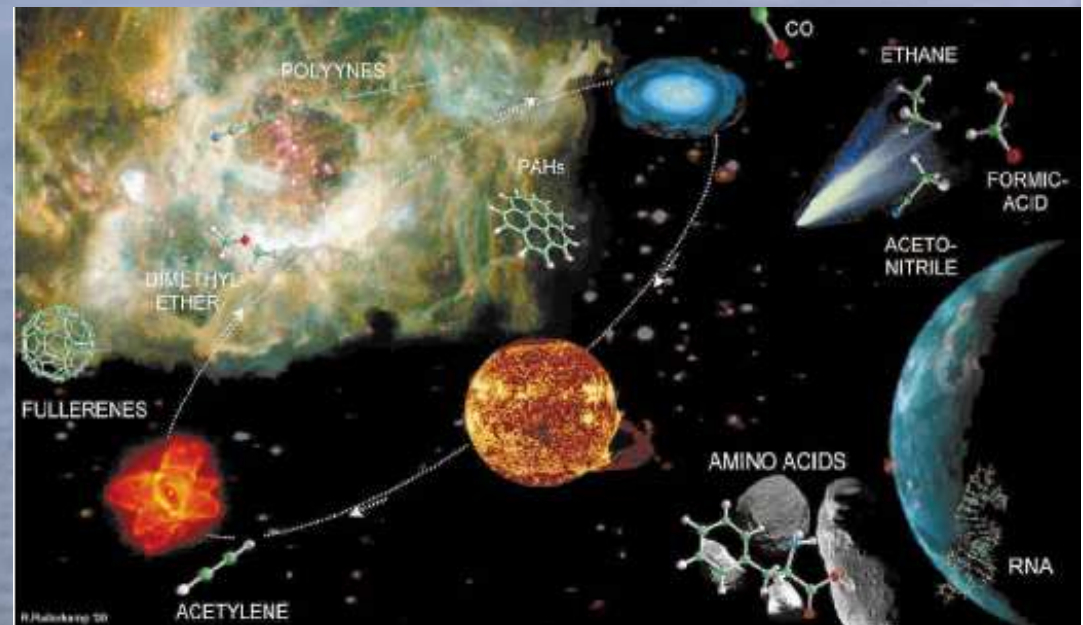
# Reading

Ehrenfreund et al. (2002) **Astrophysical and Astrochemical Insights into the Origin of Life**, Rep. Prog. Phys., 65, 1427-1487

Cronin, J.R. & Chang, S. (1993) **Organic Matter in Meteorites**, in The Chemistry of Life's Origins, J.M. Greenberg et al. (eds.), Kluwer, 209-258

Botta, O. & Bada, J.M. (2002) **Extraterrestrial Organic Compounds in Meteorites**, Surveys in Geophysics, 23, 411-467

Sephton, M.A. (2002) **Organic Compounds in Carbonaceous Meteorites**, Nat. Prod. Rep., 19, 292-311



# *Extraterrestrial Delivery of Biogenic Molecules*

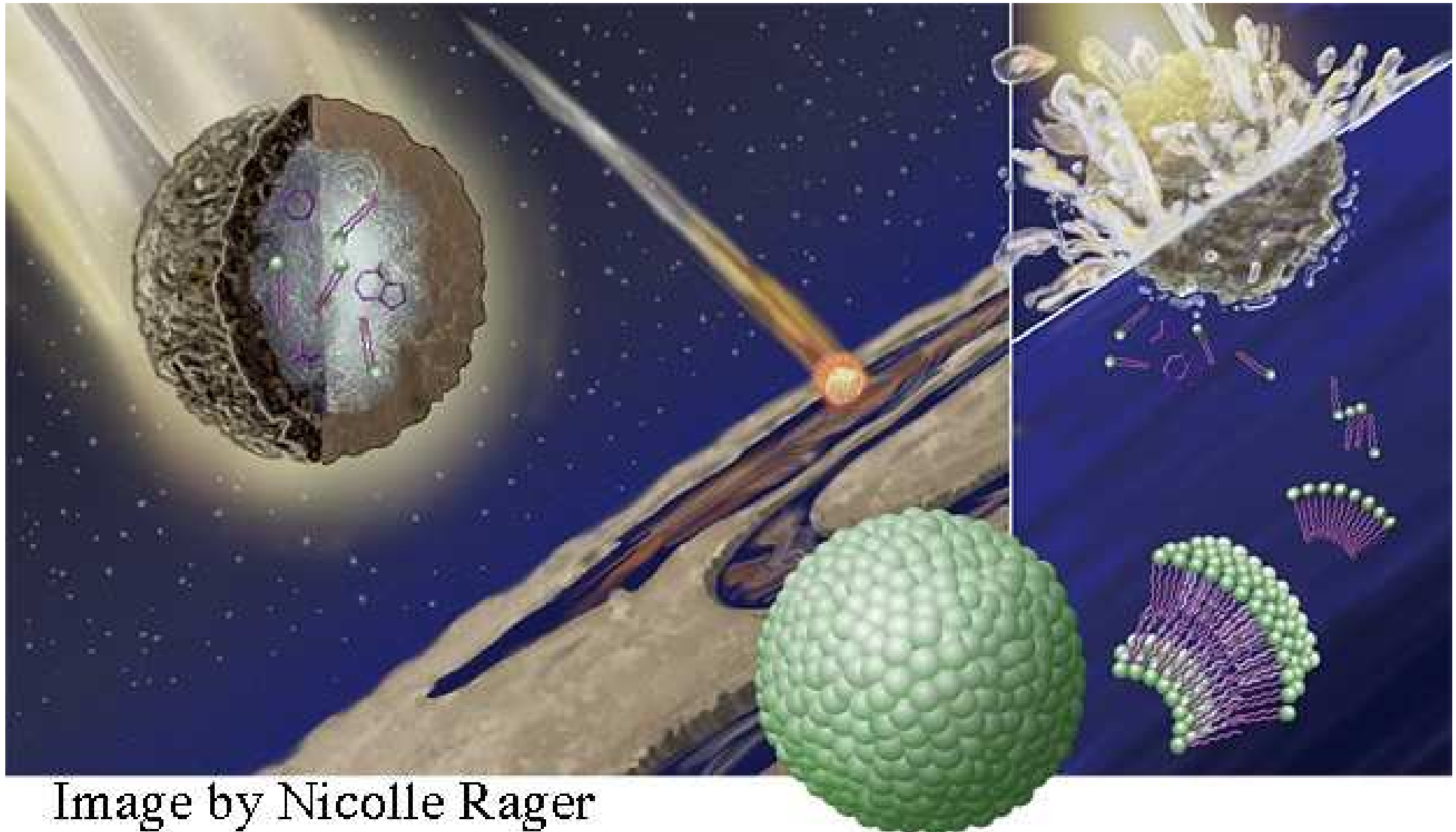


Image by Nicolle Rager

# Organics Found in Meteorites

Total Carbon Content: > 3% (by weight); Soluble Fraction: < 30% of total C

## COMPONENTS:

### ACIDS:

#### Amino acids

Carboxylic acids

Hydroxycarboxylic acids

Dicarboxylic acids

Hydroxydicarboxylic acids

Sulfonic acids

Phosphonic acids

### HYDROCARBONS:

non-volatile: aliphatic

aromatic (PAH)

polar

volatile

### FULLERENES:

$C_{60}$ ,  $C_{70}$

He@ $C_{60}$

Higher Fullerenes

### OTHERS:

#### N-Heterocycles

Amides

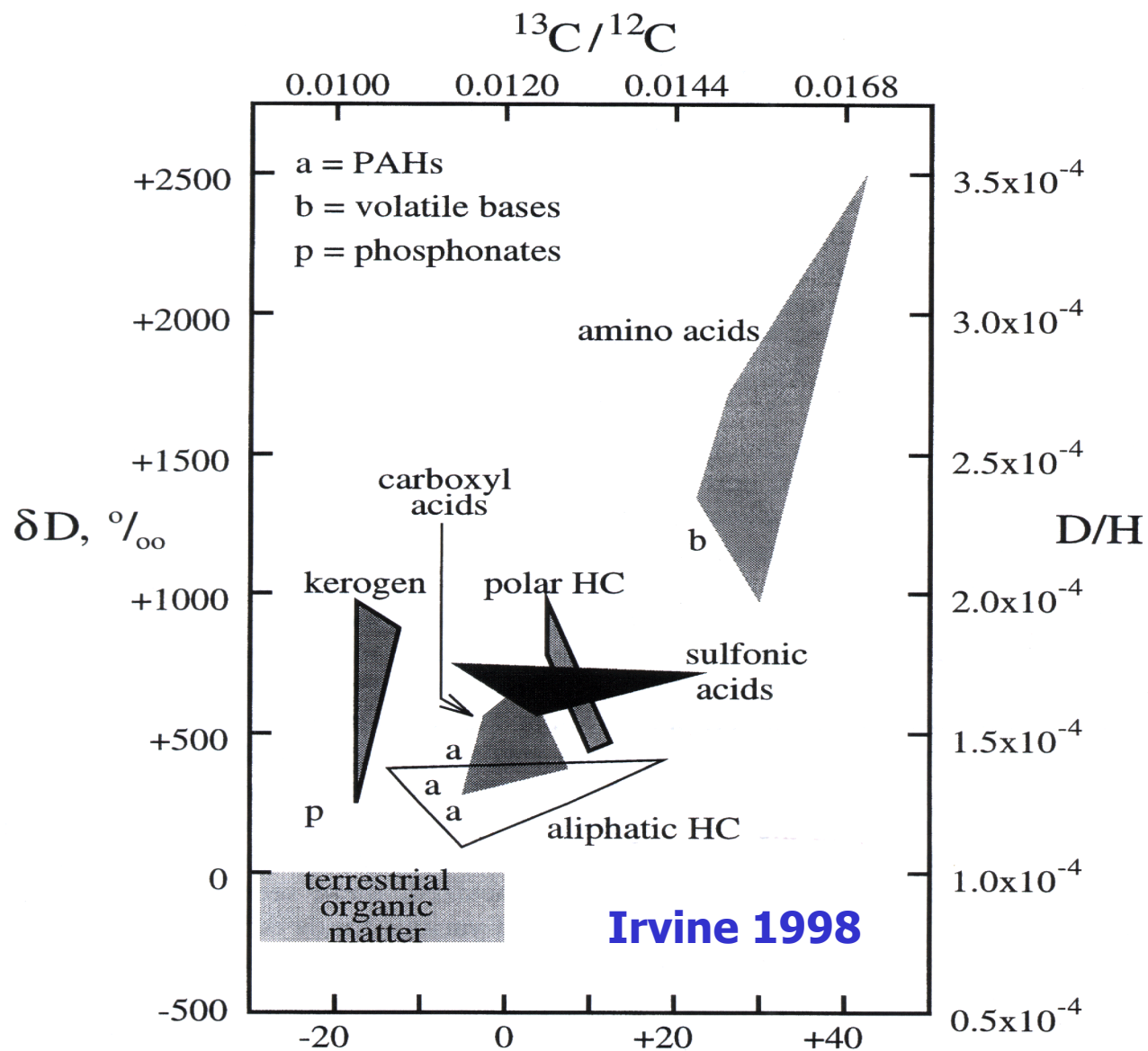
Amines

Alcohols

Carbonyl compounds



# ISOTOPIC RATIOS FOR "C" AND "H"



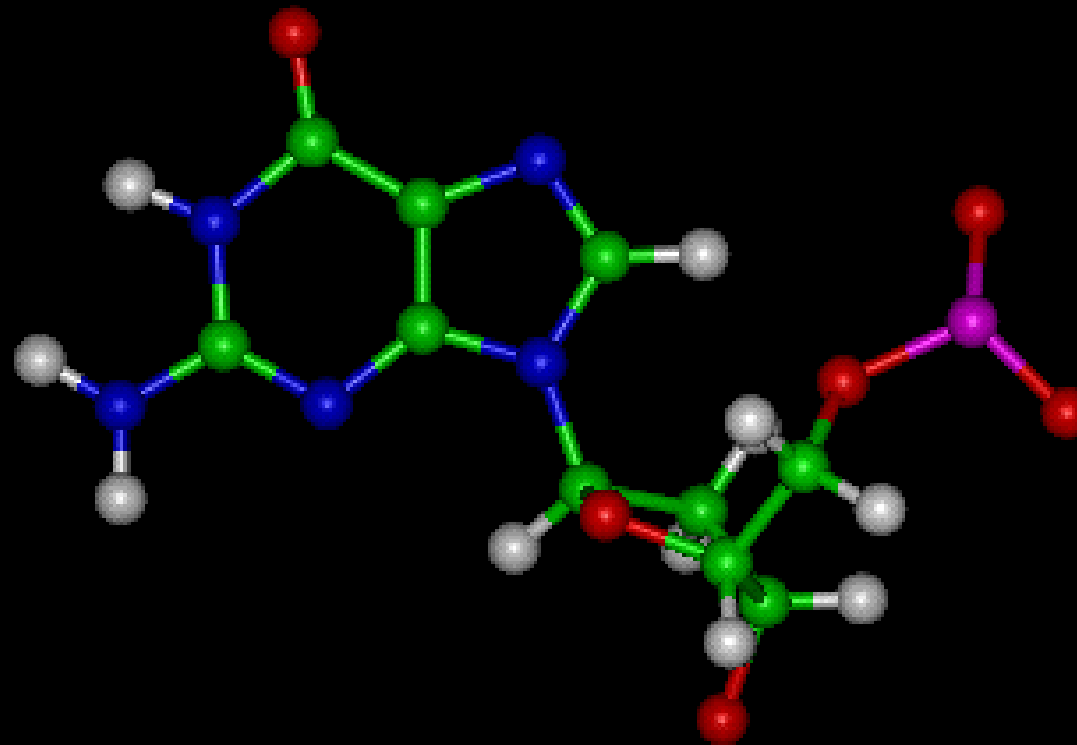
Terr.ocean =  $\delta\text{D} = 0$

$\delta^{13}\text{C}, \text{‰}$

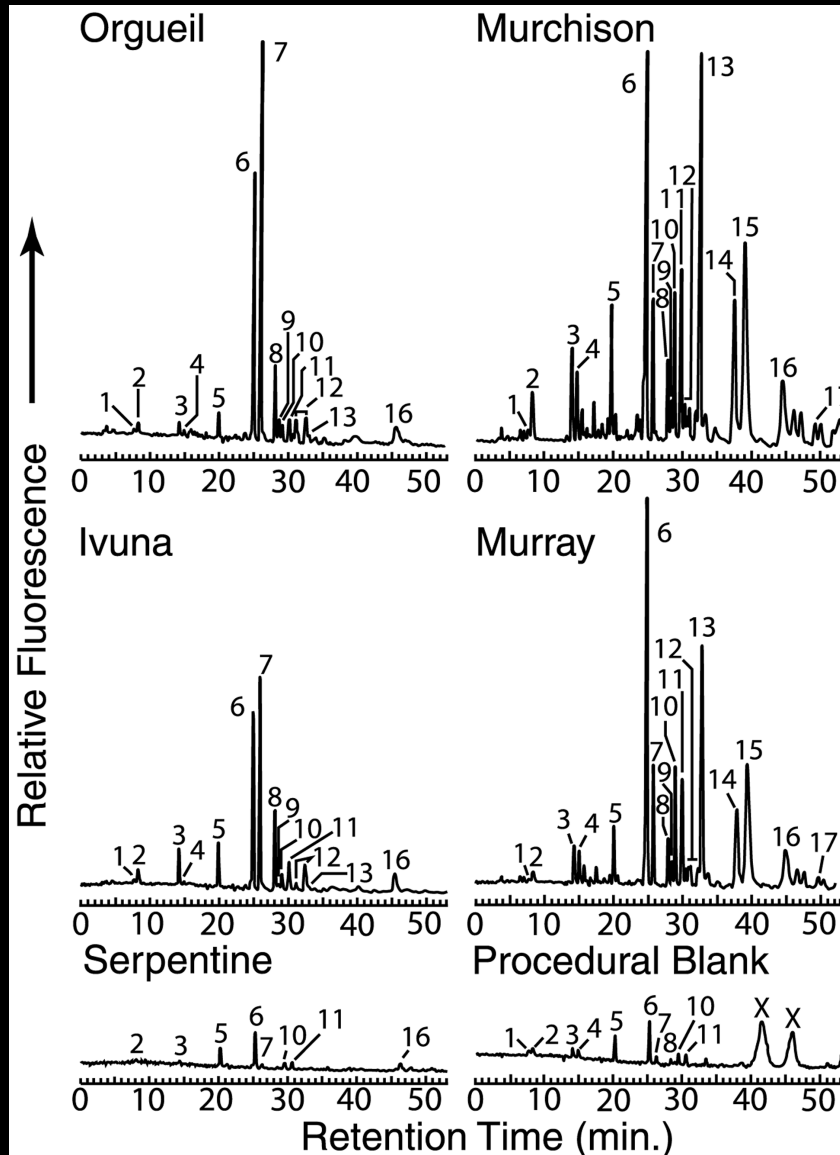
Cosmic D/H ratio  $\sim 0.8\text{-}2 \times 10^{-5}$

# DNA/RNA Components

---



# Chromatograms of Meteorite Extracts



- 1 D-Aspartic Acid
  - 2 L-Aspartic Acid
  - 3 L-Glutamic Acid
  - 4 D-Glutamic Acid
  - 5 D,L-Serine
  - 6 Glycine
  - 7  $\beta$ -Alanine
  - 8  $\gamma$ -Amino-n-butyric Acid (g-ABA)
  - 9 D,L-b-Aminoisobutyric Acid (b-AIB)
  - 10 D-Alanine
  - 11 L-Alanine
  - 12 D,L- $\beta$ -Amino-n-butyric Acid (b-ABA)
  - 13  $\alpha$ -Aminoisobutyric Acid (AIB)
  - 14 D,L- $\alpha$ -Amino-n-butyric Acid (a-ABA)
  - 15 D,L-Isovaline
  - 16 L-Valine
  - 17 D-Valine
  - X: unknown
- Ehrenfreund *et al.*, 2001**

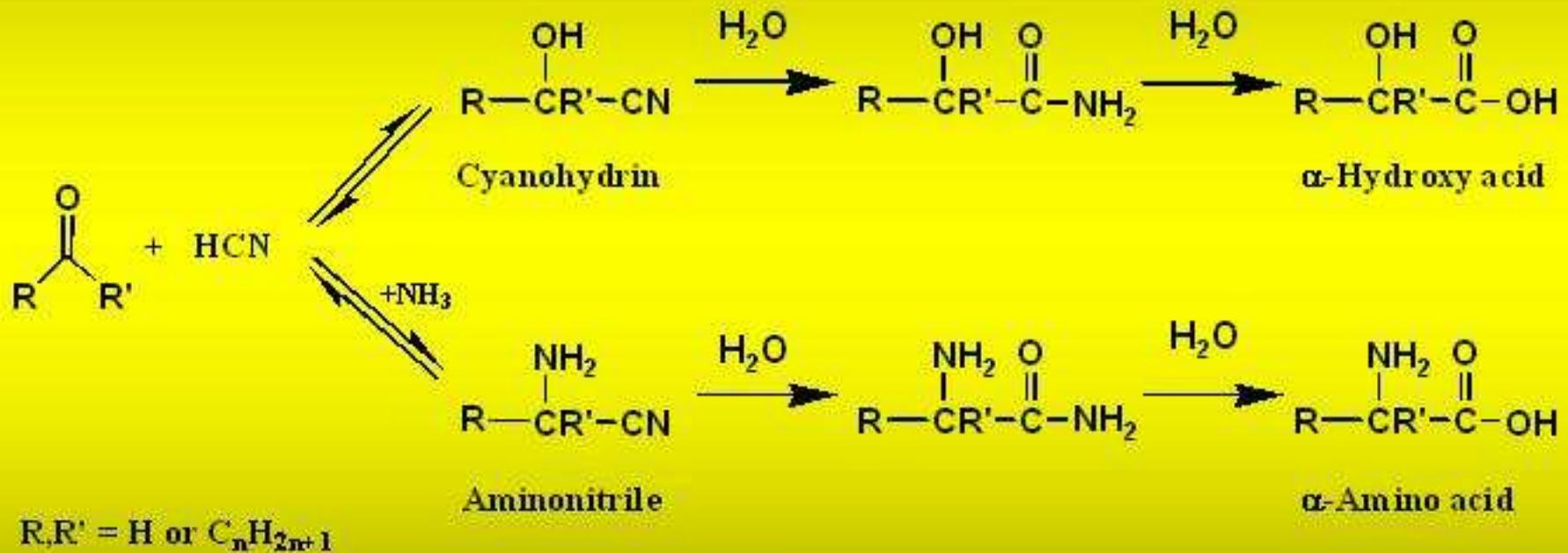


# Amino Acids in Carbonaceous Chondrites

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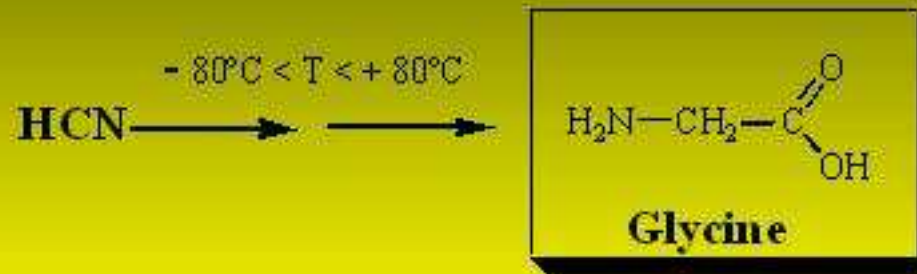
- **Amino acids are readily synthesized under a variety of plausible prebiotic conditions (e.g. in the Miller-Urey Experiment).**
- **Amino acids are the building blocks of proteins and enzymes in life on Earth.**
- **Chirality (handedness) can be used to distinguish biotic vs. abiotic origins.**
- **Most of the amino acids found in meteorites are very rare on Earth (AIB, isovaline).**

# Strecker Amino Acid Synthesis in CM-type Chondrites



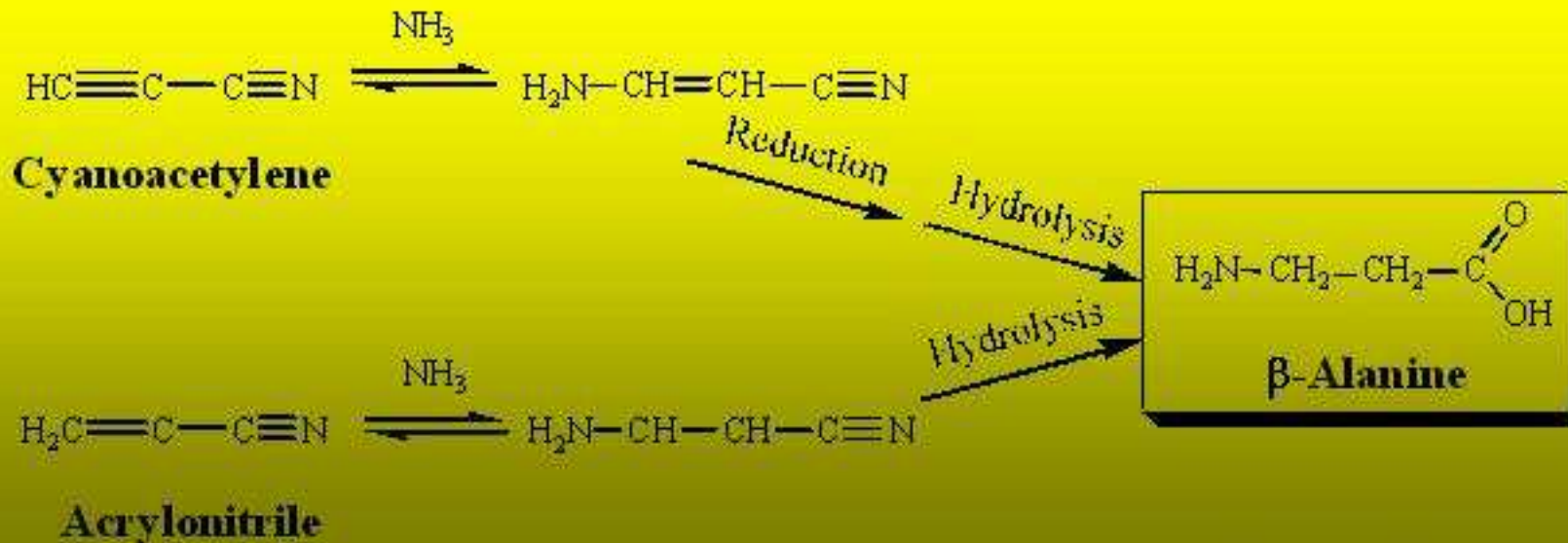
For a review: Botta and Bada, *Surv. Geophys.* 23, 411-467 (2002)

# Amino Acid Synthesis in CI-type Chondrites

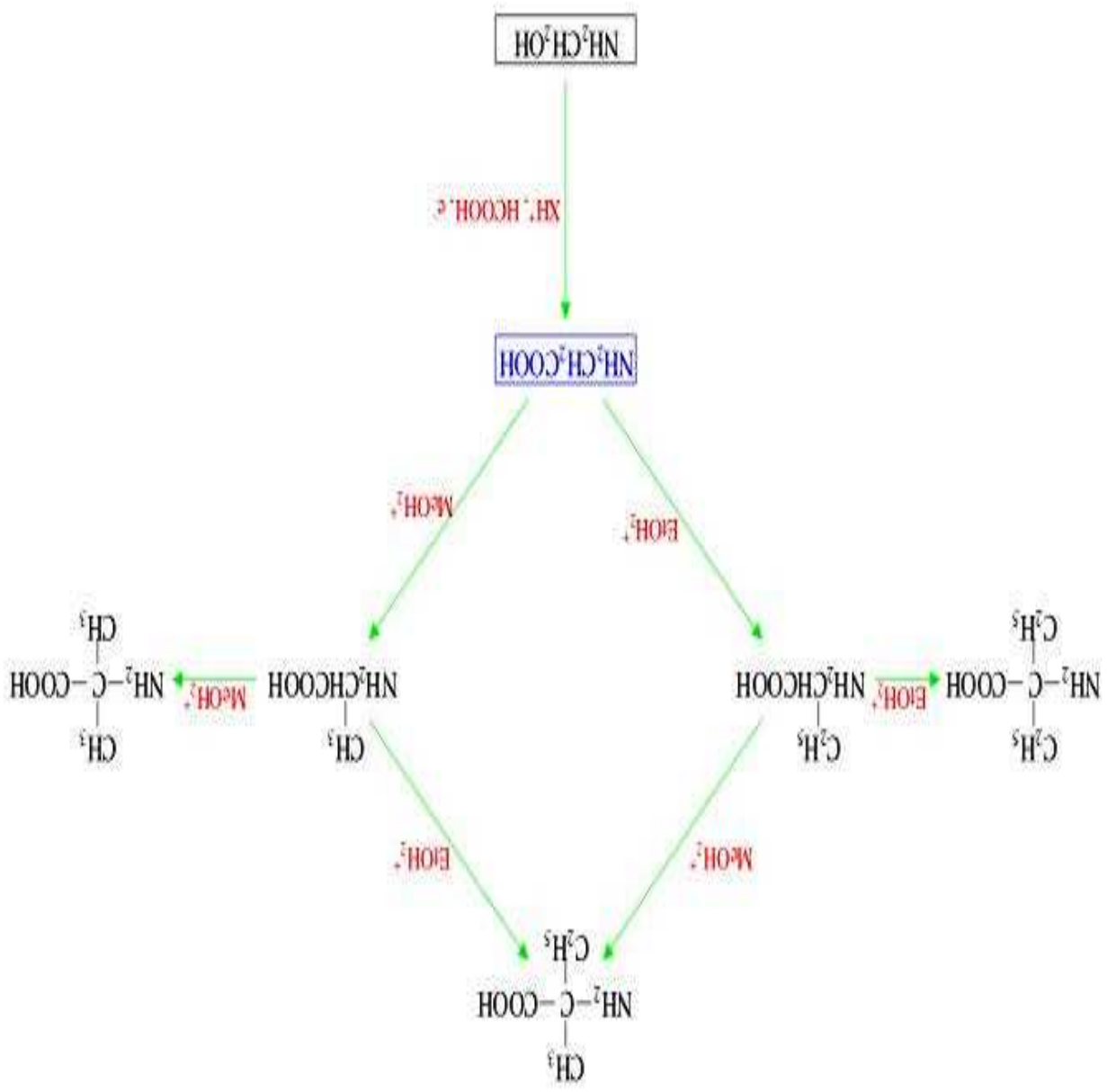


+ traces of other amino acids

*Levy et al., Icarus 145, 609-613 (2000).*

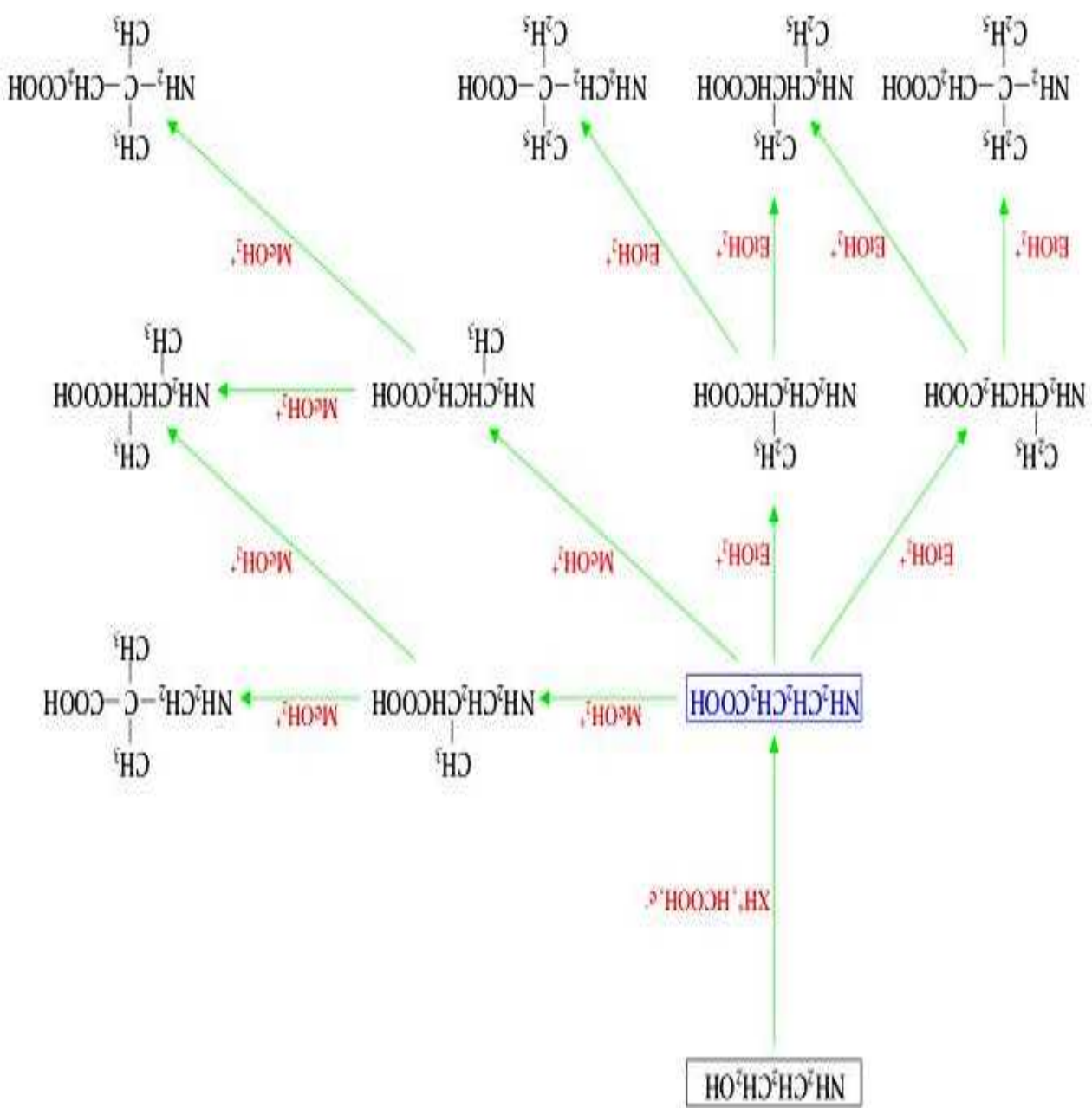


*Cronin & Chang, 1993.*

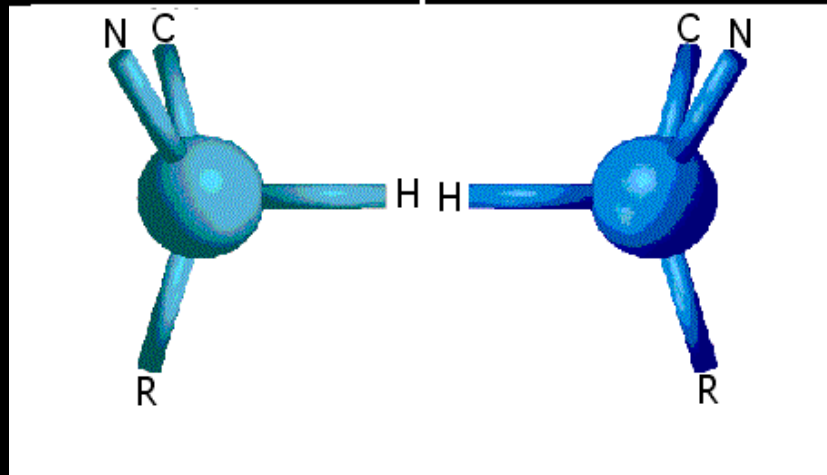


ALKYL CATION TRANSFER TO GLYCINE

# ALKYL CATION TRANSFER TO β-ALANINE



# What is Chirality?



- **Left- and right-handed mirror molecules are called enantiomers.**
- **Enantiomers possess identical physical properties (melting point *etc.*).**
- **They rotate the plane of planar-polarized light in opposite directions.**
- **They cannot be chromatographically separated on a non-chiral column.**



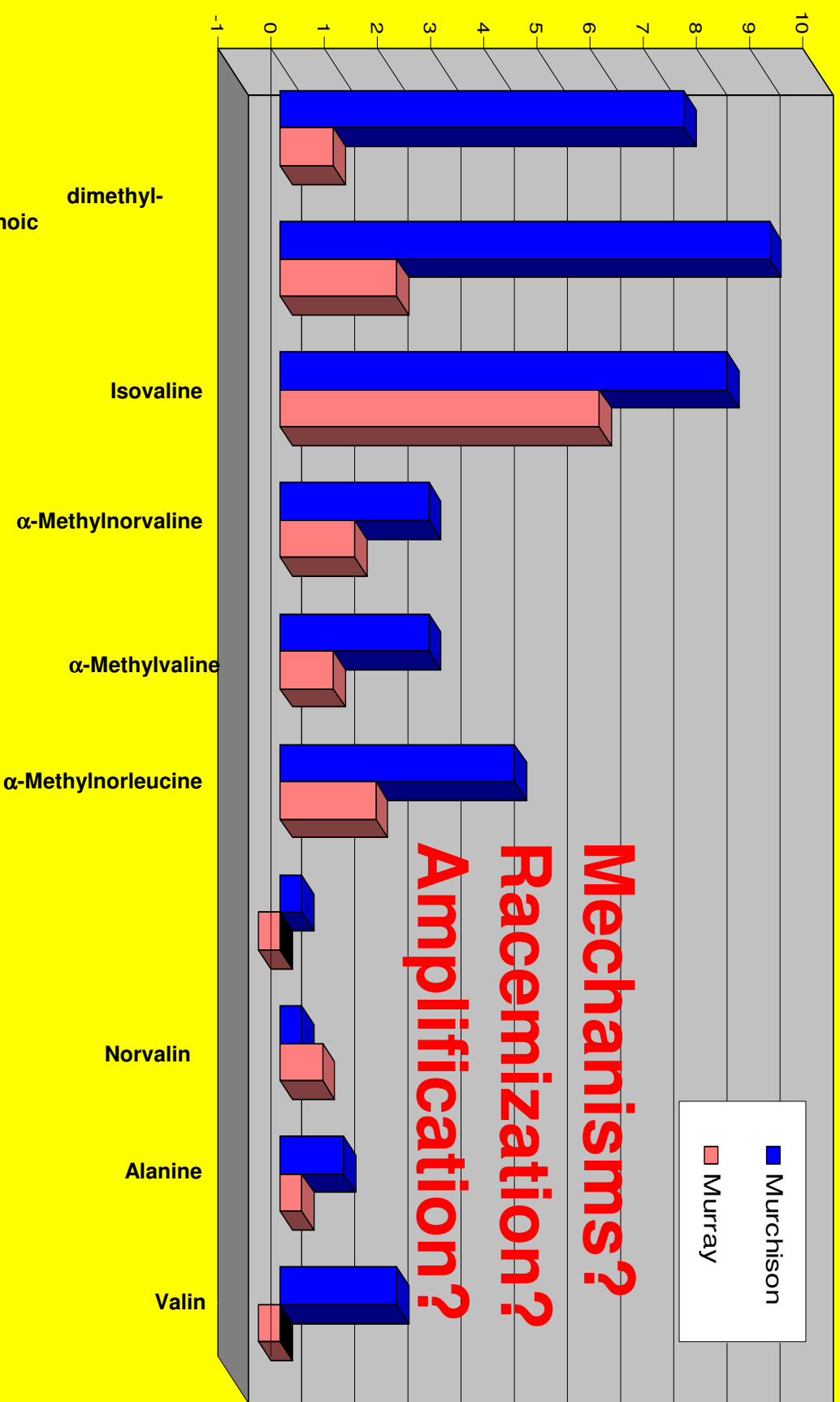
**Separation on chiral column**

**or**

**Derivatization to form diastereoisomers, separation on non-chiral column**



# Enantiomeric Excesses in Meteoritic Amino Acids

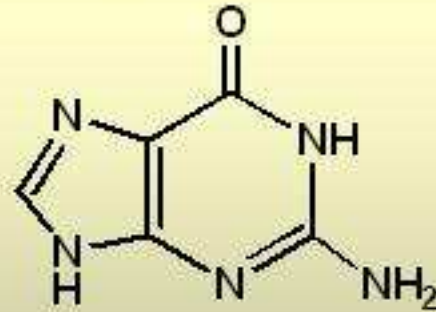


Pizzarello and Cronin, *Geochim. Cosmochim. Acta* **64**, 329-338 (2000)

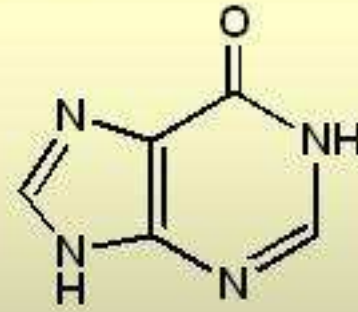
# Nucleobases in Carbonaceous Chondrites



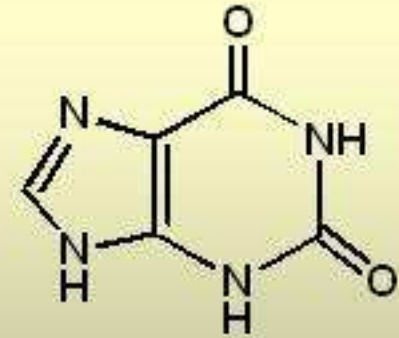
Adenine



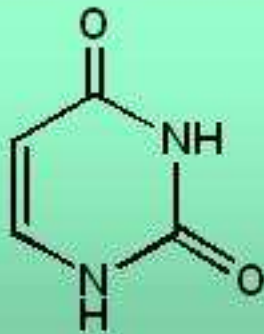
Guanine



Hypoxanthine



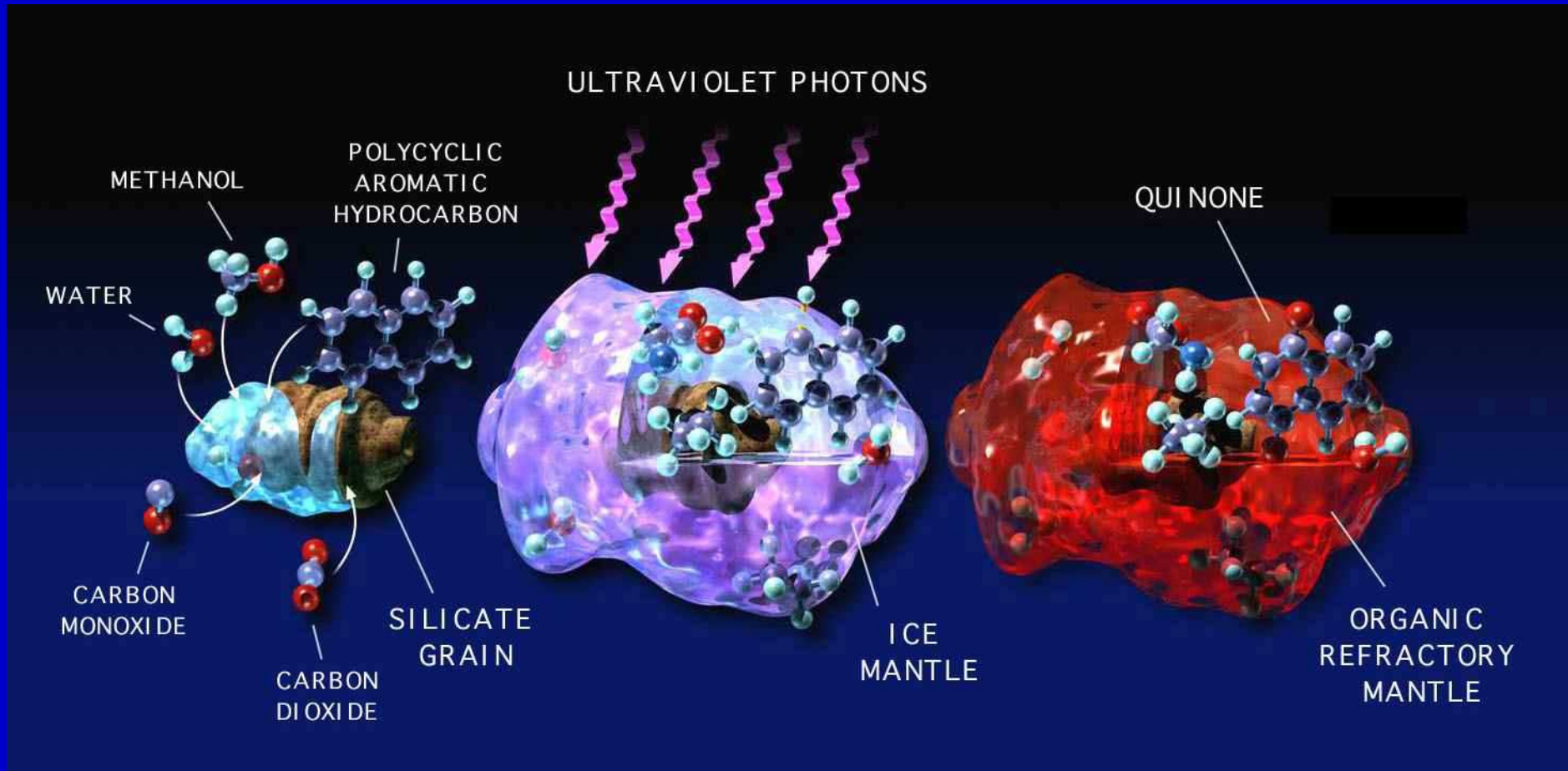
Xanthine



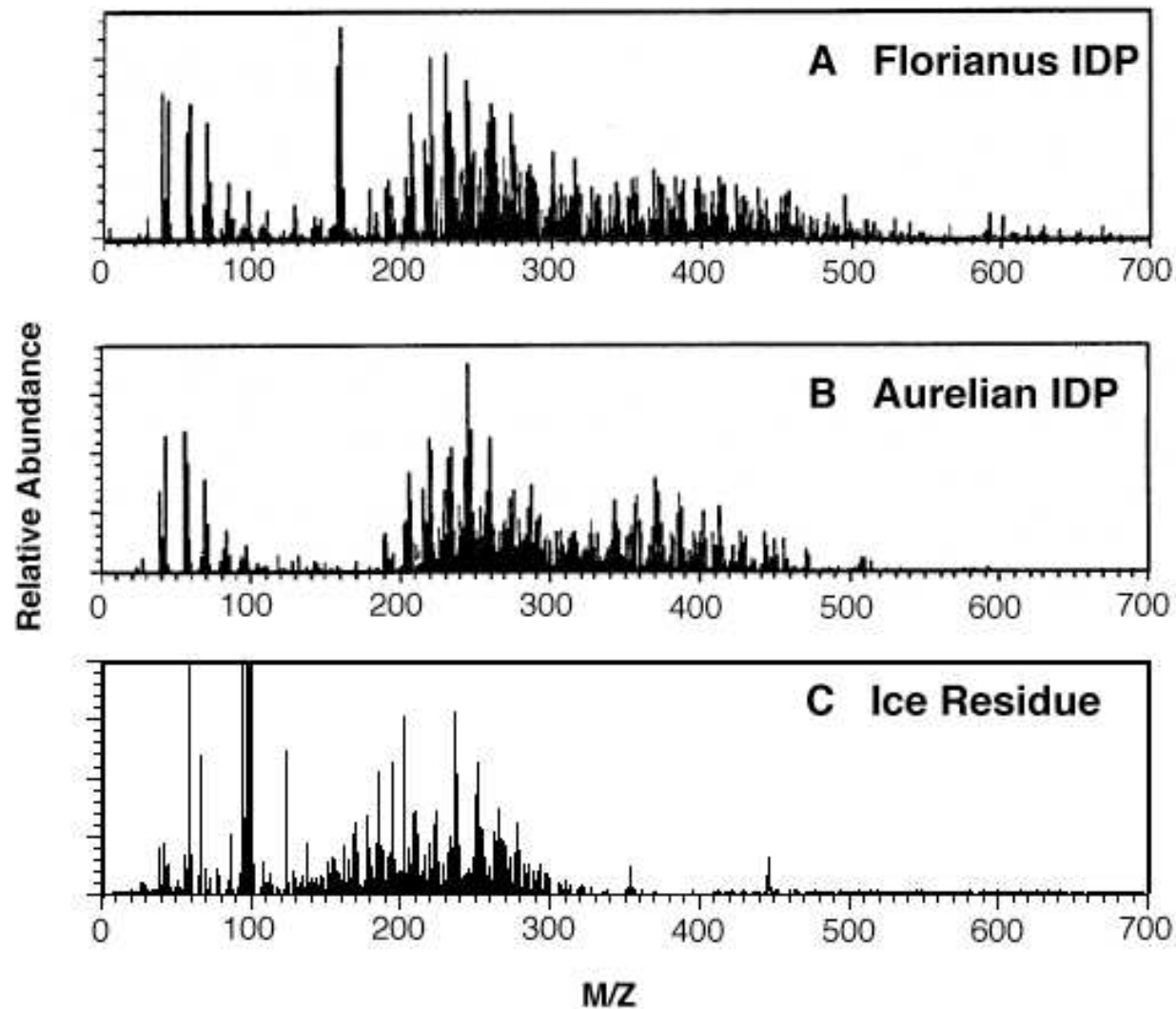
Uracil

- are very important in the replicating system of all known terrestrial organisms (in DNA and RNA)
- have been detected in Murchison, Murray and Orgueil meteorites at the 200-500 ppb level (Schwartz and coworkers, 1979-1982)
- various other (non-biogenic) N-heterocycles, including a variety of alkylated pyridines, were found in meteorites
- no isotopic measurements have been reported

# *Interstellar Dust: ice mantle evolution*



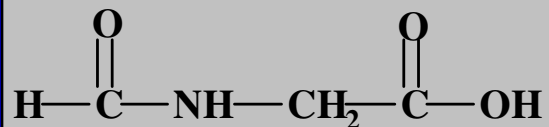
Mass Spectrum of the Room Temperature **Residue** of  $\text{H}_2\text{O}:\text{CH}_3\text{OH}:\text{CO}:\text{NH}_3$  (100:50:1:1) Ice Compared to the Mass Spectra of Two Interplanetary Dust Particles (IPDs)



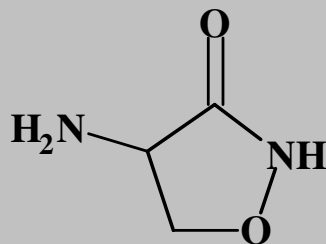
IDP Spectra - Clement et al., 1993

# *Interstellar/Precometary Ice Photolysis: Abiotic*

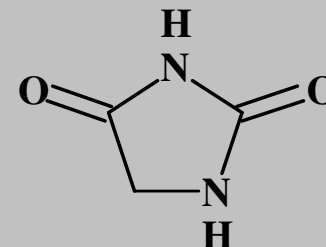
## *Synthesis of Important Prebiotic Organics*



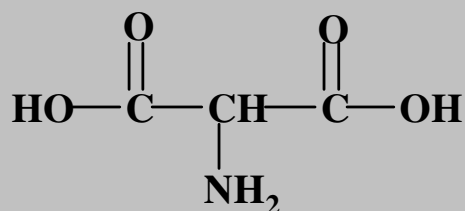
N-Formyl glycine



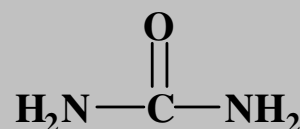
Cycloserine



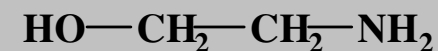
Hydantoin



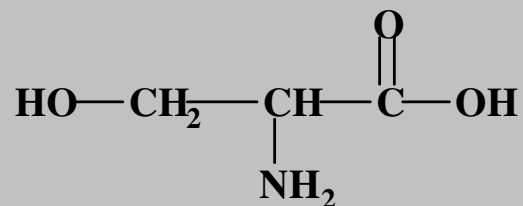
Amino malonic acid



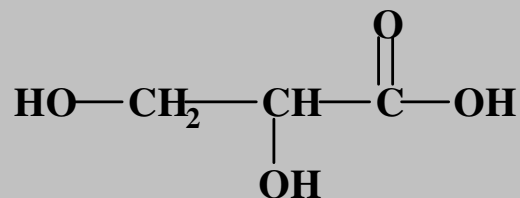
Urea



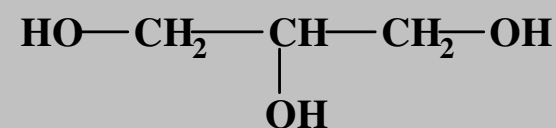
Ethanolamine



Serine

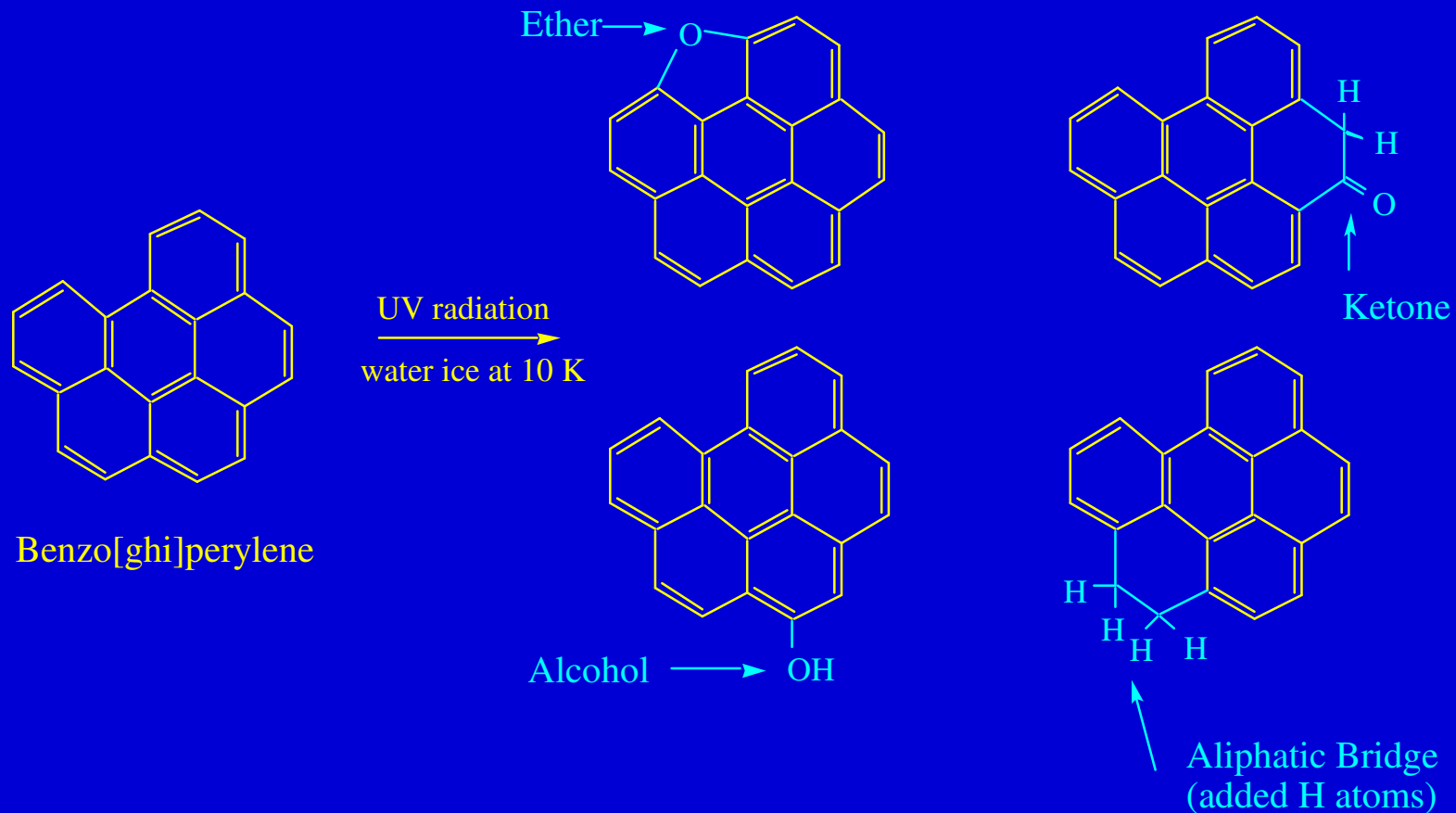


Glyceric Acid



Glycerol

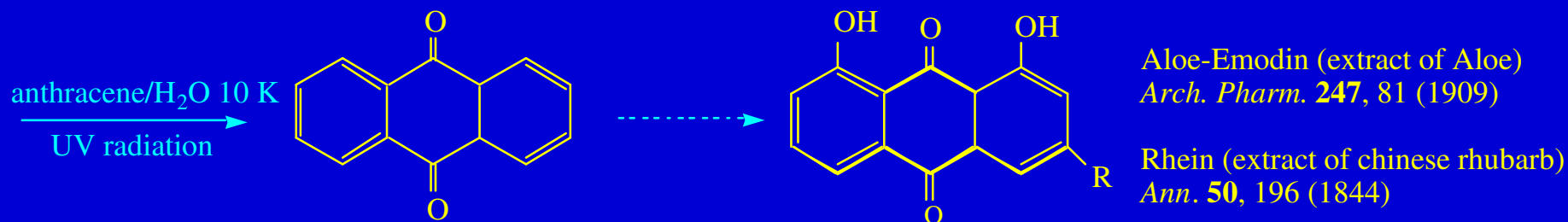
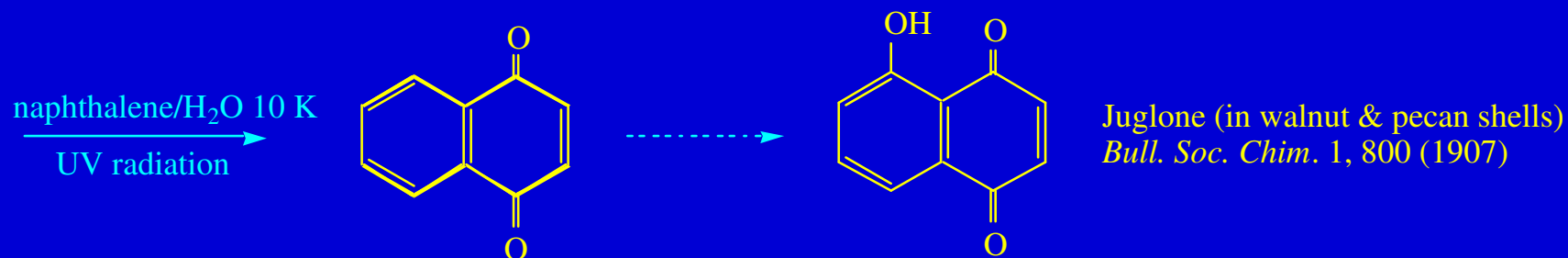
# Photolysis of PAHs in Water Ice Produces Alkanes, Ketones, Ethers, and Alcohols.



*Both oxidation (alcohol, ether and ketone formation) and reduction (addition of hydrogen) reactions occur on photolysis of water ices containing PAHs. These are the same kinds of compounds observed in meteorites, fit spectra of emission objects and, in some cases, have biochemical significance.*



# Photochemistry of PAHs in Ice: Abiotic Synthesis of Biogenic Compounds

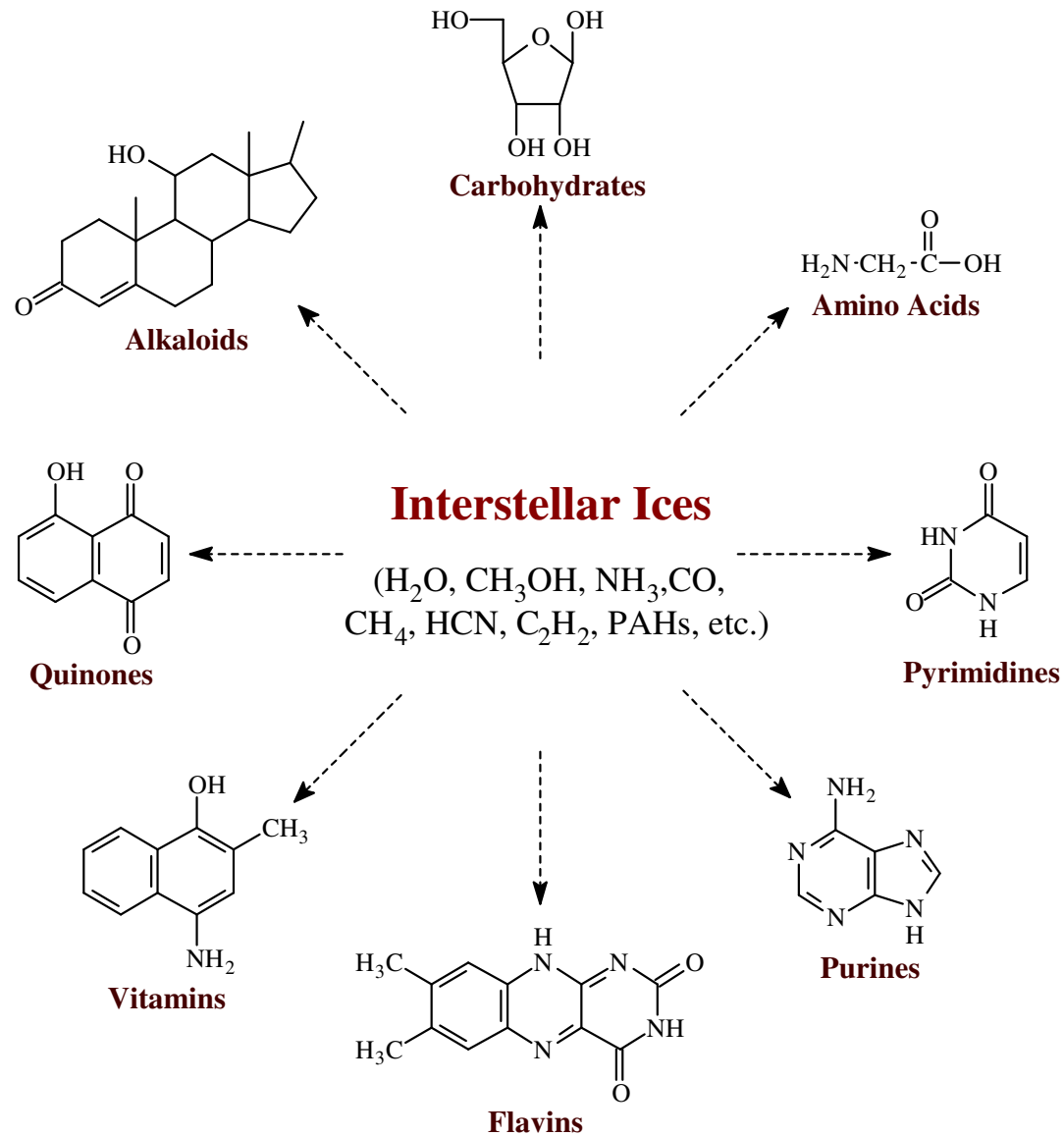


*These quinone type structures are very important in many living systems. For example, naphthaquinones (like juglone above) are essential for electron transport in simple organisms.*

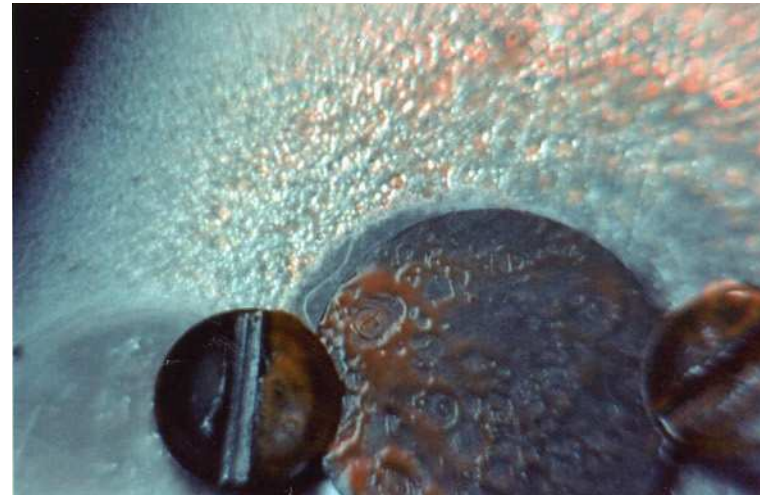
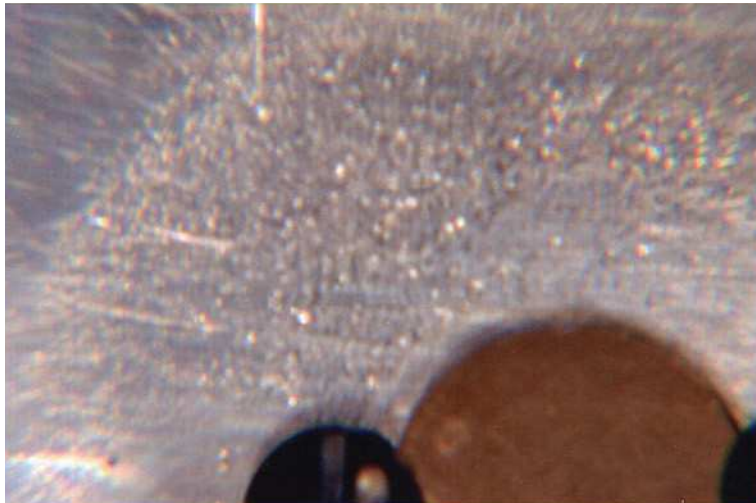
*Juglone: Bernstein et al. *Met. & Planet. Sci.* 36, 351 (2001)*

*Anthroquinone: Ashbourne et al. in prep*

# Abiotic Synthesis of Biogenically Useful Molecules in Cometary and Interstellar Ices



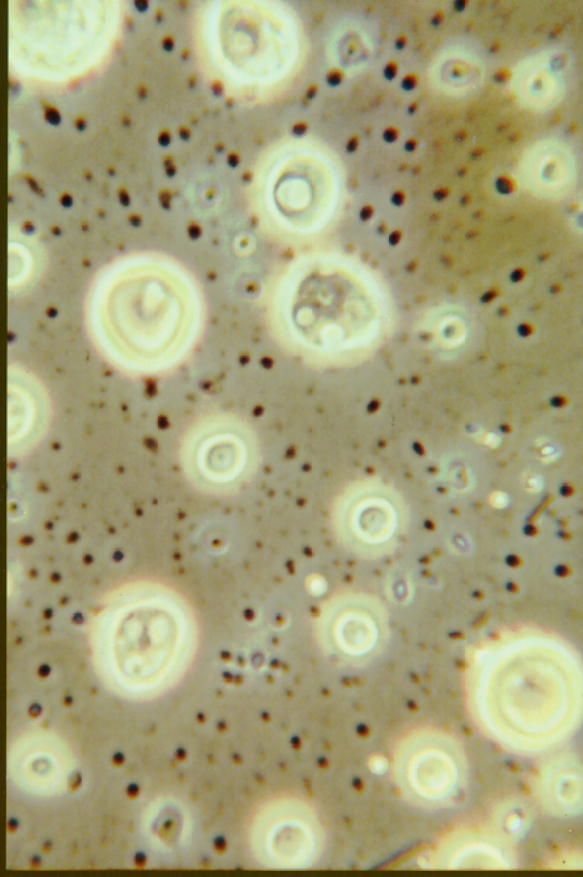
# Organic Residue Remaining After the Low Temperature UV Irradiation of the Ice





SELF ORGANIZED WATER INSOLUBLE DROPLETS  
PRODUCED WHEN THE ORGANIC RESIDUE FROM A  
UV RADIATED INTERSTELLAR / PRE-COMETARY ICE  
ANALOG IS ADDED TO WATER

*Visible Light*

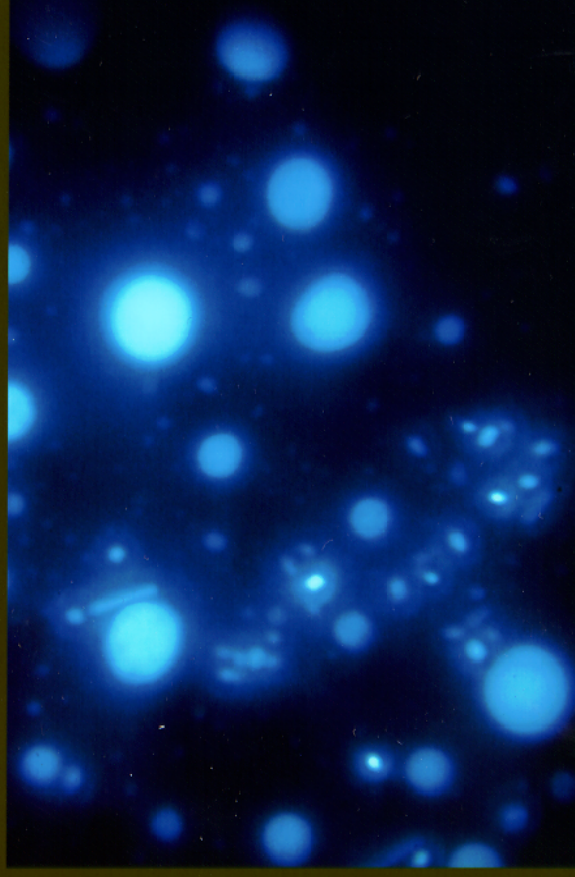


Smallest substructures are  
about 5 microns across

≈10 Microns

Initial Ice Composition=  
 $\text{H}_2\text{O}:\text{CH}_3\text{OH}:\text{CO}:\text{NH}_3$   
(100:50-10:10:10)

*Fluorescence* ( $\lambda_{\text{em}} \approx 410\text{nm}$ ;  $\lambda_{\text{ex}} \approx 330\text{nm}$ )

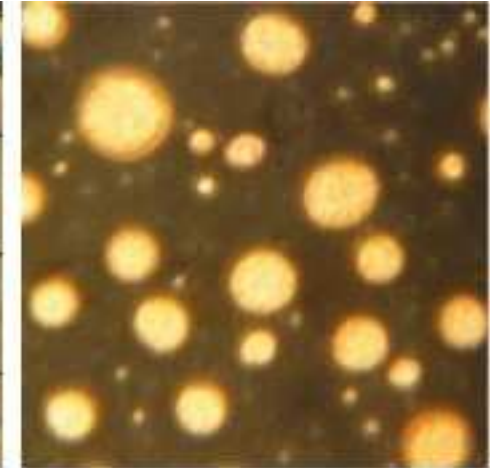
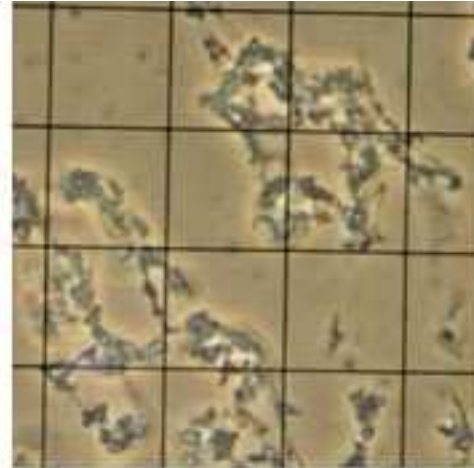
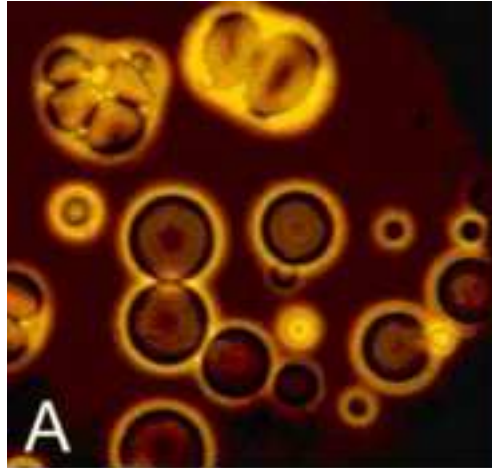


Fluorescent components have partitioned  
into the non-polar droplets

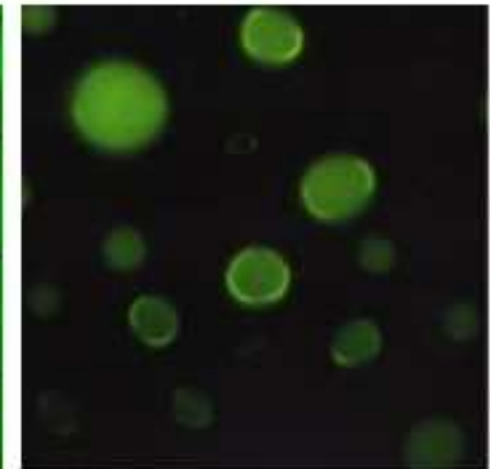
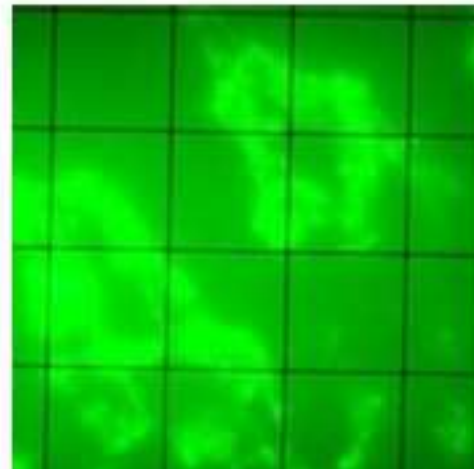
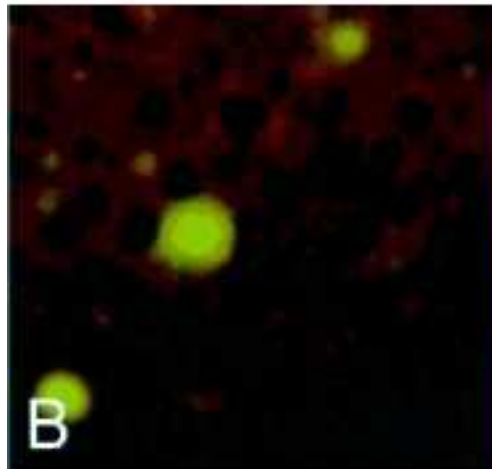


# Formation of Various Vesicular Structures from Meteorite and Ices

Phase Contrast  
Microscopy



Fluorescence  
Microscopy



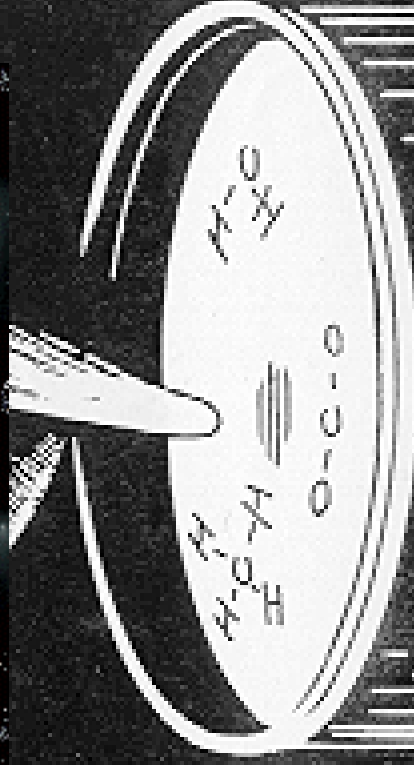
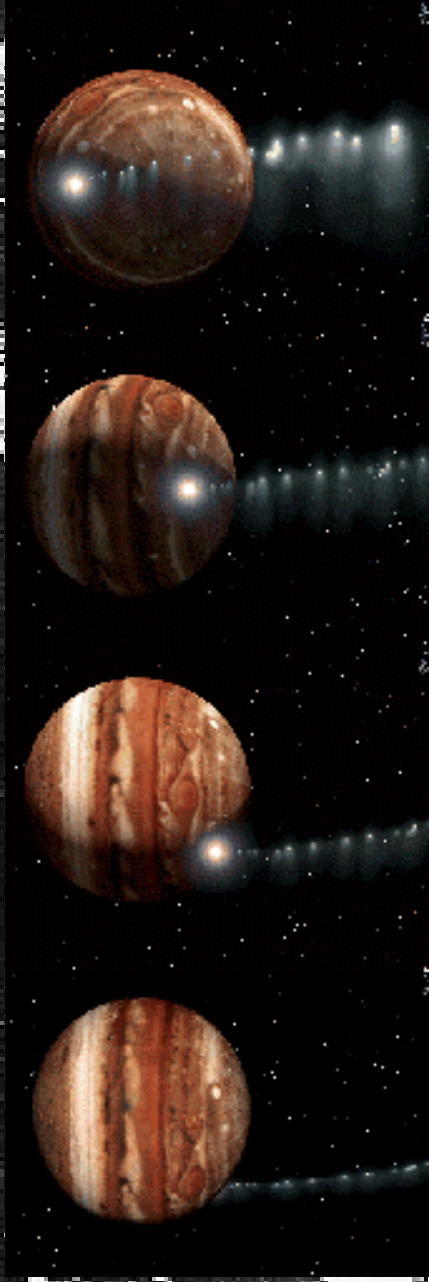
30  $\mu\text{m}$



Murchison Meteorite  
Deamer et al. 2003

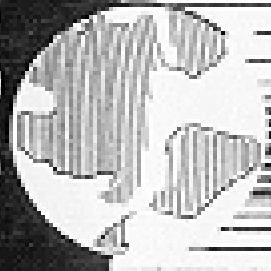
Proton Irradiated Ice  
Dworkin & Moore  
Work in progress

UV photolyzed Ice  
Dworkin et al. 2001



Primordial

S O U P



Primordial

S O U P



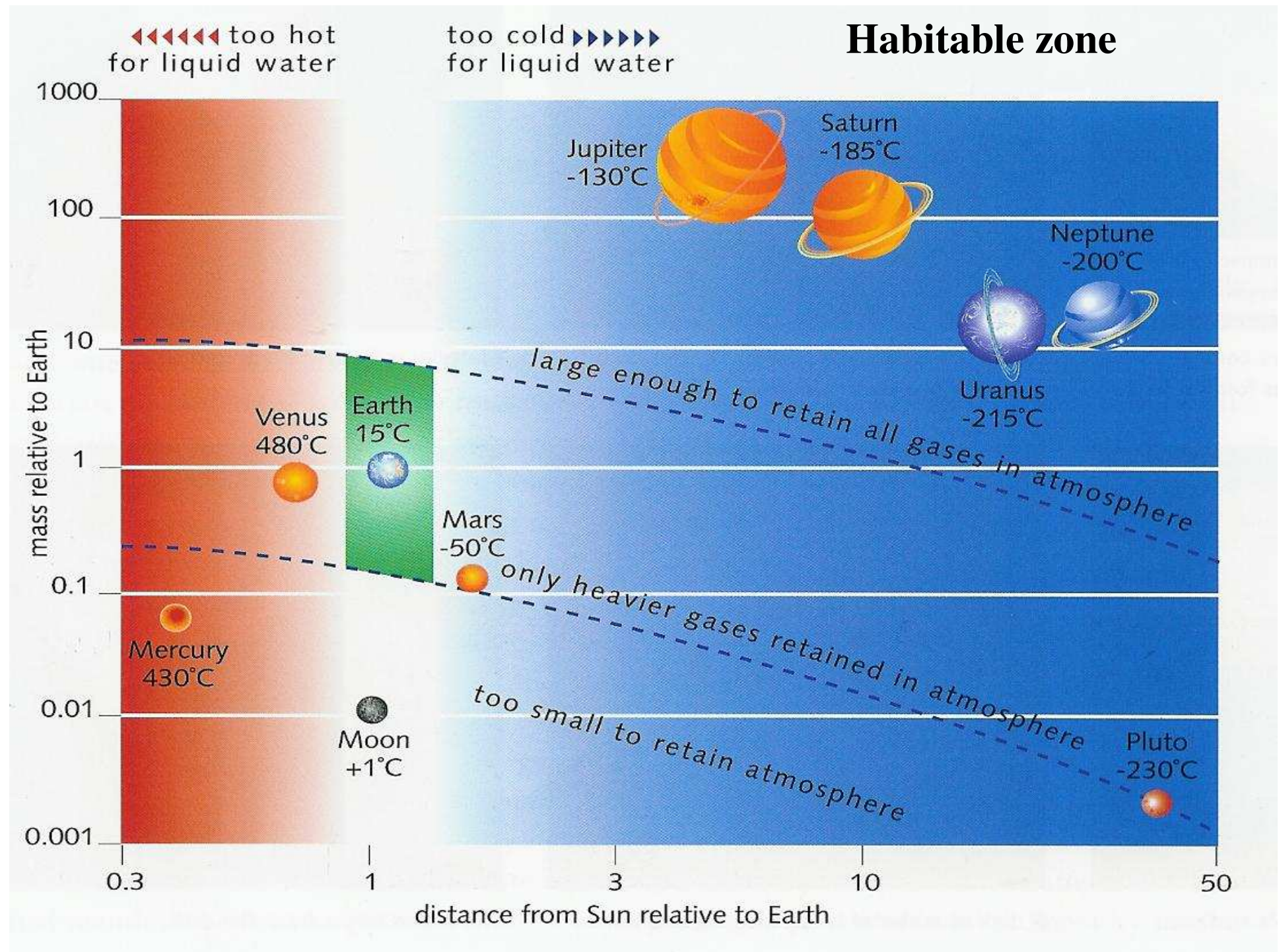


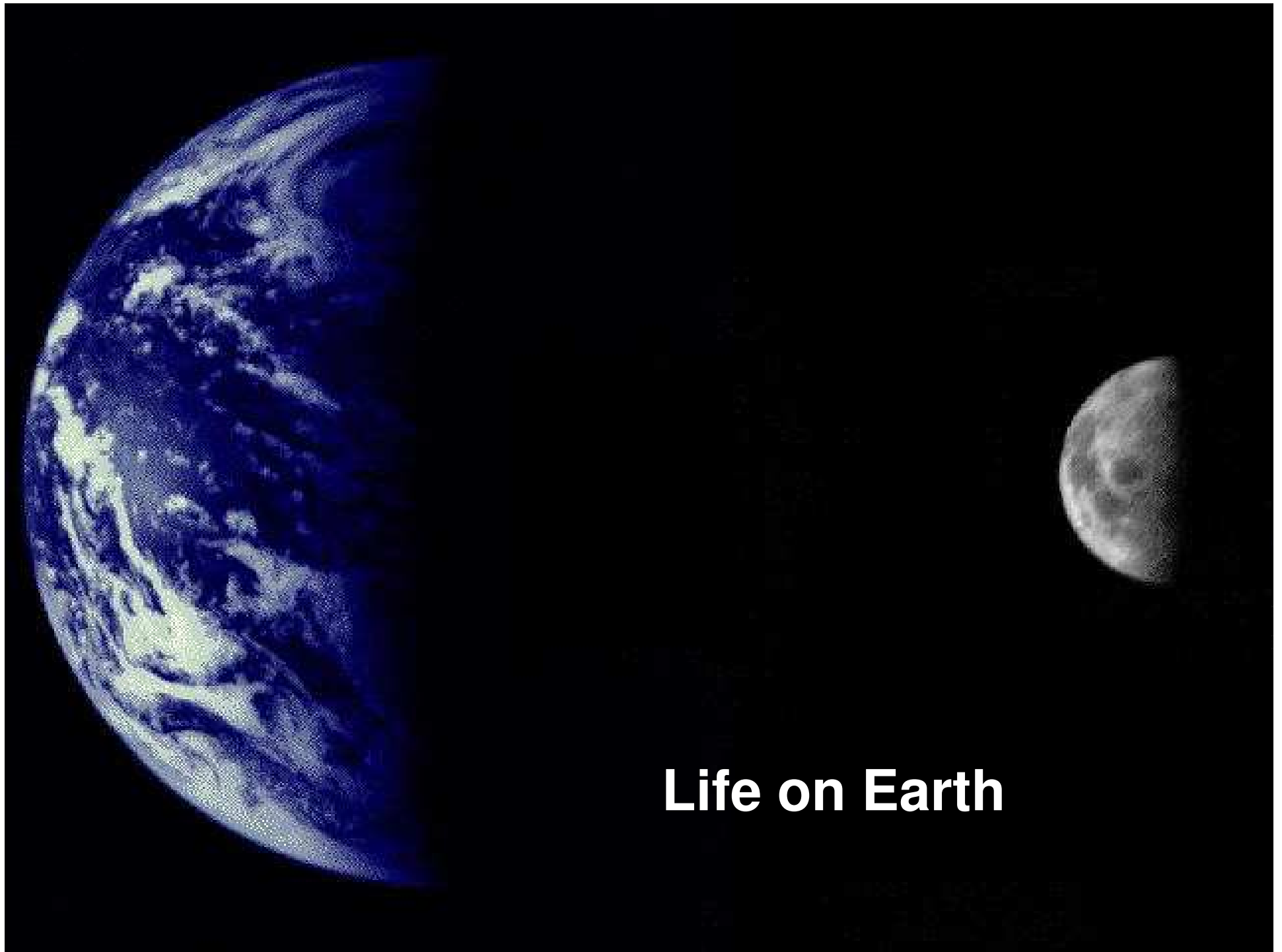


## Major Sources (in kg/yr) of Prebiotic Organic Compounds in the Early Earth

Terrestrial Sources	kg/yr*
UV Photolysis	$3 \times 10^8$
Electric Discharge	$3 \times 10^7$
Shocks from impacts	$4 \times 10^2$
Hydrothermal Vents	$1 \times 10^8$
Extraterrestrial Sources	adapted from Chyba & Sagan (1992)
IDP's	$2 \times 10^8$
Comets	$1 \times 10^{11}$
Total	$10^{11}$







**Life on Earth**

# EARLY EARTH

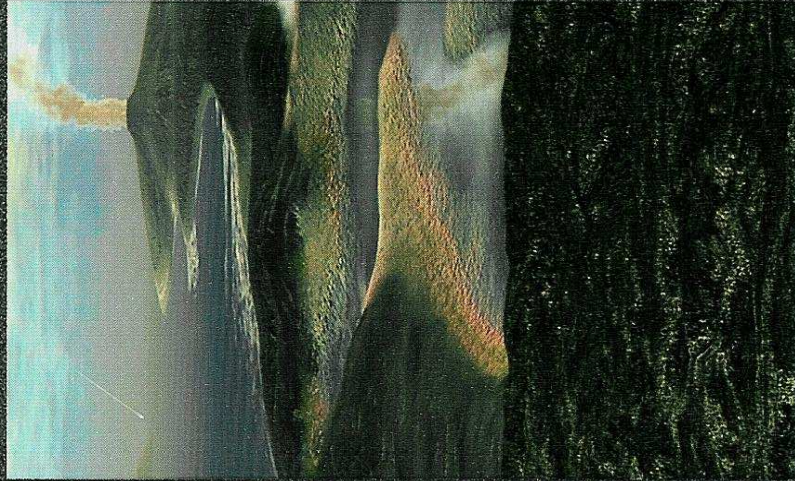




# PRIMORDIAL EARTH?



FREEZING?



TEMPERATE?



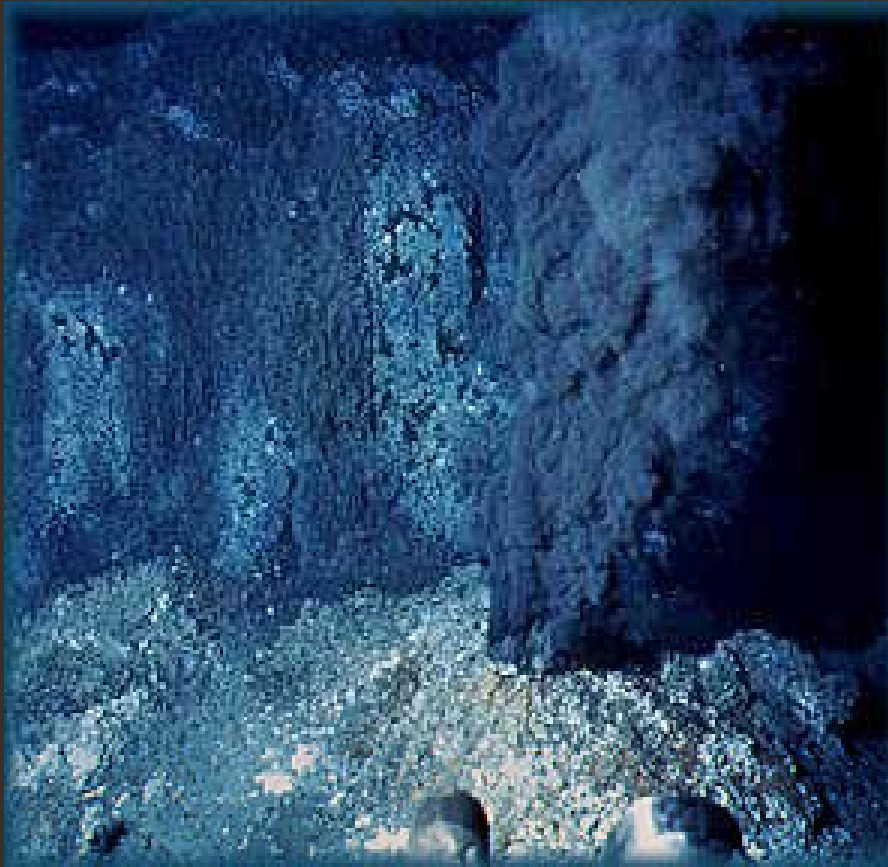
STEAMY?



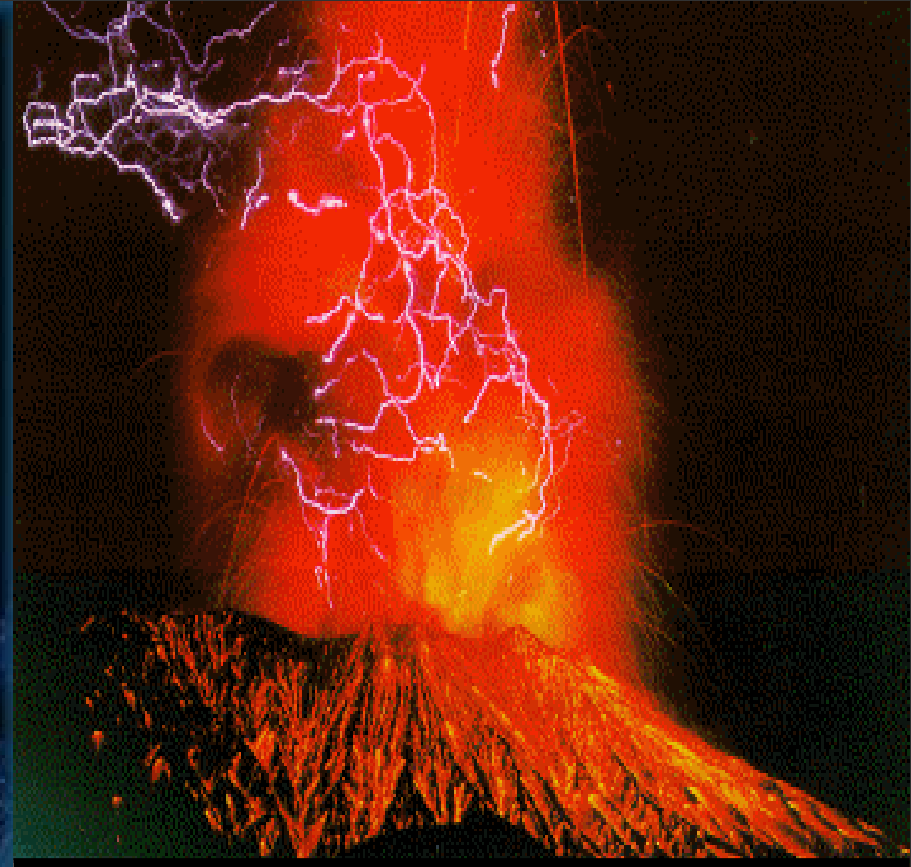


# Alternative abiotic synthesis routes

- Black Smokers

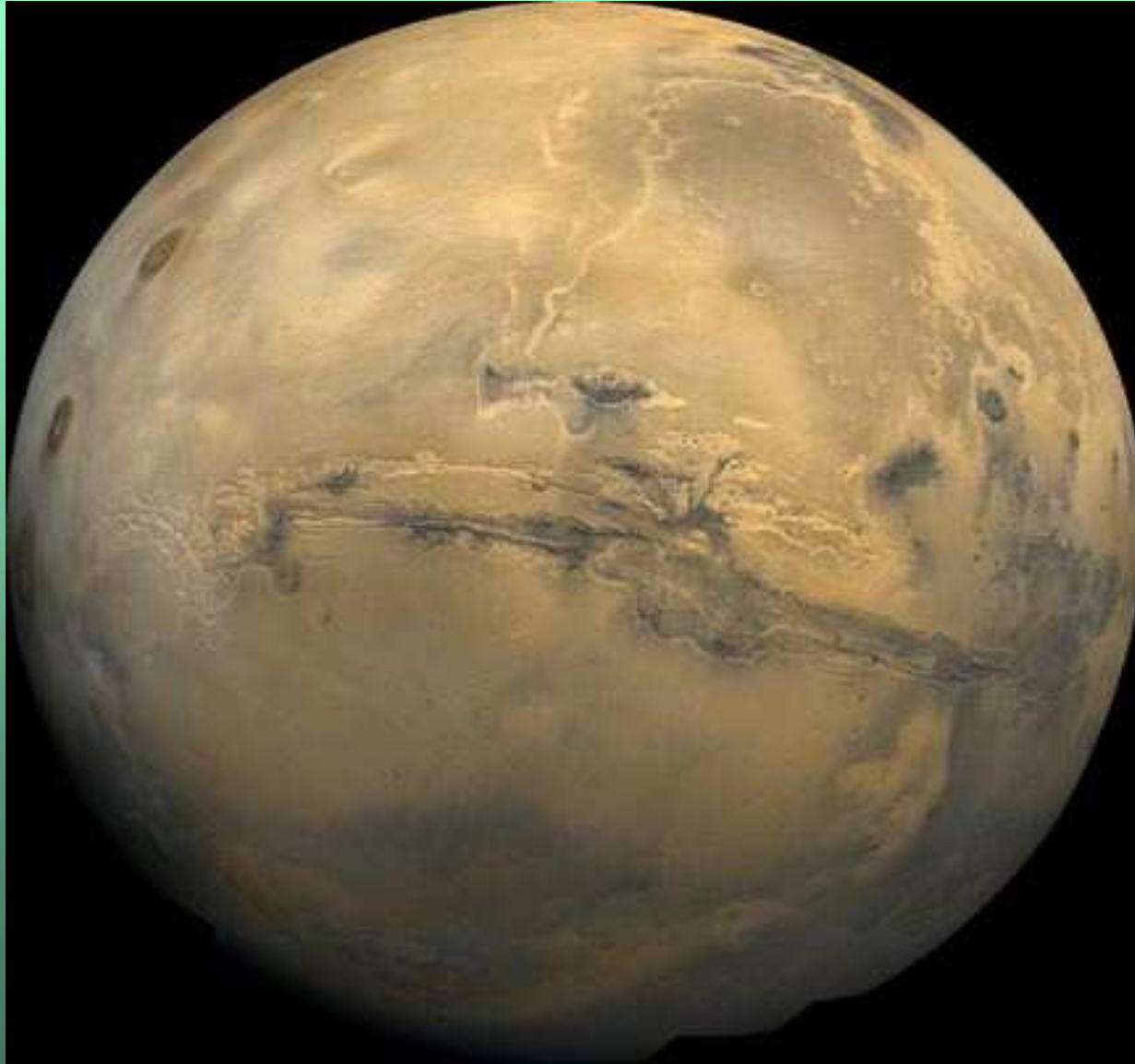


- Volcanic outflows



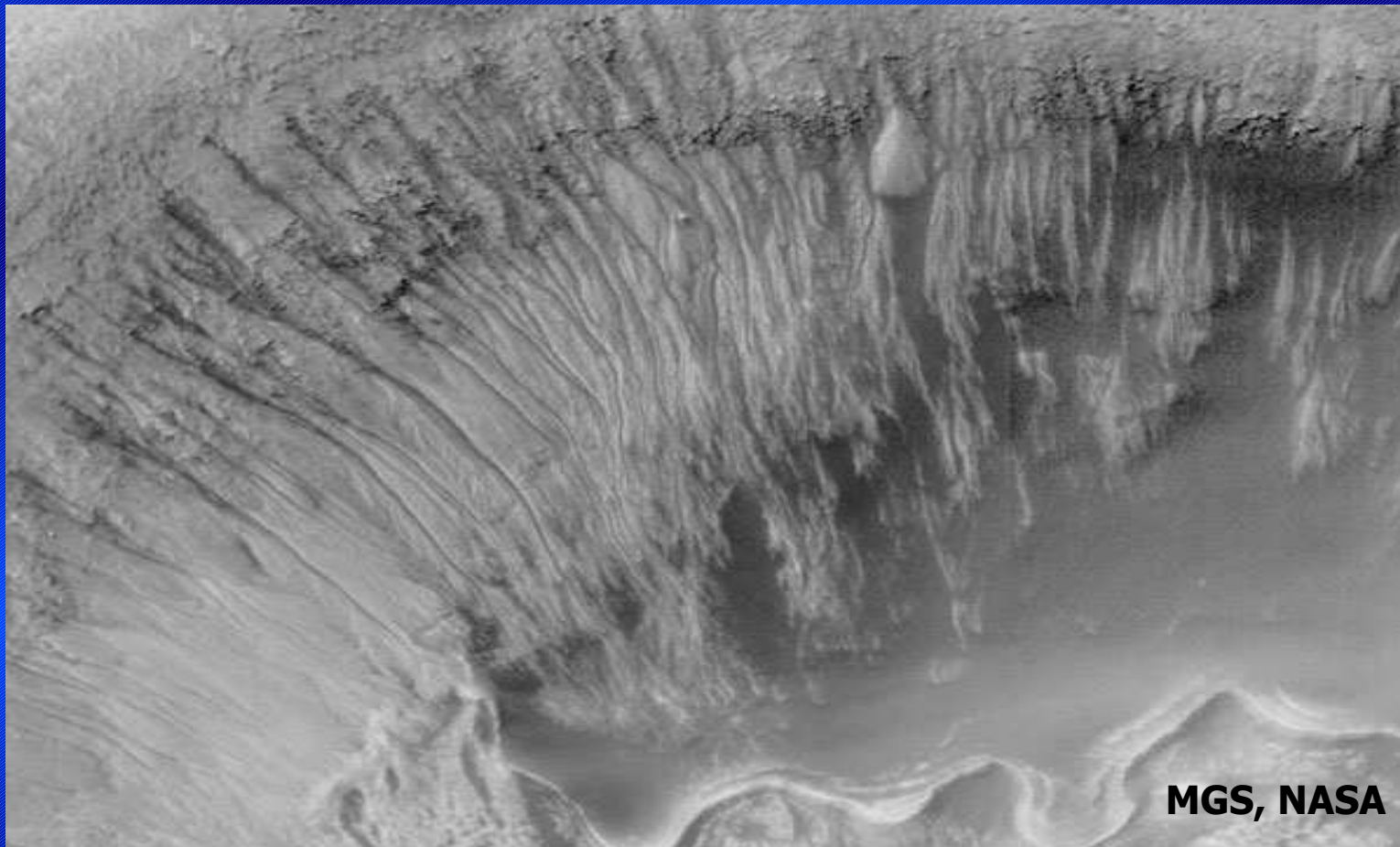


# Life on Mars

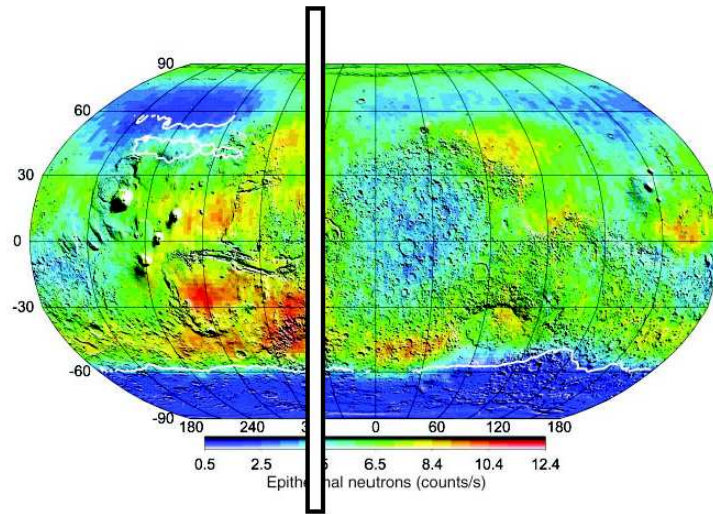


## Evidence for Recent Water on Mars

Many narrow channels run from top down to the crater floor in the Newton crater → outburst of subsurface water ?



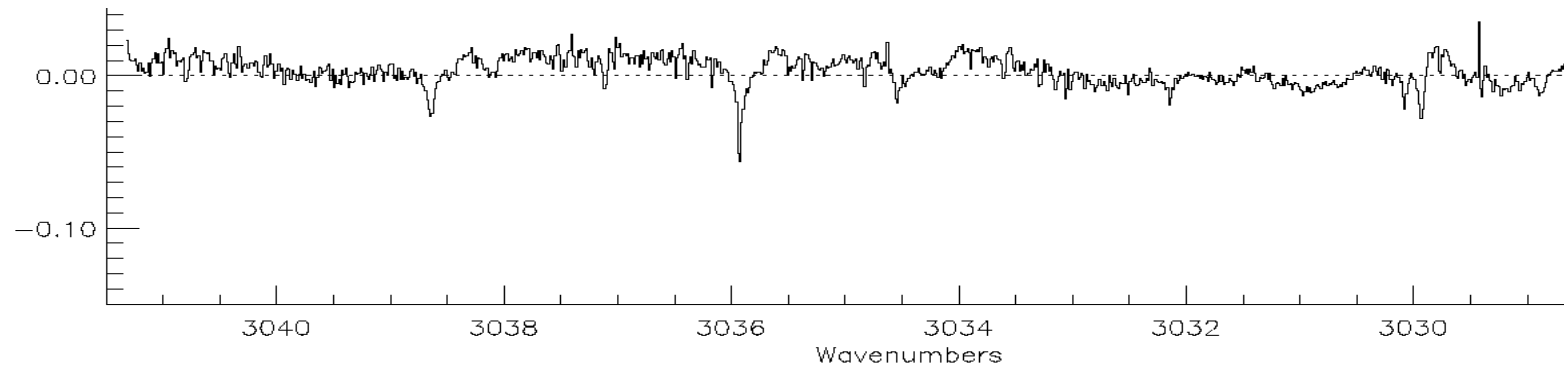
# Organics and Water on Mars



(left) sub-surface hydrogen

Mars Odyssey Boynton et al. 2002

(below) Phoenix spectra showing  $\text{CH}_4$  and  $\text{H}_2\text{O}$  (Mumma et al. 2003)



mumma\_010704.35

# Organic Compounds in Martian Meteorites

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- **ALH84001 was reported to show evidence for past life on Mars (McKay *et al.*, 1996).**
- **Upon analysis, only trace amounts of amino acids were detected (few ppb).**
- **The identified composition showed characteristics that pointed to the majority of the amino acids being terrestrial contamination, with a (possible) Martian contribution at the sub-ppb level (Bada *et al.*, 1998).**
-

# MARS EXPLORATION ROVERS

– robot rovers, 90 days mission, robots can crawl 100 m/day



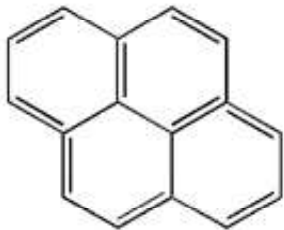
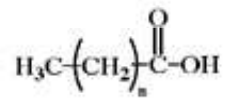
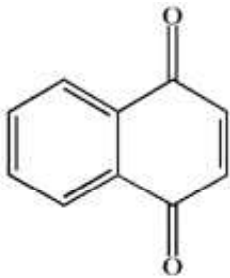
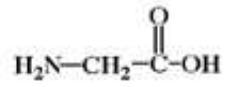
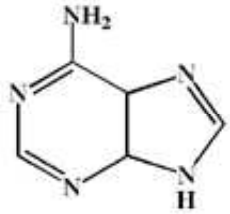


# ASTRObiology

**Could it be that our configuration of planets is extremely rare, perhaps even unique ?**

- **The right distance from the star**
- **The right mass of the central star**
- **Stable planetary orbits**
- **A Jupiter-like neighbour**
- **The right planetary mass**
- **Plate tectonics**
- **An ocean .....**

# astroBIOLOGY



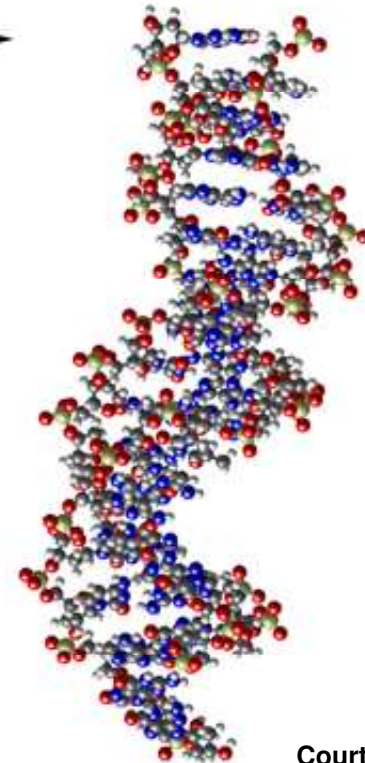
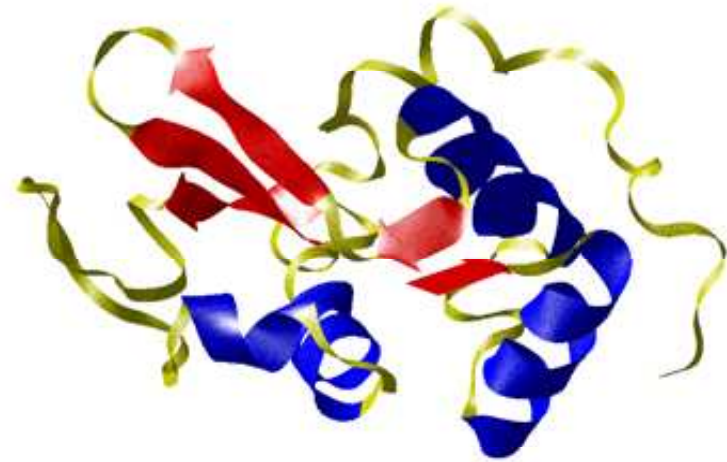
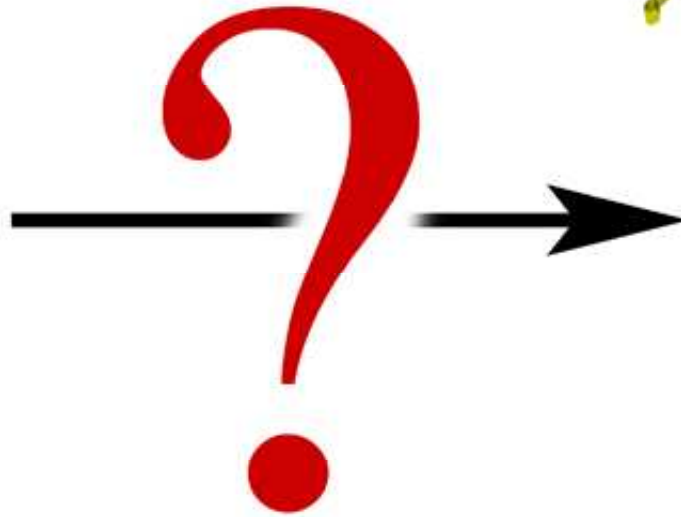
HCN

CH<sub>3</sub>OH

CO

NH<sub>3</sub>

$\begin{matrix} \text{O} \\ \parallel \\ \text{HCH} \end{matrix}$



Courtesy Jason Dworkin

**END**

